

Gurevich, L. E. and Lebedinskiy, A. I. - "The pulsations of the Cepheids Part I," LEBEDINSKIY, A. I.

Astron. zhurnal, 1949, Issue 2, p. 97-103, - Bibliog: 5 items.

So: U-3736, 21 May 53, (Letopis 'Zhurnal 'nykh Statey, No. 17, 1949).

First of two articles gives a brief introduction of cepheids, also known as pulsating stars. Lists people who have studied this phenomenon. Part II will present mathematical proofs of some of the statements made in Part I. PA 42/49T4

CIA-RDP86-00513R000929110(APPROVED FOR RELEASE: Monday, July 31, 2000

LEBEDINIKIY, A. I.

FA 170T97

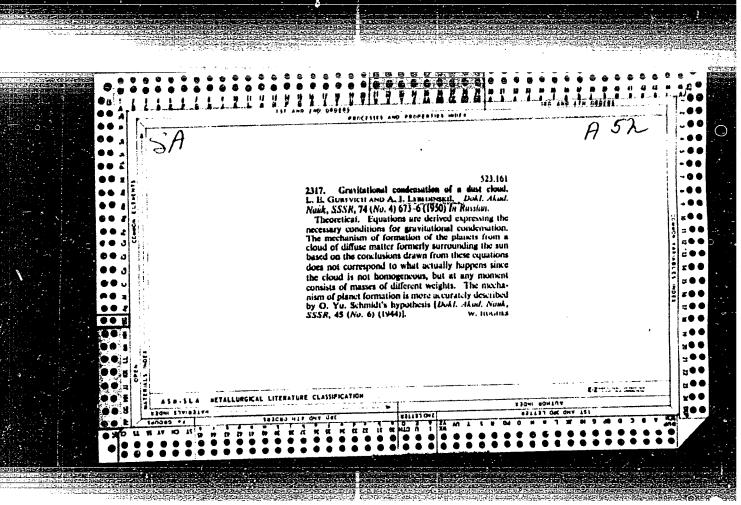
USSR/Physics - Astronomy

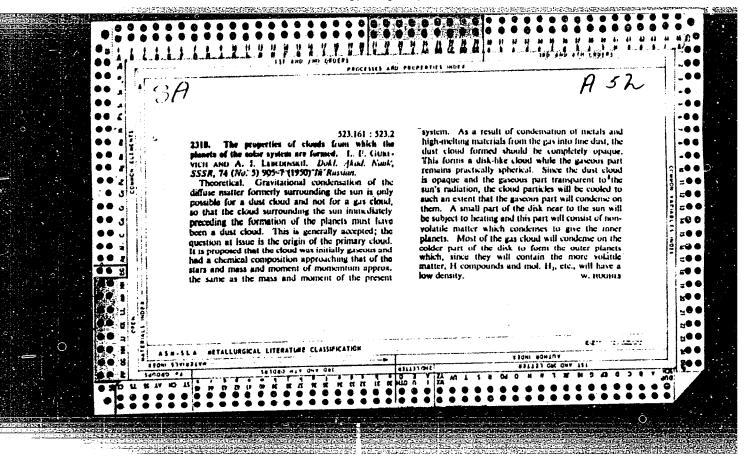
Nov/Dec 50

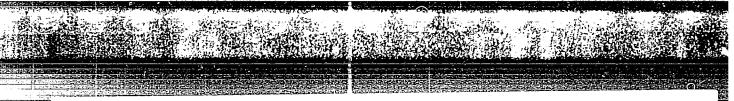
"The Formation of the Planets," L. E. Gurevich, A. I. Lebedinskiy, Leningrad State U imeni Zhdanov

"Iz Ak Nauk SSSR, Ser Fiz" Vol XIV, No 6, pp 765-803

Three reports delivered at meeting of the Depart of Physicomath Sci, Acad Sci USSR, 21 - 23 Jun 50 in Riga: "I. Gravitational Condensation," "II. Law Governing Distances of Planets and Their Rotation," and "III. Structure of the Primordial Cloud and Separation of the Planets Into the Inferior and Superior Planets."







LEBEDINSKIY A. I.

USSR/Astronomy - Planets

21 Oct 50

"Law Governing Planets' Distances and Rotation," L. E. Gurevich, A. I. Lebedinskiy, Leningrad State U imeni Zhdanov.

"Dok Ak Nauk SSSR" Vol LXXIV, No 6, pp 1061-1064

Derives general "Bode" formula for distance of nth planet to the Sun, which involves av eccentricity e:

 $R_n = \frac{1 + e}{1 + 2e} R_1 (1 + 2e)^n$

Submitted 25 Jul 50 by Acad O. Yu. Shmidt.

17272



USSR/Astronomy - Cosmogony

1 Jul 5

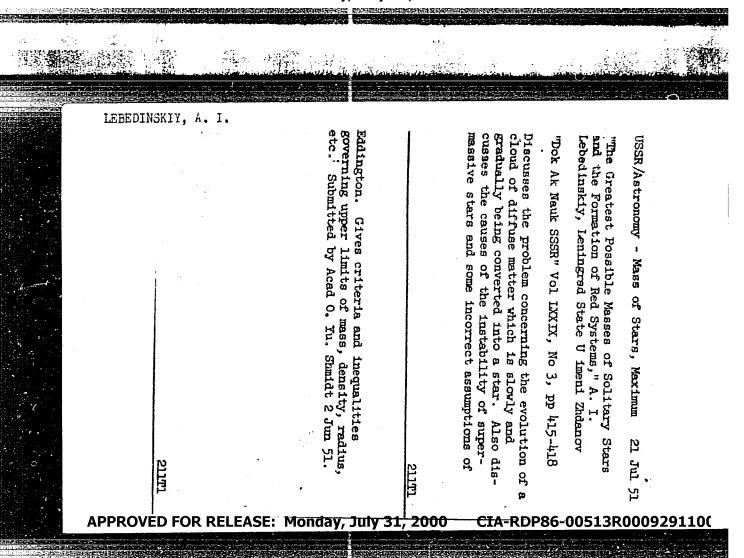
"Hypothesis Concerning the Formation of Stars," A. I. Lebedinskiy, Leningrad State U imeni A. A.

