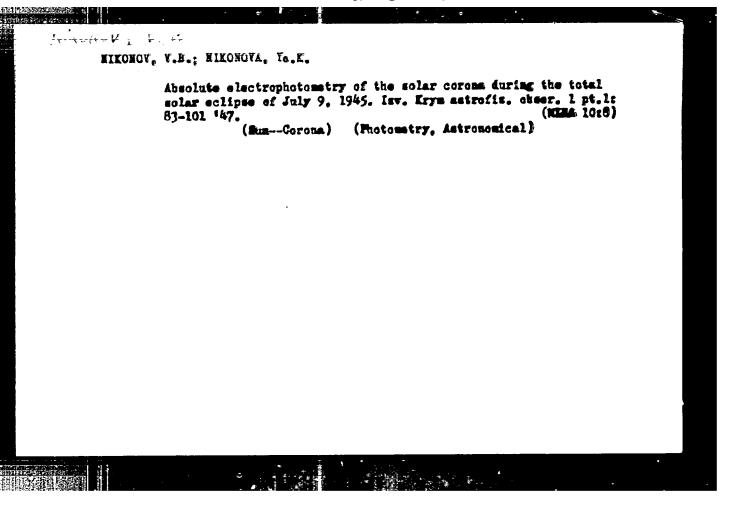
Astronomy

"Absolute Electrophotometry of the Solar Corona at the Time of the Complete Solar Eclipse of 9 July 1945." Izvestiya Frymskoy Astrofizioneskoy Observatorii, 1, 1948

Report No. -W-19569, RR 5 059020



HIKONOV, Y.B.

<u>Mikonov. V.B.</u> and Mikonova, E.K. "Experiment in photoelectric comparison of brightness of nocturnal skies in Simiens and Partisanovka," Izvestiya Krymsk. astrofis. observatorii, Vol. III, 1948, p. 109-11

SO: U- 2688, Letopie Zhurnal'nikh Statey, No. 1, 1949

HIKOHOT, Y. 3. AND MIKOHOTA, Ye. K.

Astronomy

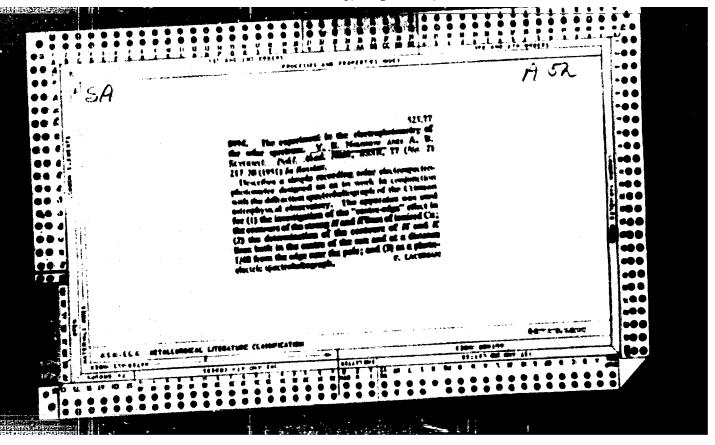
*Experiments with PhotoelectricalCom arison of the Brightness of the Night Sky in Simeiza and Partizanovka, Izvestiya Krymskoy Astrofizicheskoy Observatorii, 1948

Report No. -- W-19569. RR 52059020

HIROMOV, V. B., KRASOVEKIY, V. I. and KALINYAK, A. A.

Nablyudeniye oblasti galakticheskogo tsentra v infrakrasnykh luchakh (Observation of the Galactic Center Region in Infrared Rays). Akademiya Nauk SSSR. Doklady, 1949, v. 66, no. 1, p. 25-28, diagr., 6 refs.

AS262.83663 v. 66



MIKOFONOV, V.B.

USSE/Astronomy - Infrared Photography of Galaxy

Apr 52

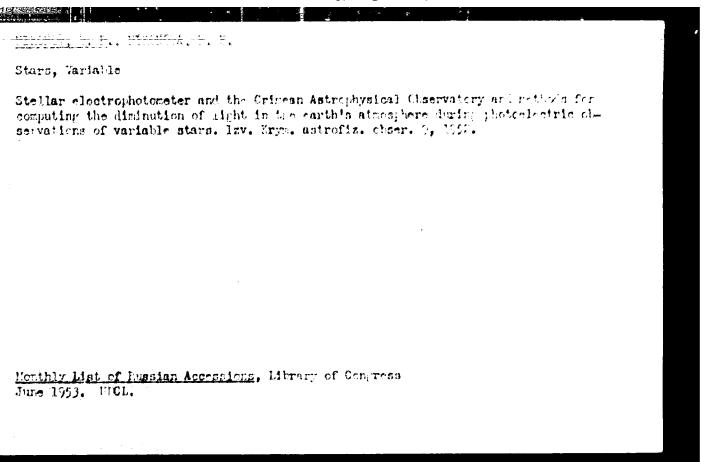
"Infrared Radiation of the Milky Way," Ye.W. Pavlov, Phys. Inst, Leningrad State U

Priroda No 4, pp 107-109

Real nature and structure of Galaxy were explained in 1948 at Crimean Astrophys Obs by observations in infrared of A.A. Kalinyak, V.I. Krasovskiy and V.B. Mikefeney, using photosensitive cathodes. In 1950 at the rame observatory, S.F. Rodionov and I. G. Frishman used photocells to photograph in infrared the Galaxy and found this radiation to be 2-10% of the background flow.

21511

"APPROVED FOR RELEASE: Tuesday, August 01, 2000 CIA-RDP86-00513R001137.



"APPROVED FOR RELEASE: Tuesday, August 01, 2000 CIA-RDP86-00513R001137.

Stars, lariable

Photoelectric observations of a variable star of the Selbei type, IW Vuljecular.

Inv. Krym. astrofiz. obser. 5, 1652.

Monthly list of Russian Accessions, Library of Congress
June 1953. UNCL.

D-1442 Ten Profess

Stellar Astronomy, Stellar Catalogs (3987)

Byull. AN Gruz, ESR, Abastumanskaya Astrof. Observatoriya, No 14, 1953, pp 1-233 Nikonov, V. B.

Accurate color equivalents are deduced for 652 stars in the galactic zone of latitude 20°, limited in the region of galactic center by declination -24° and in the region of anticenter by longitude 160°. The atmospheric effect was elimated, but not light absorption by interstellar media.

Referativnyy Zhurnal -- Astronomiya i Geodeziya, No 6, 1954 (%-30976)

KIKOHOV, V. B.

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USSR/Astronomy - Bibliography

Fa1/AUB 53

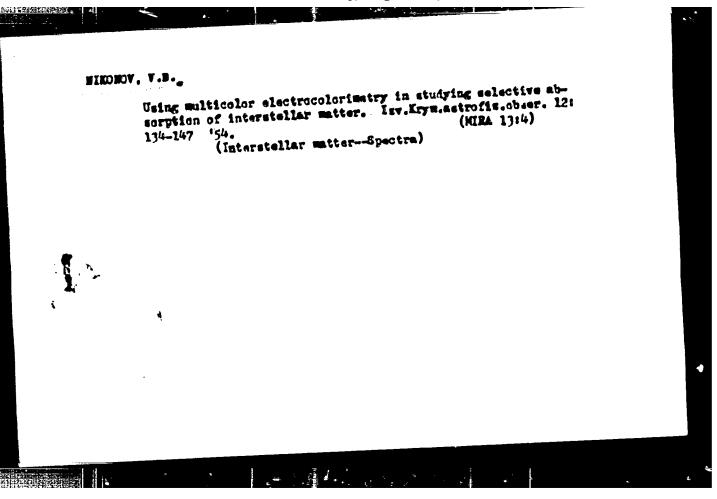
"Index to Astronomical Literature Published in the USSR During March, April 1953," Yu. G. Perel'

Astr Zhur, Vol 30, No 4, pp 475-478

Lists 60 articles on astronomy which appeared during Mar, Apr 53 in 15 books and symposia, 1 ephemeris, 6 institute organs (e.g., "Trudy"), 17 periodicals, and 5 abstracts of dissertations. For example, one author's abstract of a dissertation, the only one mantioned for the degree of Dr Phys-Math Sci, is: An Attempt to Construct a Fundamental Catalogue of Electrical Chromatic Equivalents of

262132

Stars of the Spectral Types B8 and B9," by V. B. Hikonov, Main Astron Observatory of Acad Sci USSR (expts were completed at Abastuman Astrophys Observatory and Crimean Astrophys Observatory), Leningrad, 21 pp, 100 copies.



"APPROVED FOR RELEASE: Tuesday, August 01, 2000

CIA-RDP86-00513R001137

HIKOHOV, V.B.

USSR/Astronmay - Instruments

Card 1/1

Pub. 43 - 39/97

Authors :

Dobronravin, P. P., and Nikonov, V. B.

Title

Instrument for recording the energy distribution in spectra of stars

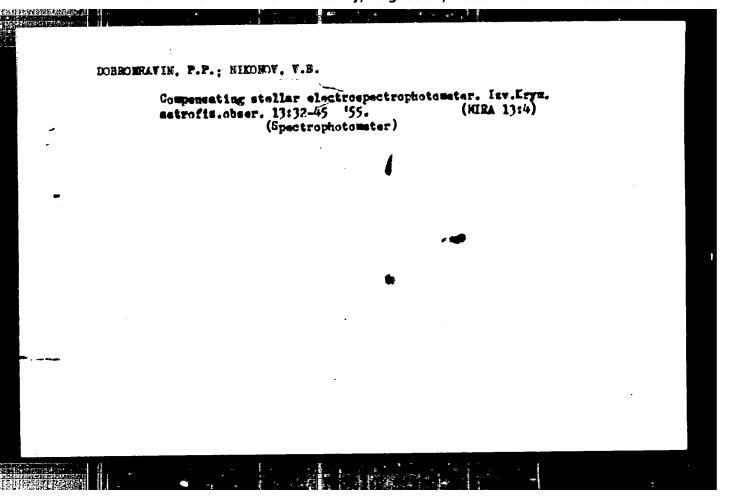
Feriodical : Izv. AN SSSR. Ser. fiz. 18/2, page 268, Mar-Apr 1954