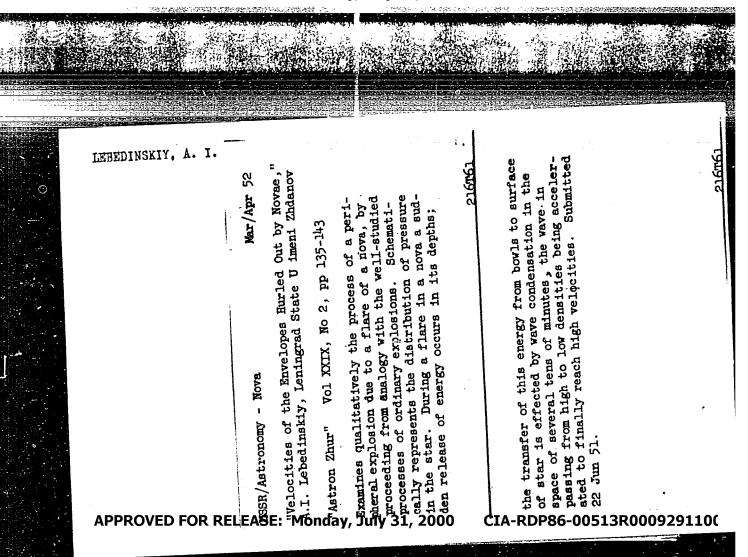
"Dok Ak Nauk SSSR" Vol LXXIX, No 1, pp 41-44

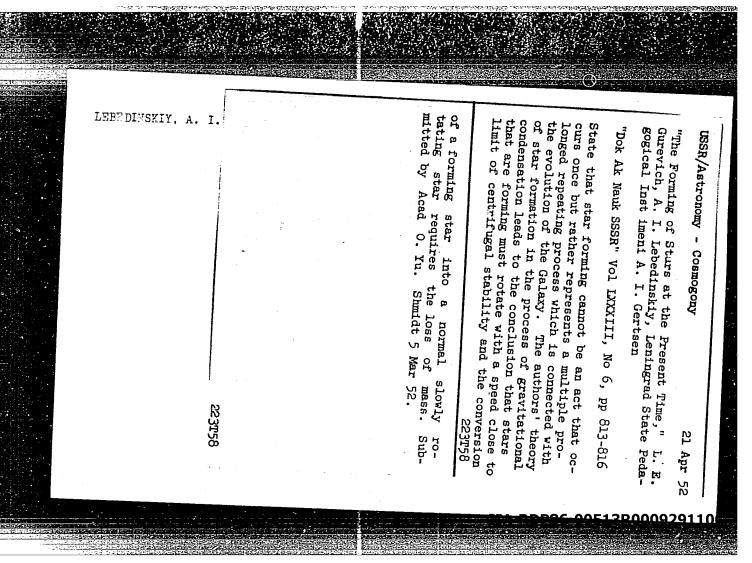
Author bases his hypothesis on Jeans' inequality const, A = gas const, \(\mu = \text{mol weight.} \) If big masses condense they may sep into smaller ones, each satisfying Jeans' inequality. Author attempts to find limit of max mass condensing into a star. He tentatively assumes max mass of hydrogen star equal to 100 solar masses. Submitted 3 May 51 by Acad 0. Yu.

LEASE: Monday, July 31, 2000

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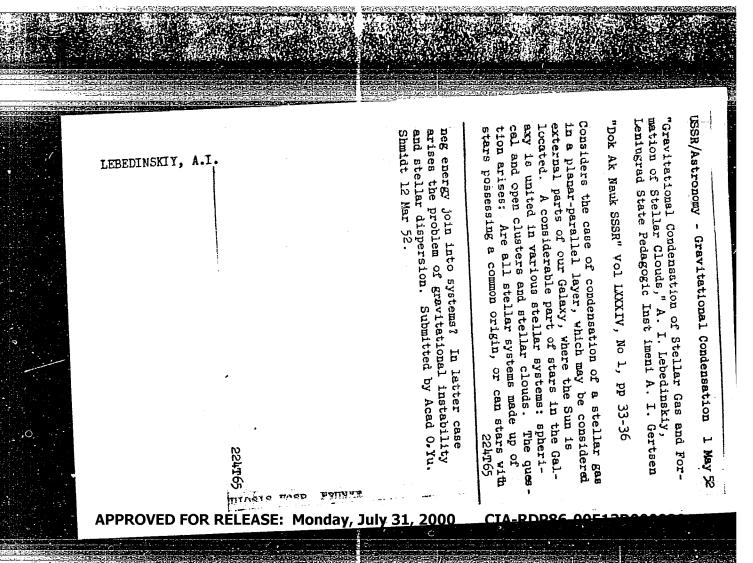