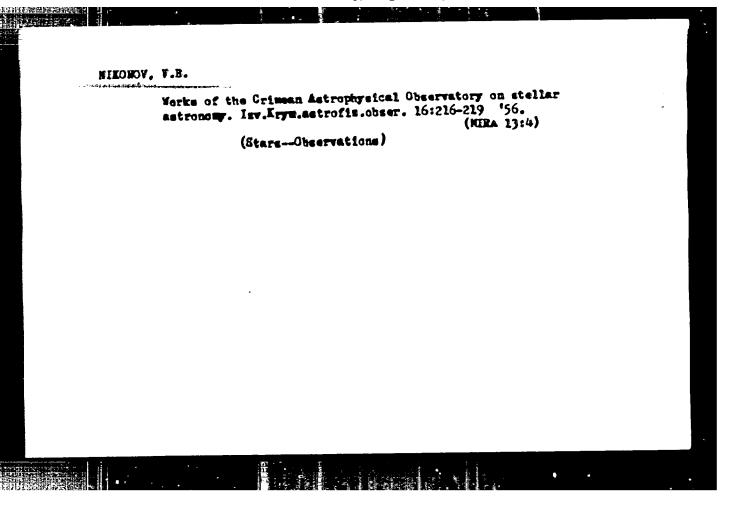
Abstract

Brief description is given of a device for recording energy distributions in spectra of stars. The instrument apploys Soviet made photo-multipliers and is intended for operation in a 500 mm meniscus telescope. The device records the ratio of the photocurrent produced by the light of a star at a given wave length and the photocurrent produced by a large part of the spectrum. The light escillations caused by the flickering and vibration of the image on the slit are completely eliminated by the new recorder. This also includes the chromatic flickering of stars at greater zenith врасер:

Institution : Academy of Sciences WoR, The Crimean Astrophysic Observatory

Submitted





HINDROY, T.B.: HERRASOVA, S.V.: POLOSUKHIHA, H.S.: HACHKOVS:IT, D.N.:
CHUTAIN, E.K.

Color-luminosity diagram for stars in the vicinity of the
sum. Inv.Krym.astrofiz.obser. 17:42-88 '57.
(NIRA 13:4)

(Stars)

· 3(1) AUTHORS:

307/20-121-5-13/50 Batslov, M. M., Zavoyskiy, Ye. K., Corresponding Member, Academy of Sciences, USSR, Kalinyak,

A. A., Nikonov, V. B., Prokof'yeva, V. V., Smolkin, G. Ye.

TITLE:

The Use of Multistage Electron-Optical Light Amplifiers in Astrophysics (O primenenii mnogokaskadnykh elektronno-

opticheskikh usiliteley sveta v astrofizike)

Doklady Akademii nauk SSSR, Vol 121, Nr 5, PERIODICAL: pp 815 - 818 (USOR) 1408

This paper investigates some problems connected with the ABSTRACT:

application of electron-optical light amplifiers in astrophysics. The authors estimate the increase in efficiency of the utilization of the photon flux with

respect to the usual photographic method. Under the investigated conditions, and in the case of equal dimensions of the pictures, the efficiency of the electron-optical method is by 4.100 times higher than in ordinary photo-

graphy. An increase in scale on the photocathode of the light amplifier reduces the increase in sensitivity of the electron-optical method compared with a usual photographic

plate by 160 times. An estimation of the sensitivity

Card 1/3

The Use of Multistage Electron-Optical Light Amplifiers SST/20-121-5-13/50 in Astrophysics

of the light amplifiers gives a value of the order of 1000. The use of an electron-optical amplifier usually cannot increase the penetration range of the telescope. But the reduction of the times of exposure by hundreds of times of its amount due to the high sensitivity of the light amplifier essentially changes the possibilities of the astrophysical investigation. The short times of exposure permit the investigation of rapidly varying processes of very faintly visible objects and a considerable increase of the utilization coefficient of the astrophysical instruments. The reduction of the times of exposure is very important for autrospectroscopy. The above-discussed considerations are confirmed by the results obtained by experiments carried out by the authors in the Krymskaya astrofizicheskaya observatoriya AN SSSR (Crimean Astrophysical Observatory AS USSR). The proper noises of the light amplifier may be neglected in comparison with the background of the sky. According to the experimental values, the use of the light amplifier permitted a reduction of the timer of exposure approximately to a thousandth part of their former amount

Card 2/3

The Use of Multistage Electron-Optical Light Amplifiers SOV/20-121-5-13/50

which satisfactorily agrees with the above-given estimate. A figure shows the photographs of 2 extragalactic nebulae which were taken by means of a light amplifier. There are 4 figures, 1 table, and 6 references, 3 of which are Soviet.

ASSOCIATION: Krymskaya astrofizicheskaya observatoriya Akademii nauk SSSR

(Crimean Astrophysical Observatory AS USSR) Glavnaya astronomicheskaya observatoriya Akademii nauk SSSR (Astronomical

Main Observatory, AS USSR)

SUBMITTED: April 14, 1958

Card 3/5

30264

5/035/61/000/010/001/034 101A\101

3,1510 (1114,1129,1166)

Mikonov, V.B.

TITLE

Photoelectrical determination of equivalent widths of H 70 in early

stare

Referativnyy zhurnal. Astronomiya i Geodeziya, no. 10, 1961, 22, ab-stract 10A167 ("Izv. Krymak. astrofiz. observ.", 1960, v. 22, 176-PERIODICAL:

186, Engl. sussery)

The authors discuss the possibility of direct photoelectrical determination of equivalent widths of speatral lines by the "differential storing" method. In this method, light fluxes are measured from the portion of speatrum which includes the line being studied, as well as those from two portions of continuous spectrum to both sides of the line. The measurements of the both light fluxes are conducted by means of a single photomultiplier and alternate every 0.5 sec. Photogurrents from the line and continuous spectrum are integrated independently. The method described was applied to determination of equivalent widths of H pin spectra of 22 stars of spectral classes 09 - AO. Observations established the possibility of organizing the large-scale photoelectrical determina-

Card 1/2

APPROVED FOR RELEASE: Tuesday, August 01, 2000

CIA-RDP86-00513R0011372

\$/035/62/000/010/016/128 A001/A101

AUTHORS:

Mekrasova, S. V., Nikonov, V. B., Polosukhina, N. S., Rybka, Ye.

TITLE:

Photoelectric magnitudes and colors of reference photometric stars in Kapteyn areas. I. Some problems in methods of com-

piling fundamental photometric catalogues

PERIODICAL:

Referativnyy zhurnal, Astronomiya i Geodeziya, no. 10, 1962, 30, abstract 10A244 ("Izv. Krymsk. astrofiz. obzerv.", 1962, v. 27,

228 - 240) ,

TEXT: A catalogue of photoelectric magnitudes and colors of reference photometric stars in Kapteyn's areas is necessary to reduce zero-points of scales of stellar magnitudes to a single system, as well as in allowance for atmospheric extinction. The authors set forth the task of observation of all reference photometric stars in 139 Kapteyn's areas ($\delta > -15^{\circ}$). In the future, observations should be extended to the entire southern half of the sky. Two methods are briefly described (Ye. Rybka and V. B. Nikonov) for compiling such a catalogue. Both of the methods are applied to the same observational data obtained in Crimea by means of an A3T-7 (AZT-7) meniscus telescope. In more Card 1/2

VIS

Photoelectric magnitudes and colors of ...

S/035/62/000/010/016/128 A001/A101

detail these methods were described earlier. Observations of 14 reference stars in 7 northmost Kapteyn's areas are utilized (results are tabulated), as well as of 17 stars of spectral classes BO-M2 from Johnson's list. Methods of observations and processing are described. It turned out that both of the methods yield errors of the same order (OMOI), however Nikonov's method is more economical in time consumption and makes it possible to control more reliably the constancy of the photometric system. It was decided to use the latter method for the further work on the catalogue (individual observations are directly extrapolated beyond the atmosphere). It is established that relation should be used in compiling catalogues of stars with a wide range of colors. There are 14 references.

B. Pesenko

[Abstracter's note: Complete translation]

Card 2/2

37397

\$/033/62/039/002/011/014 E032/E314

3.1260 AUTHORS:

Card 1/3

Butslov, M.M., Kopylov, I.H., Nikonov, V.B.,

Severnyy, A.B. and Chuvayev, K.K.

TITLE: Experiments in electron-optical photography of

galaxies in hydrogen light using the 2.6 m

reflector of the Crimean Astrophysical Observatory

PERIODICAL: Astronomicheskiy zhurnal, 4. 39, no. 2, 1962, 315 - 322 + 3 plates

TEXT: Detailed studies of extragalactic nebulae require the use of large telescopes. As regards detecting apparatus, the use of ordinary photographic techniques in conjunction with narrow-band filters necessitates long exposures and is therefore inconvenient in practice. The authors have investigated therefore the possibilities of image-converters as a means of avoiding these disadvantages. An image-converter was set up in the direct focus of the 2.6 m reflector of the Crimean Astrophysical Observatory. The immediate object was to investigate the hydrogen emission in a number of galaxies. Four light colour filters were introduced in front of the converter and

Experiments in electron-optical ...

5/033/62/039/002/011/014 B032/B314

the screen of the latter was photographed by a motion-picture camera. Altogether 58 galaxies were photographed in Ha and other light. Photographs of 10 of these are reproduced and their features are described (NGC 604, 1569, 4214, 4449, 4490, 4736, 5194, 5457, 6822 and 6946). Many unknown clouds of hydrogen-emission were detected in the galaxies. In many cases there is no correspondence between hot-star clusters and hydrogen clouds. The hydrogen component shows greater concentration in the equatorial planes than the stellar component. In some galaxies the nuclei consist of isolated condensations. The dimensions of the nuclei in H_{α} light are in some cases appreciably larger than in other light, although in a number of cases the reverse situation obtains. In several galaxies, streams or ejections from the nucleus, which are visible only in H_a light, were detected.

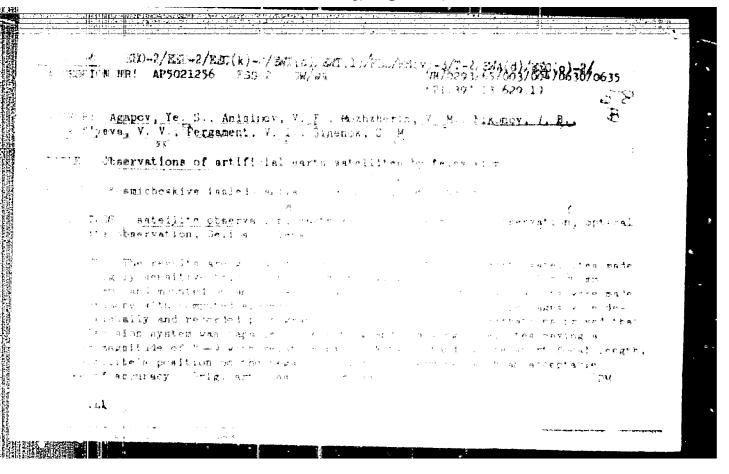
Card 2/3

AGAPOV, Ye.S.; ANISIMOV, V.F.; NIKONOV, V.B.; PROKOFYYEVA, V.V.; SINEHOK, S.M.

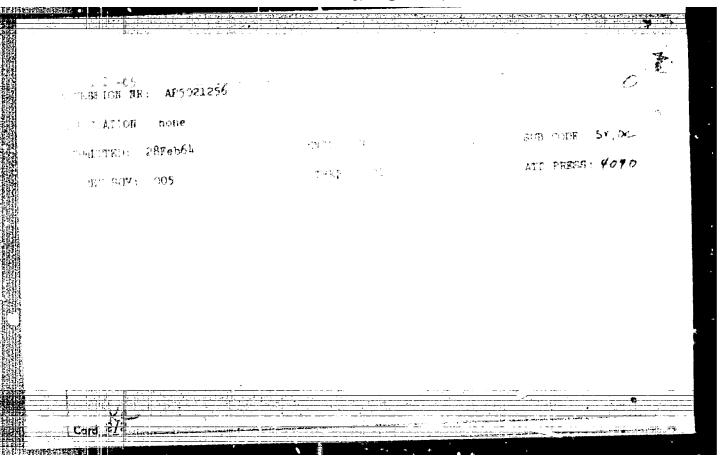
Experimental application of television technique for observations of stars. Izv. Krym. astrofiz. obser. 30:3-18 '63.

(MIRA 17:1)

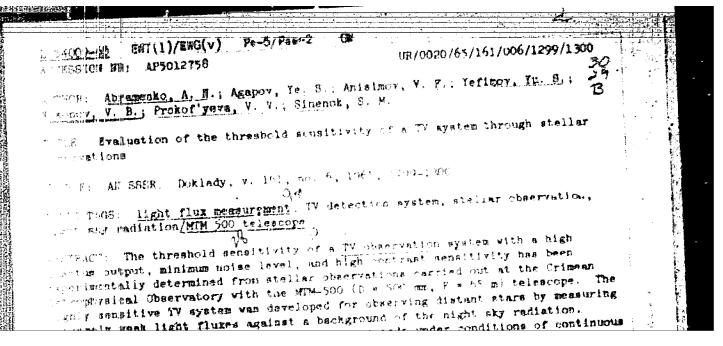
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HIKCKOV, V. F.

Nikonov, V. F. "Problems of mathematical training in the "technicums"." Min Education RSFSR. hoseow Oblast Pedagogical Inst. Moscow, 1955. (Dissertation for the Degree of Candidate in Pedagogical Science)

So: Knishnaya letopis', No. 27, 1956. Moscow. Pages 94-109; 111.

NIKONOV, V. F.

A TRANSPORTER

The first Meld. Minkevich prize was given to the following teams:
Candidate of Technical Sciences A. D. Assenov, Engineers W. I. Tereshchin,
V. P. Mikonov, D. I. Kostenko, S. G. Marinchev, I. S. Yurkov, M. M. Inshakova,
M. M. Yanchuk, A. A. Bulstnikov and G. Ye. Litvin (Automobile Works imeni
Likhachev) for their paper "Investigation and Introduction of the Process of
Mitrocementation by Direct Isothermal Hardening in an Alkali Inside Muffleless
Equipment", their design of a muffleless furnace heated by vertical radiation
tubes is of interest.

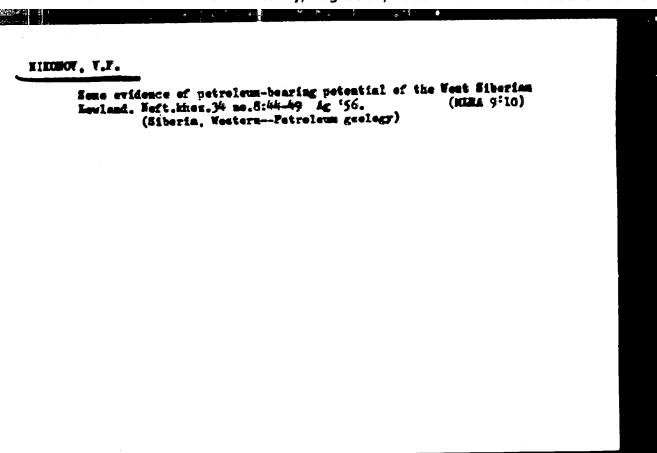
Results of the 1958 Competition for Obtaining imeni D. K. Chernov and imeni N. A. Minkevich Prizes, Metallovedeniye i termicheskaya obrabotka metallov, 1959, No. 6, pp 62-64

General Company, P.Ya.; Bikomov, V.P.; Vorlinekiy, A.G.

Weing 40 Rhoff steel for the helf-exists of automobile driving axists. Netallored. i term. obr. met. mc.6:15-19 Je '63.

(Automobiles—Axists)

(Steel alleys—Testing)



5(0) AUTHOR:

Mikonov. Y. P.

807/20-124-2-45/71

TITLE:

Authigenic Iron, Sulfur, Organic Carbon, and Bitumina in the Mesosoic Sediments of the Eastern Ural Region (Autigentoye sheleso, sera, organicheskiy uglerod i bitumy v osadkakh mesosoya Vostochnogo Priural'ya)

PERIODICAL:

Doklady Akademii nauk SSSR, 1959, Vol 124, Er 2, pp 392 - 394 (USSR)

ABSTRACT:

As is known, the organic substance is oxidized by authigenic oxide iron and sulfates. This oxidation may take place to a considerable extent within short periods (Refs 1,2). The process mentioned in the course of which oxidation iron and sulfate-sulfur are reduced is widespread (Refs 3,4). The authigenic iron is precipitated from the solution as oxidation iron (Refs 5,6). Sulfur is carried away from the regions of denudation and accumulates first in the sediments as 50^{2-} unless it migrates as a component of the organic substance and is embedded. Part of the organic substance is decomposed by redox processes. This takes place the more

Card 1/4

Authigenic Iron, Sulfur, Organic Carbon, and Bitumina in the Mesoscic Sediments of the Eastern Ural Region 507/20-124-2-43/71

quickly the thicker the original mass of the oxidizer is. For this reason there exists a certain interdependence between the stages of accumulation of authigenic iron, sulfur and the organic substance. The content of authigenic iron increases along the cross section of the Mesozoic from the Upper Jurassic to the Cenomanian, it then rapidly decreases and later on increases again. Due to this fact 2 periods of accumulation of authigenic iron are expressed in the denudations (Fig 1). The distribution of Cors

and inversely proportional to the iron concentrations (Ref 7). Sulfur, as one of the elements which are the most capable of migration (Refs 11,12) was washed away and accumulated during the great transgressions (Jurassic-Valenginian, Albian, Turonian-Oligocenic) in the sediments, i.e. at times when iron was little mobile. There is a high sulfur content in all loamy strata (0.4-0.9%). The concentration of Abitumina decreases in the Jurassic until the Cenomanian, while the iron content increases. In rocks from the Turonian to the Oligocenic which contain less iron but more sulfur,

Card 2/4

Authigenic Iron, Sulfur, Organic Carbon, and Bitumina in the Mesozoic Sediments of the Eastern Ural Region

507/20-124-2-43/71

the concentration of bitumen is strongly reduced. This took place in connection with an accumulation of large quantities of sulfate while C is not less than in older sediments. The rocks of the Jurissic-Valanginian rich in Corg and bitumen, however, contain much sulfur, i.e. even more than the organic substance itself. Sulfates are practically lacking. In the sediments of the Turonian-Oligocenic sulfur is mainly of mineral origin (about 0.25% of pure sulfur). The content of sulfate sulfur and that of bitumen are inversely proportional to each other. The iron, sulfur, Corg and bituminous content on the whole increases towards the central part of the Khanty-Kansiysk basin; the maximum amounts of these components are found towards the western edge of the basin (Fig 2). The strongest concentrations of authigenic iron are found in loasy sediments at greatest depths, while they are bound to the edge in sandy deposits. There are 2 figures and 12 Soviet references.

Card 5/4

Authigenic Iron, Sulfur, Organic Carbon, and Bitumina in the Mesoscic Sediments of the Eastern Ural Region

807/20-124-2-43/71

ASSOCIATION:

Tyumenskoye territorial'noye geologicheskoye upravleniye

(Tyumen'Territorial Geological Administration)

PRESENTED:

1931×00720×0070×04.

September 20, 1958, by S. I. Mironov, Academician

SUBMITTED:

September 19, 1958

Card 4/4

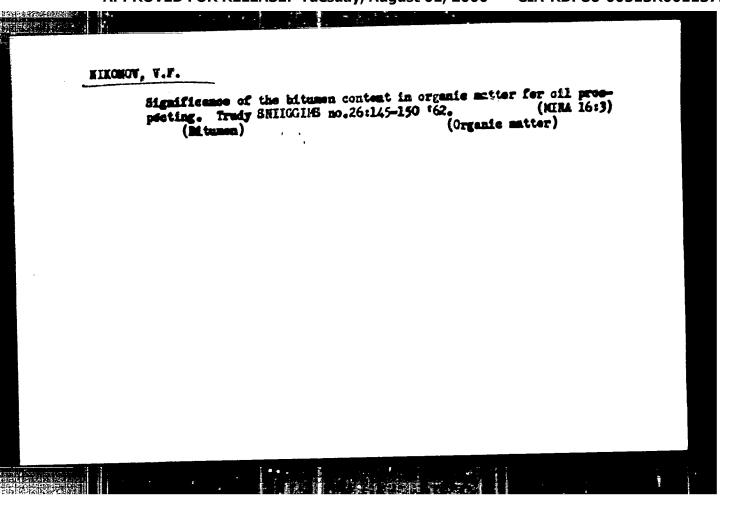
Distribution of organic carbon, bitumens, and heavy hydrocarbons throughout the profile of Neco-General deposits in the eastern part of the trans-Ural region with regard to the occurrence of petroleum and gas. Bokl. AN SEER 134 no.3:654-657 S '60. (NIRA 13:9)