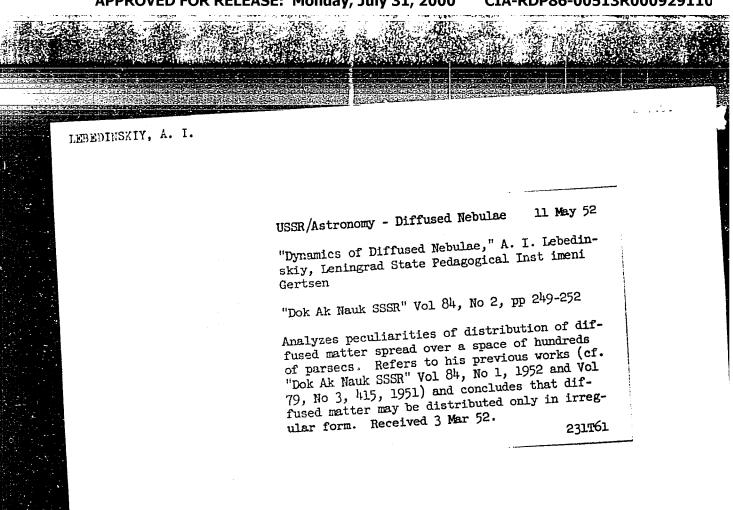
LEBEDINSKIY, A. I., and GUREVIEH, L. Te.

Spirals develop into elliptic Salaxis. II All-Union Conference of Conceguny May 1952. V. 8 (52) 99-102.

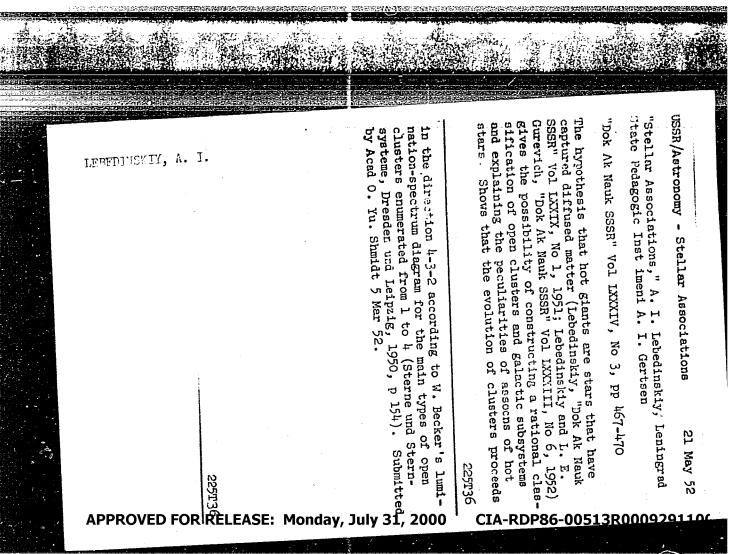
"APPROVED FOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R000929110





APPROVED FOR RELEASE: Monday, July 21





LEBEDINSKIY, A. I.

PA 245T31

USSR/Geophysics - Northern Lights

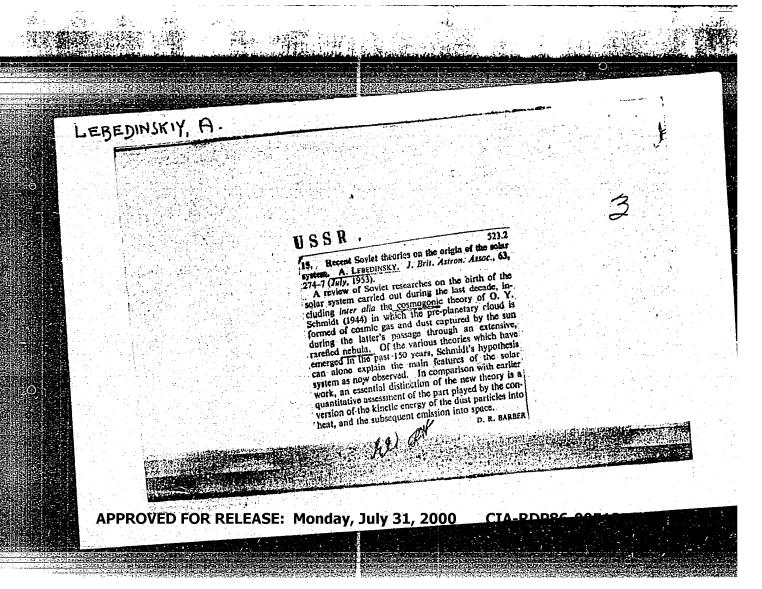
11 Oct 52

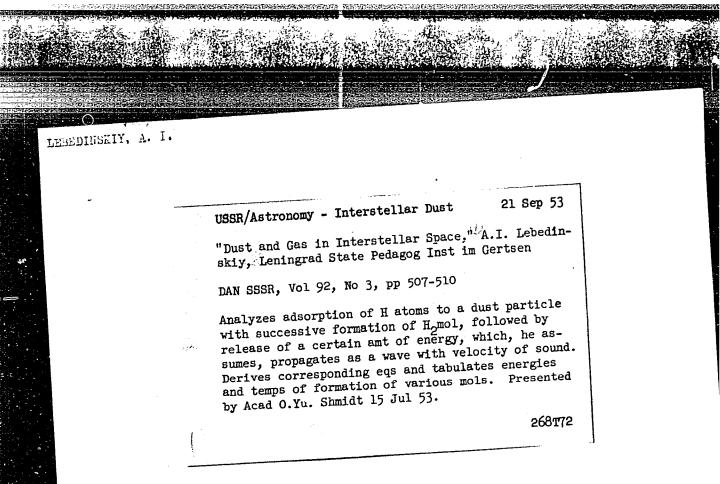
"Radiant and Arc Forms of Northern Lights," A. I. Lebedinskiy, Leningrad State Pedagogical Inst imeni Gertsen

"Dok Ak Nauk SSSR" Vol 86, No 5, pp 913-916

Assumes that northern lights are produced in the upper atmosphere, which consists of plasma. Radiative auroras are generated by electron impact and are strong in luminosity, while arc-shaped lights are due to excitation by ions and are weak in visibility. Submitted by Acad O. Yu. Shmidt 11 Aug 52.

245T31

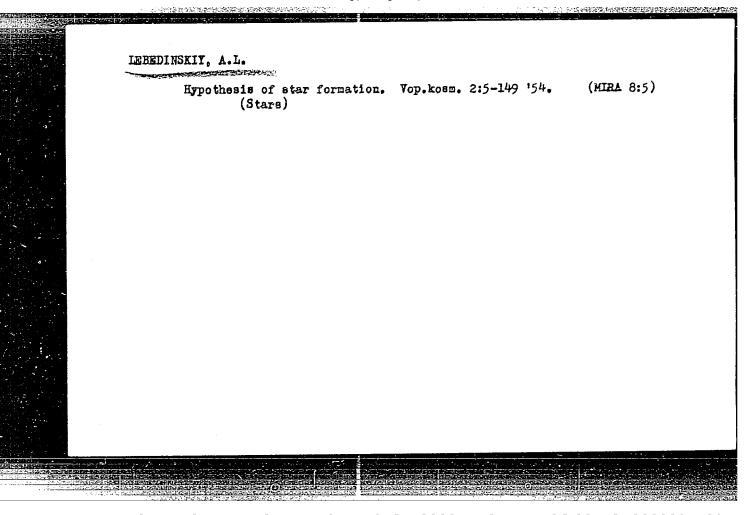






LEBEDINSKII, A.I.

RT-952 (Dynamics of diffuse nebulae) Dinamika diffuznykh tumannostei. DOKLADY AKADEMII NAUK SSSR, 92(5): 911-914, 1953.



"APPROVED FOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R000929110

Lebedinskiy, A. I.

USSR/ Astronomy - Slitless spectrographs

Card 1/1 Pub. 22 - 14/62

Authors : Lebedinskiy, A. I.

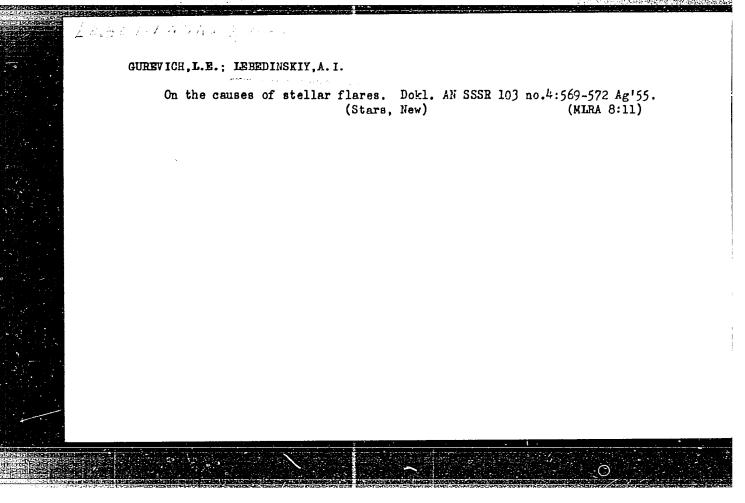
Title : Some applications of the wide-angle mirror camera

Periodical : Dok. AN SSSR 102/3, 473 - 475, May 21, 1955

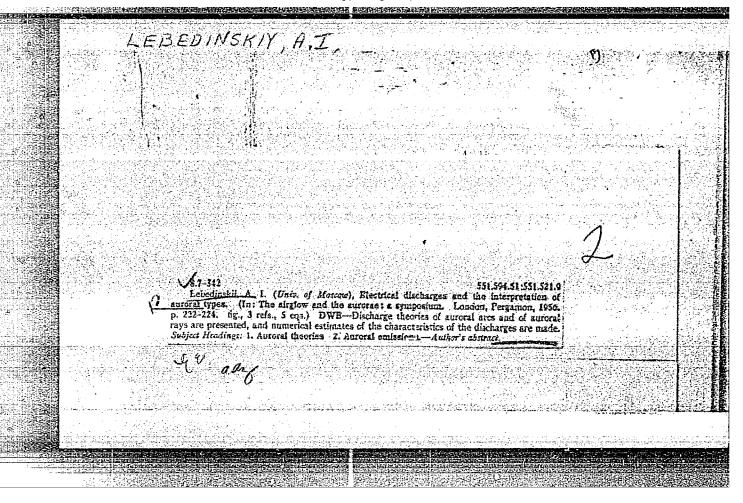
Abstract A description is given of a wide-angle camera assembly with a mirror slitless-spectrograph and its applications to taking pictures of fast changing shapes (aurora borealis, corona, bands, etc.). One Germ. reference (1951). Spectrograms.