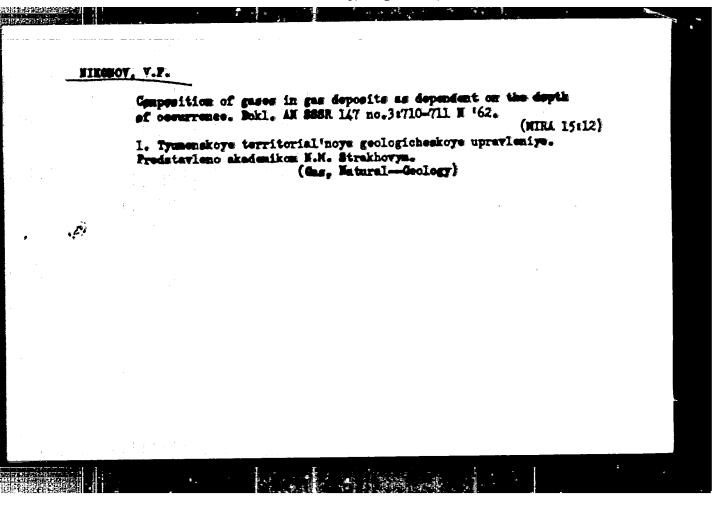
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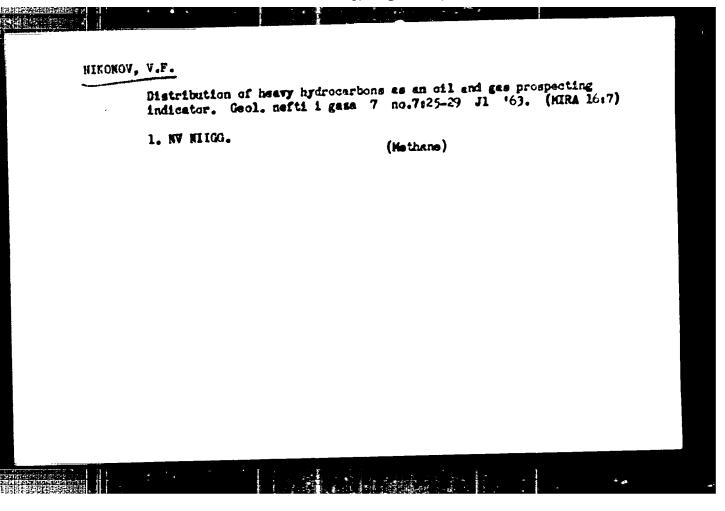
BOYARSKIKH, G.K.; MIKONOV, V.F.; PROKUPENKO, V.I.; RUVNINA, L.V.; ROMANUV, F.I.;
YASTREBOVE, T.A.; SVENCHKOV, G.P.. nauchnyy red.; MEVEL SPITIL, V.I.,
veducichly red.; YASHCHURZHINGKAYA, A.B., tekhn.red.

[Key wells of the U.S.S.R.; Beresovo key well (Tyumen* Prevince)]
BeresoveMin operacia skvashina (Tiumenskaia eMlast*). Lemingred
Gos. nauchno-tekhn. isd-vo neft. i gorno-tephivnei lit-ry, Lemingr.
otd-ie. 1962. 120 p. (Lemingrad. Vsesoiusnyi neftianoi nauchnoissledovatel*skii geologorasvedochnyi institut. Trudy, no.195)
(MIRA 15:12)

(Beresovo region (Tyumen* Province)—Geology)

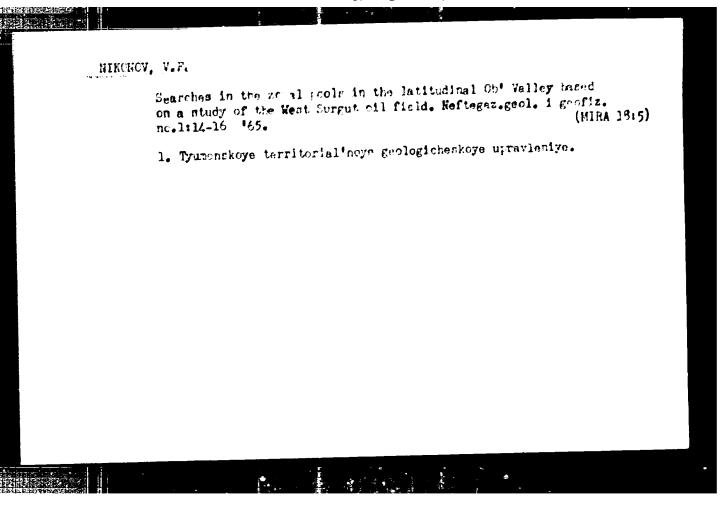


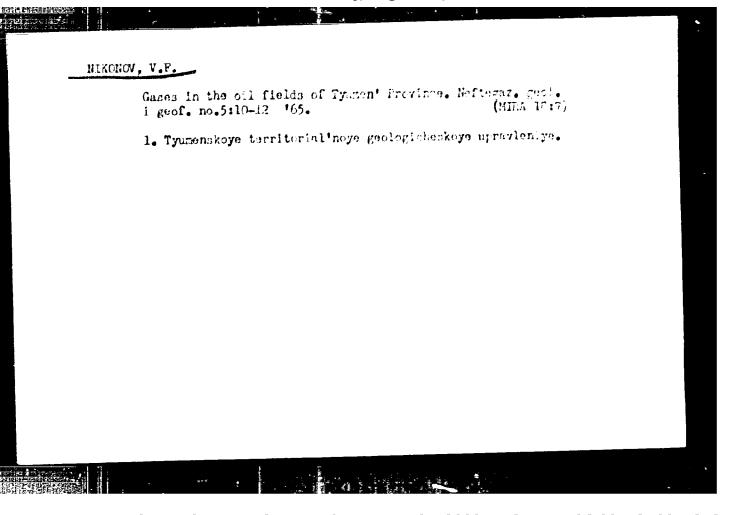


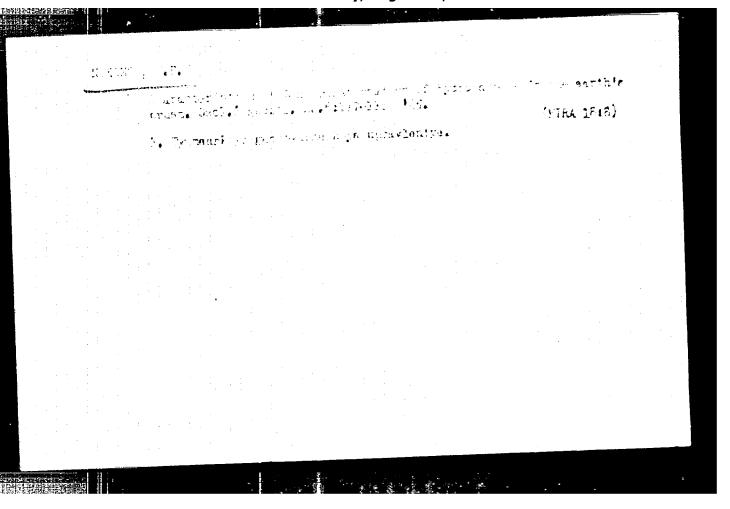


Composition of the crudes of the northern Ob'valley.
Geol. nefti i gaza 8 no. 1:20-23 Ja *64. (MIFA 17:5)

1. Tyumenskoye territorial*noye geologicheskoye upravleniye.







SOLODIKHIN, Alekeanar origon/pevich; BIKCGOV, V.F., reteenzent;
VE.T., G.Ya., inch., red.

[Technical and economic principles of the hoat treatment of metals] Tekhnicheskie i ekonomicheskie osnovy termichesko obrabotki metallov. Moskva, Mashirostroenie, cheskoi obrabotki metallov. Moskva, Mashirostroenie, (MIRA 18:11)

Variation of gas composition within gas fields. Lokl. AN SSSR 165 no.41927-929 D '65. (MIRA 18112) 1. Tyunenskoye territorial noye geologicheskoye upravleniye. Submitted July 3, 1965.

"APPROVED FOR RELEASE: Tuesday, August 01, 2000

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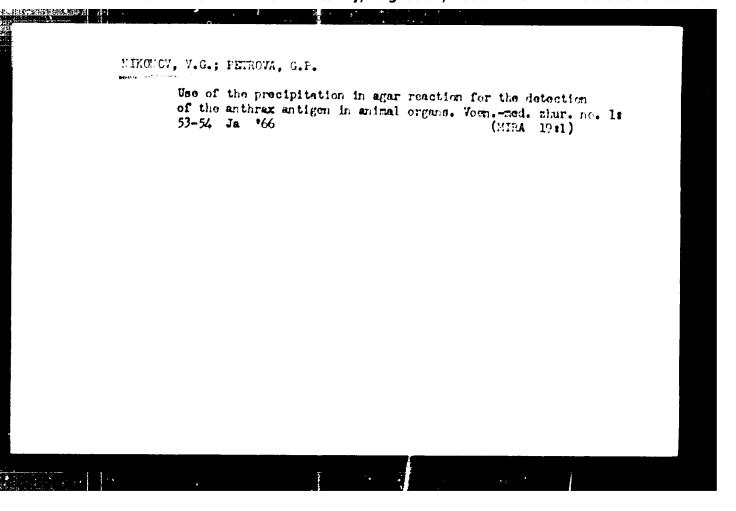
15213-66 ENT(m)/EWA(d)/T/EMP(t)/EMP(s)/EMP(b)/EMA(h) JD CC NR: AP6002912 SOURCE CODE: UR/0286/65/000/024/0074/0074	
NVENTOR: Shepelyakovskiy, K. N.; Strogonov, K. V.; Shklyarov, I. H.; Orlov, I. V.; ikonov, V. F.; Assonov, A. D.	
RG: none	
OURCE: Hyulleten' izobreteniy i tovarnykh znakov, no. 24, 1965, 74	
OPIC TAGS: steel, surface hardened steel, manganese containing steel, silicon con- sining steel, chromium containing steel, shallow hardenable steel	
BSTRACT: This Author Certificate introduces a steel for surface-hardened parts conaining 0.4—1.2% carbon and alloyed with manganese, silicon, and chromium. To obtain teel with a specified hardenability, one of three alloying elements is added in a pecified amount and the content of the other two is limited. For example, in steel ontaining 0.3—1.4% manganese, the chromium and silicon contents are limited to 0.15% and 0.17%, respectively. Steel with 0.3—1.4% silicon should contain 0.15% chromium and 0.20% manganese, and steel with 0.3—1.8% chromium should contain 0.20% manganese and 0.17—0.27% silicon.	
UB CODE: 11/ SUBM DATE: 29Dec60/ ATD PRESS: 4190	
iari 1/1	,

ZAYKOV, S.T.; KRAVISOV, P.Ya.; NIKIFOROV, B.V.; KOVAL', V.Ye.; THIGULIN, V.I.;
RUBINSKIY, P.S.; LIFSHITS, S.I.; YEVSTAF'YEV, Ye.I.; NIKONOV, V.F.;
VOZLINSKIY, A.G.