Institution:

Presented by: Academician O. Yu. Shmidt, March 24, 1955



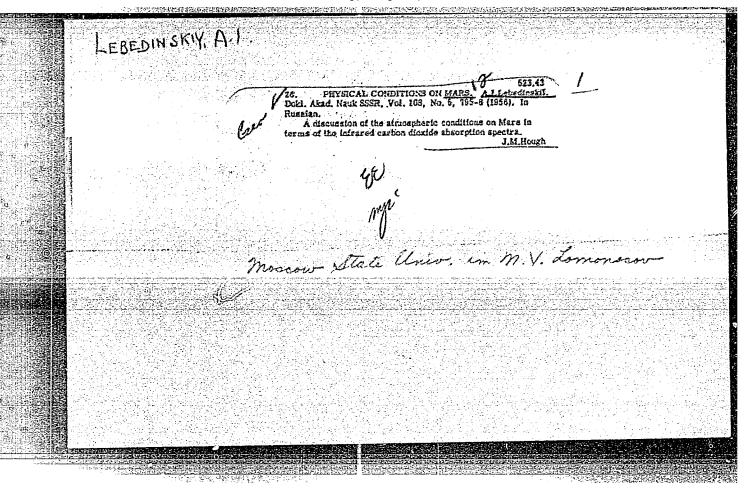
"APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110



LEBEDINSKIY, A.I.; KHOROSHEVA, O.V. Motion of stars in associations. Astron.zhur.33 no.1:54-61 (MLRA 9:6)

Ja-F 156.

1. Moskovskiy gosudarstvennyy universitet imeni M.V. Lomonosova. (Stars)



SHMIDT, Otto Yul'yevich, akademik: LEBRDINSKIV, A.I., otvetstvennyy redaktor; KOZLOVSKAYA, S.V., redaktor izdatel'stve; RYLINA, Yu.V. tekhnicheskiy redaktor

[Four lectures on a theory of the origin of the earth] Chetyre lekstsii o teorii proiskhozdeniia zemli. Izd. 3-e. dop. [Moskve] Izd-vo Akad.nauk SSSR, 1957. 138 p. (MIRA 10:9) (Cosmogony)

LEBEDINSKIY, A. Name

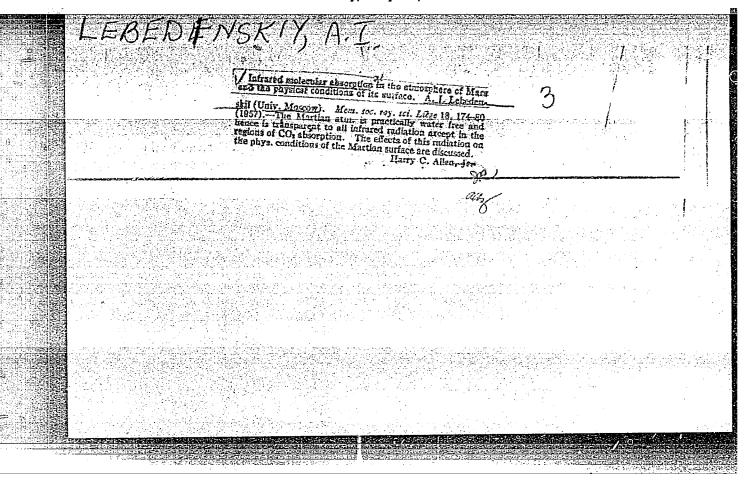
Professor Title

Remarks:

Professor A. Lebedinskiy has written an article,
"Triumph of the Human Mind", in connection with the
October 4, 1957 launching of Sputnik No. 1, which
demonstrates considerable knowledge of upper-stratosphe-

ric phenomena.

Source: N: Promyshlenno-Ekonomicheskaya Gazeta, No. 120, 6 October 1957, p.1, c.7



ISAYEV, Sergey Ivanovich; PUSHKOV, Nikolay Vasil'yevich; LEBEDINSKIY, A.I., prof., doktor fiz.-met. nauk, otvetetvennyy red.; SAKRONKHKO, L.V.; red. izd-va; POLKNOVA, T.P., tekhn. red.

[Northern lights] Poliarnye sianiia. Zarisovki poliarnykh siianii G.N. Gamon-Gemana. Moskva, Izd-vo Akad. nauk SSSR, 1956, 111 p. (Auroras)

(Auroras)

SHMIDT, Otto Yul'yevich, akademik [deceased]; KUROSH, A.G., doktor fiz.-matem.
nauk, otv.red.toma; GRIGOR'YEV, A.A., akademik, red.; DELONE, B.N.,
red.; KALASHNIKOV, A.G., doktor fiz.-matem.nauk, red.; KOZLOVSKAYA,
S.V., red.; LEEEDINSKIY, A.I., doktor fiz.-matem.nauk, red.; LEVIN,
B.Yu., doktor fiz.-matem.nauk, red.; MAL'TSEV, A.I., red.; KHIL'MI,
G.F., doktor fiz.-matem.nauk, red.; SHEVELEV, M.I., general-leytenant, red.; POLENOVA, T.P., tekhn.red.

[Selected works; mathematica] Izbrannye trudy; matematika. Moskva,
Izd-vo Akad.nauk SSSR, 1959. 315 p.

1. Chlen-korrespondent AN SSSR (for Delone, Mal'tsev).

(Groups, Theory of)

KOGAN, Ya.T., red.-sostavitel'; ALEKSANDROV, akademik, otv.red.; KALASHNIKOV, A.G., doktor fiz.-mat.nauk, red.; GRIGOR'YEV, A.A., akademik, red.; IELONA, B.N., red.; KOZLOVSKAYA, S.V., red.; KUROSH, A.G., doktor fiz.-mat.nauk, red.; LEBEDINSKIY, A.I., doktor fiz.-mat.nauk, red.; lEVIN, B.Yu., doktor fiz.-mat.nauk, red.; MAL'TSEV, A.I., akademik, red.; KHIL'MI, G.F., doktor fiz.-mat.nauk, red.; SHEVELEV, M.I., geroy Sovetskogo Soyuza, red.; PROKOF'YEVA, N.B., red.izd-va; POLENOVA, T.P., tekhn.red.