Using oxygen-blown converter steel in automobile manufacture.

Met. 1 gornorud. prom. no.4:26-31 Jl-Ag '64.

(MIEA 18:7)



Heating of jover transfermers with untiffied correct. Enangetik.
13 no.4:27-24 Ap 165.

(M.RA 10:4)

Production of fodder yeast at the Kotlass sulfite alcohol plant.

Gidroliz. i lesokhim. prom. 18 no.3424 *65. (MIRA 18:5)

ADELISON, S.V.; NIKOMOV, V.I.

Effect of dilution on the dekydrogenation of isopentame in the presence of ledime. Kkim, i tekh, tepl. i masel 10 no.3: 11-13 Hr '65.

1. Nackevskiy ordena Trudevege Krasnego Zmameni institut meftekkimieheskey i gasevey promyehlennosti im. akad. Gubkirm.

HIKOHOV, V. L.

NIKONOV, V. L.: "Principles of Constructing a System of Forest Parks in the Suburban Forest-Park Zone of the City of Leningrad." Him Higher Education USSR. Leningrad Order of Lenin Forestry Engineering Academy imeni S. H. Kirov. Leningrad, 1956. (Dissertation for the Degree of Candidate in Agricultural Science)

So: Knishnaya Letopis', No. 19, 1956.

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ARAPOV, A.D.; HIKOHOV, V.N.

Homoplasty of arteries in an infected wound; experimental observations.

Youn. med, shur. mc,2:23-30 F'57 (MIRA 12:7)

(ARTHRIES, experimental,
homoplasty of arteries in infected wds. (Rus.))

(WOUNDS AND INJURIES, experimental,
same)
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AEDREYIVA, Z.F., kand.khimicheskikh nauk; HIKONOV, V.N.,
ispolayayushehiy obyusannosti stafshego nauchnogo
sotradnika

Trihydroxyglutaric acid and its salts as a buffer in
separating lanthanides by means of ion exchange
chromatography. Isv. TSIHA no.3:196-205 '62. (NIRA 15:9)
(Clutaric acid)
(Chromatographic analysis) (Rare earth metals)

HIROMOV, V.N.; SAKOVICH, A.A.

ngasari asari

Electric current converter with silicon rectifier for use in electrometallurgical enterprises. Electrichestro nc.12:55-59 p 162. (MIRA 15:12)

1. Wesoyumyy elektrotekhnicheskiy institut imeni Isnina. (Kleatric current restifiers)

HINDROV, V.H. (COU, Corthly)

"The Investigation of Oscillation Fluctuation of a Klystron Oscillator."

Considered random amplitude and phase modulation of klystron oscillations, caused by random processes connected with the electron flow. Experimental measurements of oscillation parameter fluctuations showed a coincidence with the theory first developed by I. L. Bernshteyn.

report presented at the All-Union Conference on Statistical Redio Physics, Gor'kiy, 13-18 October 1958. (Izv. vyssh uchev zaved-Rediotekh., vol. 2, No. 1, pp 121-127) COMPLETE card under SIFOROV, V. I.)

807.29

E/141/59/002/06/010/024

E192/E382

9,4220

中的人工方面的特別的

likosov. Yalle

AUTHOR: TITLE:

Investigation of the Oscillation Fluctuations in a

Elystrom Generatory

PERIODICAL: Investiya vysshikh uchebnykh zavedeniy, Radiofizika,

1959, Vol 2, Nr 6, pp 915 - 926 (USSR)

ABSTRACT: First, the problem is investigated analytically. Under the assumption that the resonator of the klystron can be represented by an L. C and R equivalent circuit, the

dynamic equation of the system is:

$$\frac{di}{L} = v - Ri;$$

$$\frac{dt}{dt} = v - Ri;$$

$$c \frac{dv}{dt} = -1 + I_o \left[\frac{1}{1 + (c_o/2U_a)dv_1/dt_1} - 1 \right]$$
 (1a)

is the alternating voltage across the capacitor and is the oscillation current in the equivalent circuit.

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APPROVED FOR RELEASE: Tuesday, August 01, 2000

CIA-RDP86-00513R0011372

80129 \$/141/59/002/06/010/024 _\$182{\$352 & Klystron

Investigation of the Oscillation Fluctuations in a Klystron Generator

The second component in Eq (1a) represents the current induced by the electron beam. I_G is the cathode current, is the average electron transit time (from the resonator to the reflector and back) and U_G is the voltage applied to the resonator. In order to investigate the oscillation fluctuations it is assumed that these are due to the shot noise and the flicker effect in the electron beam; the thermal noise is neglected. If the shot—and flicker-effect directly induced currents are denoted by $f_1(t)$ and $f_2(t)$ and the components due to the reflected beam are denoted by $f_1^*(t)$ and $f_2^*(t)$, the dynamic equations of the system can be written as Eqs (2) and (2a). Employing polar coordinates r and r (defined by the last equation on r 916), Eqs (2) and (2a) can be rewritten as Eqs (3) and (3a), where r and r are expressed by Eqs (36). Expressions for the amplitude and expressed by Eqs (36).

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80129 S/141/59/002/06/01Q/024 Investigation of the Oscillation Fluctuations in a Klystron Generator

phase fluctuations are therefore given by Eqs (4) and (4a). These can also be expressed as Eqs (5) and (5a). The solution of these shows that the spectral densities of the amplitude and frequency fluctuations are given by Eqs (6) and (6a), respectively. If the generator is isochronous, the spectral densities of the amplitudes and frequency fluctuations are given by Eqs (7) and (7a). These can further be transformed into Eqs (8) and (8a). The effect of the flicker noise alone on the amplitude and frequency fluctuations is described by Eqs (9). The problem was also investigated experimentally; the equipment employed in the measurements was similar to that described in Ref 2. The results of the experiments are illustrated in Figures 3, 4 and 5 and in two tables on pp 924 and 925. Figure 3 shows the frequency dependence of the spectral density of the frequency fluctuation for a klystron operating in various mades. The spectral density of the amplitude fluctuations as a function of frequency is illustrated in Figure 4. It is found that the experimental results are in good agreement with the theory. For the normal operating modes the

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Investigation of the Oscillation Fluctuations in a Klystrom Generator

measured spectral line width is 0.1 to 0.5 c/s; the spectrum band of the amplitude fluctuations varies between 20 and 50 Mc/s and the spectral density from 10^{-16} to 10^{-15} c/s.

There are 5 figures, 2 tables and 8 references, 1 of which is English and 7 are Soviet.

ASSOCIATION: Nauchno-issledovatel'skiy radiofizicheskiy institut
pri Gor'kovskom universitete (Scientific-research
Radiophysics Institute of Gor'kiy University)

SUBMITTED:

May 17, 1959

Card 4/4

5/141/60/003/006/62+/025 E192/E382

9,4220

Nikonov, V.N. AUTHOR:

Concerning a Method of Measuring the Amplitude TITLE:

Fluctuations of an Oscillator

Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika, 1960, Vol. 3, No. 6, pp. 1129-1130 PERIODICAL:

Recent work by Haggblom (Ref. 1) described a method of measuring the spectral density $\mathbf{v}_{\mathbf{Q}}(\Omega)$ of the amplitude

fluctuations of a klystron oscillator in the low-frequency region. A block diagram of this equipment is illustrated in Fig. 1. The investigated oscillator 1 is frequency-modulated by the oscillator 7; the modulated output of the oscillator is applied to a crystal detector 2 where the doubled modulation frequency is separated. A signal is then detected again by means of a vacuum diode 4 and its spectrum is analysed by means of a spectral analyser 5. The calibration of the instrument is performed by means of an oscillator 8 which produces a calibration signal at the input of the analyser.

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Concerning a Method

The above method of measurement is discussed in some detail and it is shown that, in his experiments, Haggblom (Ref. 1) measured the spectral density of the amplitude fluctuations of the klystron oscillator over the frequency range from 30.001 to 30.1 Ec/s instead of the frequency range 0.001 to 0.1 Mc/s (as was assumed by Haggblom).

There are 1 figure and 3 references: 2 Soviet and 1 non-Soviet.

ASSOCIATION:

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Nauchno-issledovatel'skiy radiofizicheskiy

institut pri Gor'kovskom universitete

(Scientific Research Radiophysics Institute

of Gor'kiy University)

SUBMITTED:

April 29, 1960

Card 2/2

5/141/61/004/001/009/022 25949 E192/E362 9,3260 Malakhov, A.N. and Nikonov, V.N. 10 AUTHORS: Correlation of the Amplitude and Frequency Fluctuations in Oscillators TITLE: Izvestiya vysshikh uchébnykh zavedeniy. Radiofizika, 1961, Vol. 4, No. 1, pp. 104 - 112 , PERIODICAL: It is assumed that the signal produced by the oscillator is expressed by: $x(t) = A_0[1 + \alpha(t)] \cos[\omega_0 t + \int_0^t \gamma(\xi) d\xi]$ where $\alpha(t)$ and V(t) are relative fluctuations of amplitude and frequency which are in the form of stationary random processes such that |α| ≪ 1. |ν| ≪ ω, ā = 5 ÷ 0 where the line above the symbols represents statistical Card 1/8

5/141/61/004/001/009/022 25949 Correlation of the Amplitude E192/E382 averaging. In order to determine the spectral density of the power fluctuation $W_{\mathbf{x}}(\omega)$, it is necessary to determine 3.0 the correlation function $\Phi_{\mathbf{x}}(\mathbf{t})$ which, in turn, is dependent on the correlation functions of the amplitude fluctuations Ψα(T), frequency fluctuations Ψ, (T) and cross correlation These functions have been determined for some special cases (Ref. 1 - 5.M. Rytov - ZhETF, Vol. 29, 304, 315, 1955). On the other hand, in this work an attempt is made to find $\Phi_{\alpha}(\tau)$, $\Phi_{\gamma}(\tau)$ and $\Phi_{\alpha\gamma}(\tau)$ for a large range of oscillators. It is assumed that the oscillator signal is 70 governed by the following differential equation (Ref. 2 -A.N. Malakhov - Izv. vyssh. uch. zav. - Radiofizika, Vol. 3. 241, 1960): (2) $\sum_{i=1}^{n} a_{ik} \frac{d^{ik} x}{dt^{ik}} = F\left(x_{i}, ..., \frac{d^{ik} x}{dt^{ik}}, ...\right) + E(t).$ Card 2/8

5/141/61/004/001/009/022 25969 £192/E382 Correlation of the Amplitude The fluctuations vi(t) where E(t) represents the noise. and a(t) are expressed by the following equations (Ref. 2): (5) + pV = (-a 2 + a 2 + b 2 + b 2 + b 2)/640 (4) . a + pa = (a E + a E)/6A, Eq. (5) can also be written in a more convenient form: (5) V = P1 + (a121 - a 21)/640 $p_1 = (a_{ij}b_{ij} + a_{ij}b_{ij})/6$. The noise can be expressed by $E(t) = E_{\parallel}(t)\cos(u_0t) - E_{\perp}(t)\sin(u_0t)$ (6) Card 3/8

race and	24	ξ, • •	# # # # # # # # # # # # # # # # # # #		
	Correlation	25949 of the Amplitude	S/141/61/004/ E192/E382		
	where E _i (t)	and E_(t) at	re slowly changing fu onents can be express	notions of time	}•
		= Sett and the	$= \frac{(1/\omega_u) \dot{E}(\mathbf{Q} \sin(\omega_u t);}{(1/\omega_u) \dot{E}(t) \cos(\omega_u t)}.$	(5).	
	The correlat	tion function of	the noise is given !	p à:	91
•			7 f (m) cos(ms) dm	(8)	
	which can b	e approximately	be expressed by:	. :	• • •
14		Φg (τ) = A*(τ) cos (·	$m_0\tau) = A^1(\tau) \sin(m_0\tau).$	(10)	
	where A ⁰ defined by:	and Al are sle	owly changing function	ns which are	
	Card 4/8				
			• • • • • • • • • • • • • • • • • • •		4

Correlation of the Amplitude ... E192/E382 $A^*(t) = \int \widetilde{W}_E(U) \cos(Ut) dU = \int \widetilde{W}_E^*(U) \cos(Ut) dU;$ $A^*(t) = \int \widetilde{W}_E(Q) \sin(Ut) dQ = \int \widetilde{W}_E^*(Q) \sin(Ut) dQ.$ By solving Eq. (4) with respect to a(t) it is found that: $a(t) = \int f(t-\xi)e^{-p\xi}d\xi \qquad (14)$ where $f(t) = (a_{\frac{1}{2}} \sum_{i=1}^{n} e_{i+1})/\delta A_{o}$. The final expression for the amplitude correlation function, derived on the basis of Eq. (14), is in the form:

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$$\bigoplus_{\alpha}(\tau) = \frac{1}{p\delta A_o^2} \int_0^{\infty} \frac{A^o(\tau + y) + A^o(\tau - y)}{2} e^{-py} dy$$
(17)

From this it is easy to find the spectral density of the amplitude fluctuations; this is expressed by:

$$W_{\rm g}(\Omega) = 2W_{\rm E}^{\rm o}(\Omega)/6A_{\rm o}^{\rm o}(p^{\rm o} + \Omega^{\rm o})$$
 (18)

Similarly, it is shown that the correlation function for the frequency fluctuations is given by:

$$\psi_{\tau}(z) = \frac{A^{\bullet}(z)}{zA_{0}^{2}} + \frac{P_{1}^{2}}{zA_{0}^{2}\rho} \int_{0}^{z} \frac{A^{\bullet}(z+y) + A^{\bullet}(z-y)}{2} e^{-\rho z} dy + \frac{2\rho_{1}}{zA_{0}^{2}} \int_{0}^{z} \frac{A^{\bullet}(z+y) - A^{\bullet}(z-y)}{2} e^{-\rho z} dy.$$
(20)

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S/141/61/004/001/009/022 Correlation of the Amplitude ... E192/E382

while the cross-correlation function for the fluctuations is:

$$\Phi_{n}(\tau) = \frac{1}{2A_{0}^{2}} \left\{ \frac{\rho_{1}}{\rho} \int_{0}^{A^{2}(\tau + y)} \frac{A^{n}(\tau - y)}{2} e^{-\rho \tau} dy + \int_{0}^{\pi} A^{n}(\tau + y) e^{-\rho \tau} dy \right\}.$$
(22)

The above expressions are used to analyse some special cases:

- 1) WEQUE WO!
- 2) the derivative of $W_{\mathbf{g}}(u)$ in the vicinity of $u = u_{\mathbf{g}}$
- is not equal to sero, and
- has a width comparable with p or less than p. From the above it is concluded that for the existence of cross correlation between the amplitude and frequency fluctuations of an oscillator it is necessary and sufficient that one of the following conditions be fulfilled:

5/141/61/004/001/009/022 Correlation of the Amplitude ... E192/E362 a) simultaneously, by \$60 and by \$0 is the case when an oscillator is nonisochronous; b) simultaneously, $a_{ii} \neq 0$ and $a_{j} \neq 0$, which occurs when the lefthand-side portion of Eq. (2) contains even as well as odd derivatives, and c) the spectrum of the noise in the vicinity of wa asymmetrical with respect to w . It is clear that in all actual oscillators at least the third condition is fulfilled. It is therefore necessary to take into account the crosscorrelation between the frequency and amplitude fluctuations in determining the spectral line W_(w) of the oscillator. There are 1 figure and 8 Soviet references. Nauchno-issledovatel'skiy radiofisicheskiy ASSOCIATION: institut pri Gor'kovskóm universitete (Scientific Research Radiophysics Institute at Gor!kiy University) February 28, 1960 SUBMITTED: Card 8/8

S/141/61/004/006/009/017 E192/E382

9.3960 (1067, 1139, 1159)

THORS: Helakhov, A.N., Nikonov, Y.M. and Razina, T.D.

TITLE: Some methods and results of measurements of

amplitude- and frequency-fluctuations in oscillators

PERIODICAL: Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika, v. 4, no. 6, 1961, 1052 - 1064

TEXT: Two methods of measurement of the spectral density of frequency fluctuations are known (Ref. 2: D. Middleton - Trans. IRE, ED-1, 56, 1954; Ref. 3: I.L. Bershteyn, Izv. AN SSSE, ser.fis., 14, 145, 1950). The methods are discussed and evaluated and one of them is employed to measure the parameters of an experimental oscillator. In general, the measurement of the fluctuation spectra in an oscillator is based on the system illustrated in Fig. 1, which consists of: 1 - a discriminator; 2 - detector and 5 - analyser. The quasi-

chromatic signal applied to the input of the discriminator is in the form: $x(t) = A_0 \left[1 + \alpha(t)\right] \cos \left(\omega_0 t + \sqrt{1 + \alpha(t)}\right)$ (1.1)

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Some methods and results of

where $\alpha(t)$ and $\gamma(t)$ are stationary random processes having a cross-correlation function $\Phi_{\alpha \sqrt{2}}(t) = \alpha(t) \sqrt{(t+\tau)}$, such that $\bar{\alpha} = 0$, $\bar{\beta} = 0$, $\alpha^2 << 1$ and $\gamma^2 << \omega_0^2$. One of the methods of measurement is based on a discriminator containing a tuned circuit; the second method employs a delay line in the discriminator. The basic function of the discriminator consists of converting the frequency modulation of the input signal into amplitude-modulation of the output signal. The voltage at the input of the detector is therefore in the form

$$y(t) = B_0 [1 + \beta(t)] \cos \left(\omega_0 t + \int_{0}^{\infty} J_1(t) dt \right)$$
 (1.2)

The relative amplitude fluctuations $\beta(t)$ in the signal y(t) are linearly dependent on $\alpha(t)$ and $\gamma(t)$, so that the general expression for the spectral density of the fluctuations $\beta(t)$ can be expressed as:

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Some methods and results of

$$W_{\beta}(\Omega, n) = K_{\alpha}(\Omega, n) W_{\alpha}(\Omega) + K_{\gamma}(\Omega, n) W_{\gamma}(\Omega) + K_{\alpha\gamma}(\Omega, n) W_{\gamma}(\Omega)$$

$$+ K_{\alpha\gamma}^{o}(\Omega, n) W_{\alpha\gamma}^{o}(\Omega) + K_{\alpha\gamma}^{1}(\Omega, n) W_{\alpha\gamma}^{1}(\Omega)$$

$$(1.3)$$

where $W_{\mathbf{x}}(\Omega)$ is the spectral density of the signal $\mathbf{x}(t)$ at the frequency Ω , \mathbf{n} is a certain parameter dependent on the setting of the discriminator, $\mathbf{K}(\Omega_{-},\mathbf{n})$ are frequency characteristics of the discriminator and $\mathbf{W}_{\mathbf{a}\mathbf{y}}^{0}(\Omega)\mathbf{W}$; (.) are mixed spectral densities. The detector is followed by a filter which only passes a frequency lower than \mathbf{w}_{0} . The output signal of the filter contains a mean component $\bar{\mathbf{z}}$ and fluctuations $\Delta \mathbf{z}(t)$, which are preportional to $\beta(t)$. If it is assumed that the detector does not introduce any frequency distortion, the spectral density of the useful signal $\Delta \mathbf{z}(t)$ at the output of the filter is:

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Some methods and results of

$$W_{\Delta z}(\Omega, n) = z^2 \ell^2 W_{\beta}(\Omega, n) \qquad (1.4)$$

which is applied to the spectrum analyser. The symbol ℓ in Eq. (1.4) is a multiplier, determined by the type of detector. For a linear detector $\ell=1$ and for a square detector $\ell=2$. If a tuned circuit is used in the discriminator, the quantity $\beta(t)$ can be expressed by (Ref. 5; G.S. Gorelik, G.A. Yelkin - Radiotekhnika i elektronika, 2, 28, 1957).

$$\beta + 2\delta\beta + \lambda^2\beta = \lambda^2\alpha + \delta\alpha - \eta V \qquad (2.1)$$

This equation is employed to determine the amplitude, frequency and frequency-amplitude fluctuations over $\beta(t)$. In the case of a discriminator based on a delay line, the quantities $\beta(t)$, $\alpha(t)$ and $\gamma(t)$ are functionally related as follows (Ref. 5 and Ref. 6: V.S. Troitskiy - Radiotekhnika i elektronika, 1, 818, 1956):

Card 4/# 8

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Some methods and results of con-

 $\beta(t) = \frac{1}{1+k^2+2k\cos\varphi_0} \left\{ \alpha(t) + k^2\alpha(t+\tau_0) + k\left[\alpha(t) + \alpha(t+\tau_0)\right]\cos\varphi_0 - k\sin\varphi_0 \Delta\varphi \right\}$ (3.1).