[Otto IU1'evich Shmidt; his life and works. A collection devoted to a hero of the Soviet Union, Academician Otto IU1'evich Shmidt, 1891-1956] Otto IU1'evich Shmidt; zhizn' i deiatel'nost'. Sbornik, posviashchennyi geroiu Sovetskogo Soiuza akademiku Otto IU1'evichu Shmidtu, 1891-1956. Moskva, 1959. 469 p. (MIRA 12:12)

1. Akademiya nauk SSSR. 2. Chlen-korrespondent AN SSSR (for Delone). (Shmidt, Otto IUI'evich, 1891-1956)

LEBEDINSKY, A. I.

"The composition of terrestrial corpuscular radioaction and possible mechanisms of its origin"
Lebedinsky, A. I., S. N. Vernov, A. E. Chudakov, I. P. Ivanenko

Investigations conducted during the flights of the Soviet earth satellites and the cosmic rocket have yielded the following data on the composition of "terrestrial corpuscular radiation", that is, particles revolving about the earth:

1. In the outer zone (1), the overwhelming majority of particles are electrons of energy 20 to 100 kev. If we represent the energy spectrum of these electrons in 5. Extrapolation the region of maximum radiation in the form: N (E) 1/E , then of such a soft spectrum into the region of lower energies shows that at E=5 to lo kev, the total density of electron energy would exceed the energy density of the . For this reason, the electron spectrum in the region of magnetic field H2 /8 low energies (several kilovalts) should either have a maximum or at least be small at least in this region. The number of electrons of relatively high energies $(5 \times 10^6 \text{ ev})$ is negligibly small as compared with the number of electrons in the 20-100 Kev range. The overall energy released in the crystal of the scintillation counter from photons due to the bremsstrahlung of electrons is from two to three orders of magnitude in excess of the energy released by all high energy particles that produce an energy release in the crystal of over 4.5 x 100 ev per particle. The total energy of all electrons in the outwe zone is approximately equal to the energy released in aurorae during somewhat over one week.

report presented at the International Cosmic Ray Conference, Moscow, 6-11 july 1959

21(8) AUTHORS: Vernov, S. N., Corresponding Member, 507/20-124-5-17/62 AS USSR, Grigorov, N. L., Ivanenko, I. P., Lebedinskiy, A. I. Murzin, V. S., Chudakov, A. Ye. TITLE: A Possible Mechanism of the Production of "Terrestrial Corpuscular Radiation" Under the Action of Cosmic Rays (Vozmozhnyy mekhanizm sozdaniya "zemnogo korpuskulyarnogo izlucheniya" pod deystviyem kosmicheskikh luchey) PERIODICAL: Doklady Akademii nauk SSSR, 1959, Vol 124, Nr 5, pp 1022~1025 (USSR) ABSTRACT: By "terrestrial corpuscular radiation" the authors mean the fluxes of particles moving in the terrestrial magnetic field along closed orbits. According to the authors: opinion, the following radiation production mechanism deserves the most attention: Under the action of cosmic radiation, the earth, like any other celestial body, becomes a neutron source. The neutrons traverse the magnetic field without being disturbed as uncharged particles and attain great distances from the earth. The charged particles originating from neutron decay move in the magnetic field along the lines of force. The particle Card 1/4 in the course of time reaches the region of high geomagnetic

A Possible Mechanism of the Production of "Terrestrial SOV/20-124-5-17/62 Corpuscular Radiation" Under the Action of Cosmic Rays

latitudes, where fieldstrength increases considerably with increasing latitude. In this region the velocity vector of the particle will, as the particle approaches the earth, turn so long with respect to the vector H, until at the latitude λ_{max} the angle between the velocity of the particle and the vector H becomes equal to 90° . At this point the particle returns and begins to move in the rear direction along the same magnetic line of force, If conditions are favourable, the decay products of the neutrons may perform 10^8 and more oscillations between the northern and the southern turning point. Therefore, the intensity of the flux of these particles increases by the same amount. Experimental data indicate the existence of such a radiation. The present paper carries out a closer investigation in order to find out by what factors the intensity of these rays is determined. Calculation is followed step by step. The authors calculate the intensity of the "terrestrial

Card 2/4

A Possible Mechanism of the Production of "Terrestrial SOV/20-124-5-17/62 Corpuscular Radiation" Under the Action of Cosmic Rays

corpuscular radiation" for various heights and latitudes; the results obtained by these calculations are shown by a diagram. They lead to the following conclusions: Although the number of neutrons decaying in the earth is very small, they may cause intensive cosmic radiation. The experimentally determined intensity is by \sim 100 times lower near the equator than calculated intensity. According to experimental data there is no terrestrial corpuscular radiation in geomagnetic latitudes above 40°, but in the present paper $j(\lambda = 40^{\circ}) \sim j(\lambda = 0^{\circ})$ is obtained. This means non-agreement by more than 10 times the amount. In order to reestablish agreement with the experiment, it is useful to assume an additional flux of particles from "magnetic traps", which are particularly strong in large latitudes. This may be due to the existence of electric fields. This assumption also appears to be confirmed by the data concerning the considerable increase of perturbations of the termostrial magnetic field with increasing latitude. With increasing latitude, the interdictions imposed upon energy by Stoermer's theory

are being disobeyed to an ever-increasing extent. The

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A Possible Mechanism of the Production of "Terrestrial SOV/20-124-5-17/62 Corpuscular Radiation" Under the Action of Cosmic Rays

mechanism discussed in the present paper must apply also in the neighborhood of astrophysical objects having a magnetic field. Therefore, the investigation of this radiation in the neighborhood of planets may be a means of observing weak magnetic fields. The authors thank D. V. Skobel'tsyn for his advice and M. S. Rabinovich for discussions. There are 2 figures and 7 Soviet references.