Again, the expressions for the spectral density of $\beta(t)$ are derived on the basis of Eq. (3.1). The sensitivity of the measurement equipment of either type depends on the internal noise of the equipment. The noise is primarily produced by the detector and by the analyser. The detector noise consists of flicker and shot noise. The minimum detectable amplitude-spectral density and frequency-fluctuation density are determined by the equipment noise and it is shown that these quantities can be expressed by:

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APPROVED FOR RELEASE: Tuesday, August 01, 2000

3)222 **5/141/61/004/006/009/01**7

Some methods and results of

$$\mathbf{W}_{\alpha}^{\bullet V, \, \square} (\Omega) = \frac{10^{-7} - 10^{-10}}{\Omega} \tag{4.8}$$

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 $\mathbf{w}_{\mathbf{v}}^{\text{MAH}}(\Omega) = \frac{1}{\kappa_{\mathbf{v}}(\Omega, \mathbf{n})} \frac{10^{-7} - 10^{-10}}{\Omega}$ (4.9).

By analysing these formulae, it is found that the sensitivity of the two methods is identical if the equivalent quality factor of the tuned circuit is given by:

$$\mathbf{Q}_{\mathbf{K}} = \mathbf{Q}_{\mathbf{K}} = \mathbf{n} \mathbf{Q}_{\mathbf{M}} = \mathbf{ccs} \, \mathbf{\Psi}_{\mathbf{0}} \mathbf{v}_{\mathbf{0}} \mathbf{v}_{\mathbf{0}} \tag{4.12}$$

The tuned-circuit method was employed to investigate the fluctuations in an oscillator operating at 100 kc/s and an oscillator of 1.25 Hc/s. Some of the results are illustrated Card 6/

33222 S/141/61/004/006/009/017 E192/E382

Some methods and results of ... E192/E)02 in Fig. 5, where W_{β} is plotted as a function of n . The experiments showed that the relative width of the spectral line of the first oscillator was 10^{-0} when an exide-cathode tube was employed and 10^{-7} when the oscillator was based on a tungsten cathode; The corresponding figures were

5 x 10⁻⁷ and 10⁻⁷ for the oscillator operating at 1.25 Mc/s.

It is concluded, therefore, that a substantial portion of the spectral line width in the oscillator is due to the flicker noise spectral line width in the oscillator is due to the flicker noise spectral line width in the oscillator is due to the flicker noise spectral line width in the oscillator is due to the flicker noise spectral line width in the oscillator is due to the flicker noise spectral line width in the oscillator can be eliminated by of the tubes; this fluctuation component can be eliminated by of the tubes; this fluctuation cathodes. The authors thenk employing tubes with tungsten cathodes. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. There are 6 figures, I.L. Bershteyn for making useful criticism. The

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\$/141/61/004/006/009/017

Some methods and results of E192/E382

ASSOCIATION: Nauchno-issledovatel'skiy radiofizicheskiy

institut pri Gor'kovskom universitete

(Scientific Research Radiophysics Institute

of Gor'kiy University)

SUBMITTED:

February 8, 1961

Card 8/\$8

3921.6

S/141/62/005/002/008/025 E192/E382

9.4210

Nikonov. V.N. AUTHOR:

Steady-state single-frequency oscillation regime TITLE: of a magnetron oscillator

Izvestiya vysshikh uchebnykh zavedeniy. Radiofizika. v. 5, no. 2, 1962, 260 - 269 PERIODICAL:

A theoretical "model" corresponding to a singlefrequency operation of a magnetron oscillator operating continuously is constructed. This system is illustrated in Fig. 1, where the idealized "spokes" of the space charge are indicated. A cylindrical system of coordinates is used and it is shown that the single-frequency operation can be described by the following system of simplified equations:

$$\frac{dx}{dt} = \omega_0 y - \frac{\omega_0}{Q_0} x \tag{11}$$

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S/141/62/005/002/008/025 E192/E382

Steady-state single-frequency...

 $\frac{dy}{dt} = -\omega_0 x + \frac{1}{c_3} \int_V JE_{-} dv \qquad (11a)$

4

where $z_0 i_{\pi} = x$ and

y is the voltage on the equivalent capacitance C

of the resonator. Q is the equivalent quality factor of the resonator.

z is the wave impedance of the equivalent resonator

is the current density for the antiphase oscillation

The quations can be transformed into polar coordinates, in which case they can be written as:

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APPROVED FOR RELEASE: Tuesday, August 01, 2000

S/141/62/005/002/008/025 E192/E382

Steady-state single-frequency... E192/E3

$$\frac{dU}{dt} = \frac{\omega_0}{2Q_{\frac{1}{2}}} U + \frac{1}{2-C_{\frac{1}{2}}} \int_{0}^{2\pi} I_{\frac{1}{2}}(U \cos \theta) dU = \frac{1}{2}(U)$$
 (12)

$$\frac{d^{\frac{1}{2}}}{dt} = \omega_{0} = \frac{1}{2\pi C U} \int I_{H}(U \cos t, U \sin t) \sin^{\frac{1}{2}} dt = \omega_{0} + U(U) = \omega_{1}$$
(12a)

where $y = U \cos \sqrt{1}$ and $x = U \sin \sqrt{1}$ and the functions $\frac{1}{2}(U)$ and $\frac{1}{2}(U)$ are dependent on the first harmonic of the induced current $\frac{1}{2}(U)$. The induced current is calculated on the basis of determining the electron trajectories in the magnetron under the assumption that the interaction of the HF field with the space charge results in the appearance of the spokes. By determining $\frac{1}{2}(U)$ it is Card $\frac{3}{2}$

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Steady-state single-frequency.... E192/E382

easy to find the standy-state amplitude U_1 , the power generated and the frequency ω_1 of the magnetron. The dependence of the magnetron current I_0 on the oscillation amplitude is determined and the so-called limit-cycle factor p is evaluated. This factor is defined as $p = \left[\frac{\partial \psi}{\partial u}(u)/\frac{\partial u}{\partial u}\right]_{u=U_1} > 0$ and it indicates

the stability of the steady-state amplitude of the oscillations. The results of the analysis are compared with the experimental data and it is concluded that: a) the magnetron in the single-frequency regime can be adequately described by an oscillatory system with one degree of freedom, which is described by

$$\frac{d^2 \mathbf{i}_{11}}{dt^2} + \omega_{11}^2 = -\frac{\omega_T}{q_{11}} \frac{d\mathbf{i}_{21}}{dt} + \omega_T^2 \int_{\mathbf{V}} \mathbf{j} \mathbf{E}_{11}(\mathbf{r}) d\mathbf{v}$$
 (5)

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Steady-state single-frequency E192/2382

b) the steady-state value of the current I_o is proportional to the square of the oscillation amplitude; c) the magnetron current and the induced currents depend primarily on the changes of the space charge in the spokes; d) when determining the parameter p, it is necessary to take into account the radial motion of the electrons. The analysis is by no means exact or complete but it is useful in describing some of the important aspects of the magnetron operation. There are 2 figures.

ASSCCIATION:

Nauchno-issledovatel'skiy radiofizicheskiy

institut pri Gor'kovskom universitete

(Scientific Research Radiophysics Institute of

Gor'kiy University)

SUBMITTED:

July 10, 1961

Card 5/05

39247 5/141/62/005/002/009/025

9.4210

Amplitude and frequency fluctuations of a magnetron Wikonov, V. W. TITLE:

oscillator

Izvostiya vysshikh uchobnykh zavedeniy. Radiofizika, v. 5, no. 2, 1962, 270 - 286 PERIODICAL:

The equations derived in the preceding article in this issue of the journal, pp. 260 - 269, are employed to investigate the fluctuations of the amplitude and phase in a single-frequency magnetron. The fluctuation sources are the thermal fluctuations in the equivalent resonant circuit of the oscillator and the shot noise of the electron beam, The thermal fluctuations are taken into account by:

$$\frac{dx}{dt} = \omega_0 y - \frac{\omega_0 x}{Q_y} + \omega_0 \xi_1(t)$$
 (2)

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Amplitude and

where $\xi_1(t)$ is the e.m.f. developed in the equivalent resonant circuit by the resistance R_j . On the other hand, the shot noise is taken into account by:

$$\frac{dy}{dt} = -u_0 x + \frac{1}{c_2} I_H(y, x) + \frac{1}{c_3} \xi_H^* (t, x, y)$$
 (2a)

where ξ_H^+ is the random component of the induced current due to the shot effect in the electron beam; I_H^- in Eq. (2a) denotes the induced current. Eqs. (2) and (2a) are expressed in polar coordinates and are then employed to derive the equations for the random components of the amplitude fluctuation $\alpha(t)$ and phase fluctuation $\varphi(t)$. The resulting equations can be solved by first determining the random components of the induced current. It is then possible to evaluate the correlation functions and spectral densities for the amplitude and frequency Card 2/4

S/141/62/005/002/009/025 E192/E382

Amplitude and

Fluctuations. The cross correlation functions for amplitude and frequency fluctuations are also determined. The fluctuations were also measured experimentally by means of the delay-line method on magnetrons operating at 5.2 cm. By comparing the method on magnetrons operating at 5.2 cm. By comparing the experimental results with the theory it is found that the theory gives an adequate description of the amplitude-and frequency-gives an adequate description of the amplitude-and frequency-fluctuation phenomena in the magnetron. The intensity of the amplitude fluctuations B decreases inversely proportionately to the quantity $I_0\beta^2(I_0)$, where $\beta=pQ/c_0$. The strength of the frequency fluctuations is a minimum for the operating conditions corresponding to $c_0\beta$ ranging from 0 to -10°. The amplitude and frequency fluctuations are always cross-correlated, the cross-correlation function consisting of even and odd terms. There are 5 figures.

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"APPROVED FOR RELEASE: Tuesday, August 01, 2000 CIA-RDP86-00513R001137

5/141/62/005/002/009/025

Amplitude and E192/E382

ASSCCIATION: Nauchno-issledovatel'skiy radiofizicheskiy

institut pri Gor'kovskom universitete (Scientific Research Radiophysics Institute of

Gor'kiy University)

SUBMITTED: July 10, 1961

Card 4/4

FED/ENT(1)/END(v)/.20.5. 500-. ---- - 9 UP/0141/65/008/002/0219/0228 AI-5014498 ACCESSION ME AUTHOR: Earlenghaya, S. A.; Eis. yakov, A. G.; Krotikov, V. B.; Haumov, A. I.; Riko-Boy, V. Hat Porfix yev, V. A.; Plechkov, V. M.; Strezhneva, K. M.; Troitskiy, V. 6.1 Fednacyer, L. I.; Imbyako, L. V.; Sorchina, B. F. TITUE: Cheepvation of the radio eclipse of the moon at millimeter wavelengths SCURCE: WWE. Redioficies, v. 8, no. 2, 1965, 219-228 TISIC TABLE radioastronomy, lumar echipse, brightness temperature, lumar surface "SETRANT The radio emission from the moon was measured during the eclipses of 7 In y and 30 December 1963, by a procedure in which inc antenna was restained in when the emission of a section of the expositive altitude and a mountain supportanting a temperature close to that of the surrounding air. The work was done of Crayata in Armenia 30% m. of any amino 400 mg. Friending Bray o to December. Several refinaments were introduced to correct for the variation the height of the moon during the time of the eclipse. The maximum relative cop of effective temperature was ~ 175, ~ 85, 8 t 25, 5 t 25, and 3 t 25 at wave-

7 4.5 AP5014498 ACTESIEDAN IN: 'engths 1.d, 2.1, 4.0, 7.5, and 16 am in the eclipse of 7 July and 22.5 1 2.5%, 2 1%, and 8 t 2% at wavelengths 1.2, 4.0, and 6.0 mm in the collipse of 30 December ber. The bost expensent between the observation data and the trearstically preted concrete of the redio brightness tempressive twin to tellipse, for a bomogeneous model of the about 18 r/h = (6 t 1.5 and 1.0) x 10 . Y = (kpc) / (k-thermal combined wity, o--density, o--specific heat, b--tangent of dislectric loss angle the lumbs material). This value of Y/b agrees with previously obtained value measured by a different method. "We thank the Director of the Institute of Physics, Armenies Amademy of Sciences, L. I. Alikhanyas for the opportunity of performing the work on the high-mountain wase of the Institute and for help. [02] : figures and I table. ABSOCIATION: Bauchno-issledsvatel'skiy radiofisicheskiy institut pri Ger'kovekom waiversitete (Radiophysics Scientific Research Institute at the Cor kiy University) SUB CODE: AA, EC ZRC LI 00 00 RUBYLT ITO ATD PIESE: 6029 OTHER! OOM 005 RHIP HOW ard 2/2

HIKOSOV, V.P.; SMOKTIY, L.Ya.

Reduction of sulfates and selenites of alkaline earth metals by hydrogen. Ehur. neorg. khim. 5 no.8;1899-1900 Ag '60.

(MIMA 13:9)

(Alkaline earth sulfates) (Alkaline earth selenites)

3,5150

S/169/62/000/003/063/098 D228/D301

AUTHOR:

Nikonov. V. Y.

TITLE:

Astronomic methods of studying the nocturnal and the diurnal transparency of the atmosphere (Theses)

PERIODICAL:

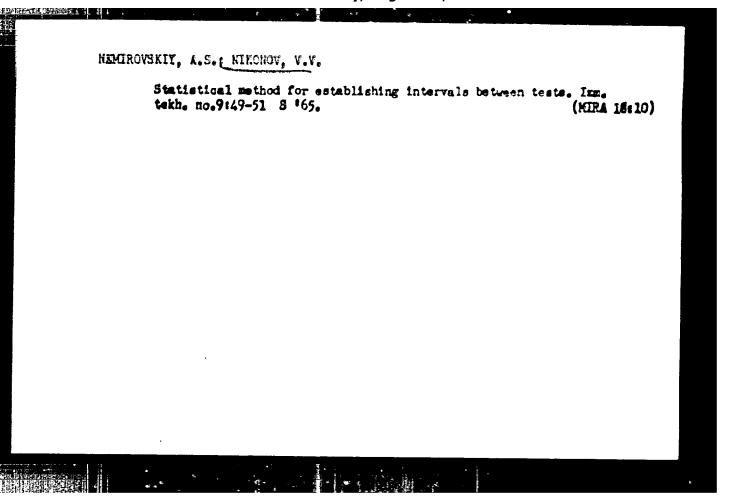
Referativnyy zharnal, Geofizika, no. 3, 1962, 29, abstract 3B235 (V sb. Aktinometriya i atmosfern. optika,

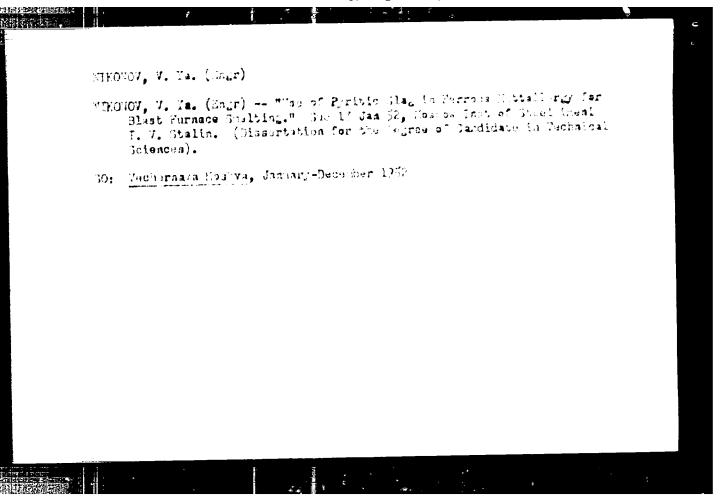
L., Gidrometeoizdat, 1961, 150)

TEXT: Instantaneous values of the nocturnal and the diurnal transparancy of the atmosphere for a certain moment of time can be obtained, if the true values of the monochromatic brilliance of stars— $m(\lambda, 0)$ — and their apparent brilliance— $m(\lambda, z, t)$ — for a certain moment of time in a certain volume of air are known. Effective procedures for determining $m(\lambda, 0)$ exist, A method of observing the apparent brightness of stars— $m(\lambda, z, t)$ — is suggested. This has been developed in astrophysics in connexion with the problem of observing the weakest stars. Abstracter's note: Complete translation.

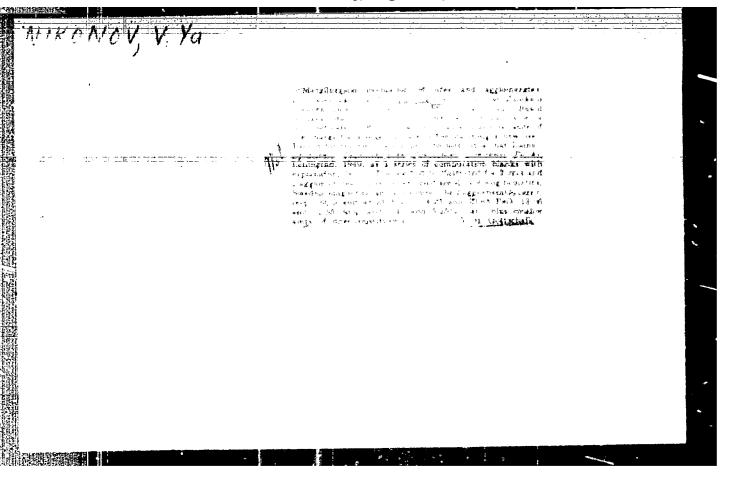
Card 1/1

VB





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Ye. Z.

137-58-4-6817

Translation from: Referativnyy zhurnal, Metallurgiya, 1958, Nr 4, p 72 'USSR'

AUTHOR: Nikonov, V. Ya.

TITLE: Hydrometallurgical Treatment of Pyrite Cinders (Gidrometal-

lurgicheskaya pererabotka pivitnykh ogarkov)

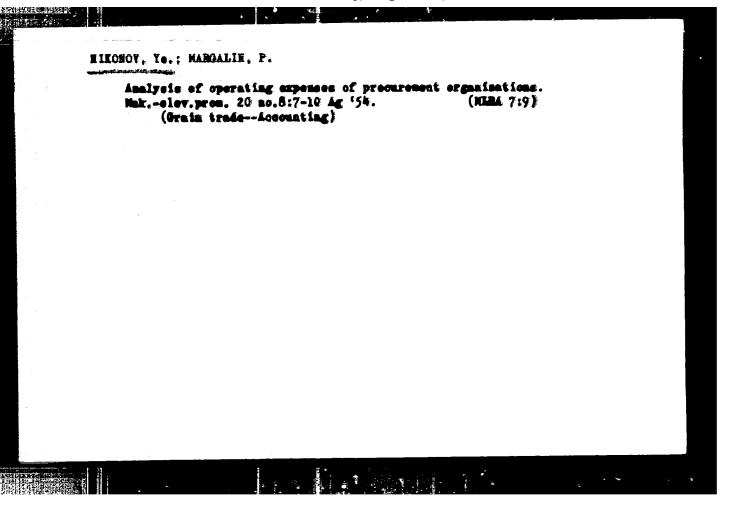
PERIODICAL: Sb. tr. Mosk. vech. metallurg. in-ta, 1957, Nr 2, pp 53-58

ABSTRACT: The degree of extraction of Cu and Zn from pyrite cinders by leaching by water or 3% H₂SO₄ may be significantly increased by preliminary sulfating roasting of the pyrites at 550-650°. 620° being the optimum temperature. A flow sheet for the pro-

cess, tested on the laboratory scale, is offered.

1. Pyrites -- Copper -- Extraction 2. Pyrites -- Zinc -- Extraction

Card 1/1



GORBITOVICH, G.D.; MEXIME, E.M.; MIMONOV, Ye.A.

Ridging drainage blocks of milled-peat fields. Terf.prom.33
no.2:11-12 '56. (MEMA 9:6)

1.Haziyeveksye terfepredpriyatiye.
(Feat industry)