ASSOCIATION:

Moskovskiy gosudarstvennyy universitet im. M. V. Lomonosova (Moscow State University imeni M. V.

Lomonosov)

SUBMITTED:

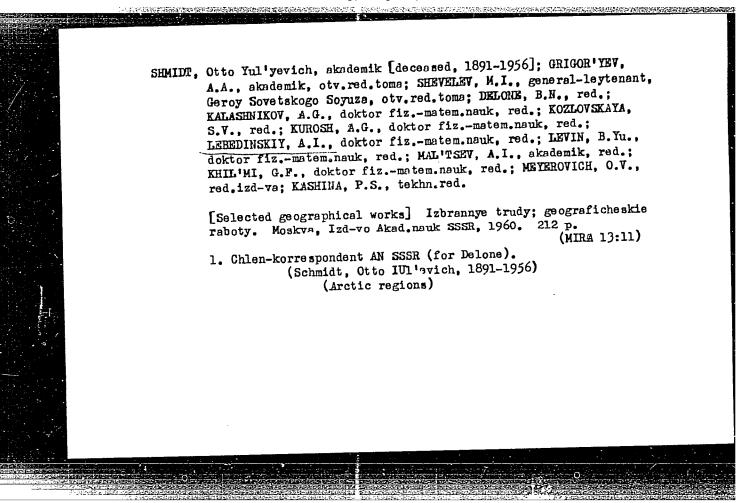
November 21, 1958

Card 4/4

SHMIDT, Otto Yul'yevich [deceased]; LERKDINSKIY, A.I., doktor fiz.-matem. nauk, otv.red.toma; LEVIN, B.Yu., doktor fiz.-matem.nauk, otv.red. toma; KHIL'MI, G.F., doktor fiz.-matem.nauk, otv.red.toma; KALASHIIKOV, A.G., doktor fiz.-matem.nauk, red.; GRIGOR'YEV, A.A., akademik, red.; DELONE, B.N., red.; KOZLOVSKAYA, S.V., red.; KUROSH, A.G., doktor fiz.-matem.nauk, red.; MAL'TSEV, A.I., akademik, red.; SHEVELEV, M.I., general-loytenant, Geroy Sovetskogo Soyuza, red.; NOVICHKOVA, N.D., tekhn.red.; KASHINA, P.S., tekhn.red.

[Selected works; geophysics and cosmogony] Izbrannye trudy; geofizika i kosmogoniia. Moskva, Izd-vo Akad.nauk SSSR, 1960. 209 p. (MIRA 14:1)

(Cosmogony) (Geophysics)
(Schmidt, Otto IUl'evich, 1891-1956)



23936 \$/035/61/000/006/018/044 ASO1/A1S1

3,2420

AUTHORS: Vernov, S.N., Chudakov, A.Ye., Lebedinskiy, A.I., Ivanenko, I.P.

TITLE: Composition of terrestrial corpuscular radiation and possible mechan-

isms of its origination

PERIODICAL: Referativnyy zhurnal. Astronomiya i Geodeziya, no. 6, 1961, 33, ab-

stract 6A287 ("Tr. Mezhdunar. konferentsii po kosmich. lucham, 1959,

v. 3", Moseow, AN SSSR, 1960, 54-58)

TEXT: Assuming that the outer radiation belt consists of electrons with energy spectrum N (> E) \sim E⁻ ?, the value of γ was determined to be \sim 5 for energies from 20 to 100 kev. Extrapolation of this spectrum to the region of lower energies (3 - 10 kev) would result in density of energy of particles which would exceed the density of energy of the constraining magnetic field. Therefore, either the spectrum of low-energy electrons must have a maximum or γ should be small. A weakening of the Earth's magnetic field was observed in the seat of a maximum filled trap. A fraction of auroras can be explained by leakage of particles from the outer belt into the atmosphere. The source of replenishment of the outer zone is solar corpuscular fluxes. At their motion, recombination is possible which gives

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23936 S/035/61/000/006/018/⁰⁴⁴ A001/A101

Composition of terrestrial corpuscular radiation ...

rise to neutral atoms. Atoms ionized about or within the radiation belt give rise to electrons captured by the magnetic field in the outer zone. The inner radiation belt is restricted by the force lines starting from the Earth's surface at latitudes $30^{\circ}-40^{\circ}$. It is filled up at β -decay of neutrons formed in the Earth's atmosphere by cosmic rays. At life time of protons in the belt being $\sim 10^{7}$ sec, the neutron mechanism is sufficient for filling up the inner zone up to intensity conserved. A sharp fall off of intensity in the belt at the geomagnetic latitude 30° can be explained by three mechanisms: a) non-conservation of magnetic moment of particles, b) effect of magnetic disturbances, c) drift of low-energy particles from the Earth as a result of being affected by ring currents presumed to exist in the ionosphere.

V. Temnyy

[Abstracter's note: Complete translation]

Card 2/2

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3,1550 (1057,1062,1129)

\$/555/60/007/000/001/007 B123/B201

AUTHOR:

Lebedinskiy, A. I.

TITLE:

Hydrogen content in large planets

PERIODICAL:

Voprosy kosmogonii, v. 7, 1960, 50-54

TEXT: An attempt has been made by the author to explain the fact that the chemical composition of giant planets differs from that of other planets. For his purposes, the author assumed heating of the outer part of the protoplanetary cloud to have been caused by the motion of solid components. A hypothesis put forth by him in 1950 explains why there is an excess of hydrogen in planets starting from Jupiter, and lack of it from Mercury to Mars. According to this hypothesis, the protoplanetary cloud was composed of hydrogen and, possibly, helium. Other elements were contained as small admixtures. Before this cloud was subjected to a high pressure, hydrogen and all its volatile compounds were in the gaseous phase. As condensation rose, dust began to concentrate in the equatorial plane, and formed a compact ring, like Saturn's, around the Sun. Because the Sun did not heat through to the outer zones of the ring, the light gases condensed on the

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Hydrogen content in

dust. The heated part gave rise to the planets of the terrestrial group on the basis of sparingly fusible materials (rocks, metals) with moderate masses. The planets of the Jupiter group were formed from light matter adding to the dust. The various chemical compositions of the Jupiter group planets can only be explained by considering the energy equilibrium of dust at the periphery, where solar radiation was practically inexistent, and stellar radiation could not heat the dust to more than 30K. Under these conditions, friction was apt to be the main source of thermal energy. The author has shown in two previous papers that the evolution of a dust system and that of a system of solid bodies exclude each other. The later paper has been published in the present periodical, Vol. 2, p. 5, 1954. When a flat disk, similar to a Saturn ring, arises, and the solid bodies act upon one another by gravitation, their volume is increased as the result of collisions of bodies of the subsystem. The equatorial diameter of the dust system is reduced by a decrease of the moment of momentum. The author considers the case where solid bodies and dust particles coexist in the subsystem. The energy of solid bodies is reduced by the friction on dust; it radiates into space as thermal energy. When bodies with comparable mass collide they are crushed in most cases, since, as a rule, their relative

Card 2/3

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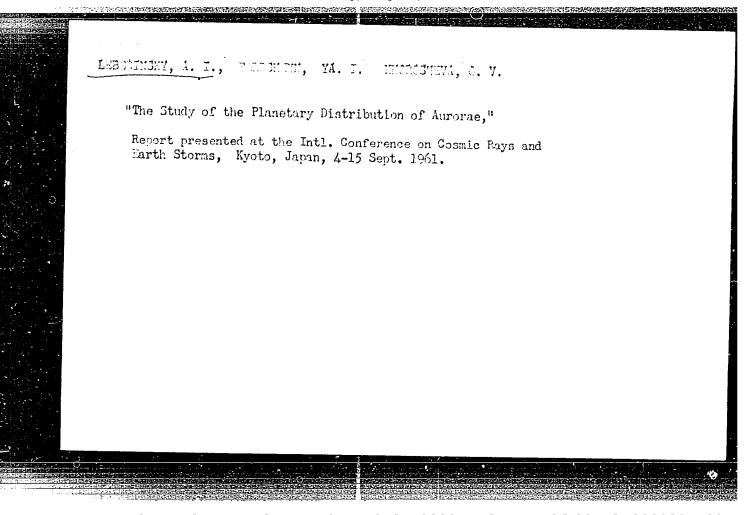
Hydrogen content in ...

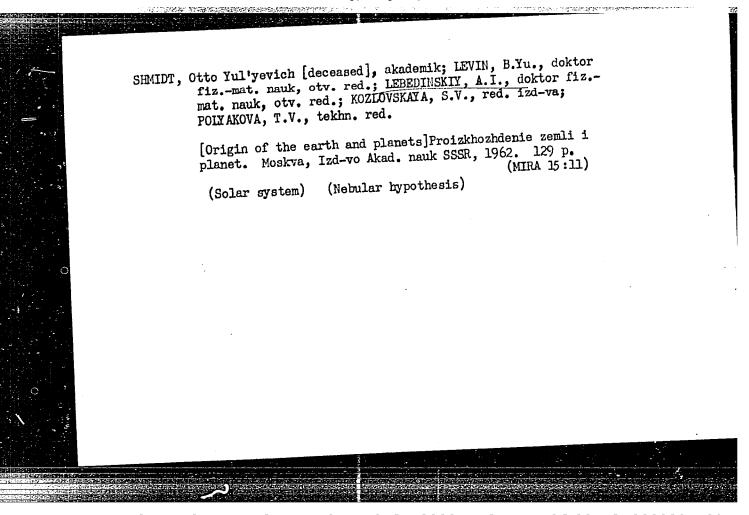
8/555/60/007/000/001/007 B123/B201

velocity is not less than several km/sec. The thus resulting small bodies go to augment the bigger bodies with which they collide, and the planetary bodies are formed on the expense of dust of the subsystem. Eventually, the disperse matter gets together to form the planets. The heat set free in the process varies from 0.1 to 100 erg/cm²sec, depending on the distance from the Sun, and corresponds to a dust temperature of 5-30°K. The process of planet formation takes place in two stages. In the first one, the solid component accumulates, and as the next stage, the chemical composition undergoes a change, once the planet has grown sufficiently massive. The gas density passed through a maximum at some distance from the Sun; this explains the maximum in the Jupiter - Saturn zone. Moreover, a monotonic dependence of the density on the orbital radius is observed with the Jupiter satellites. It is concluded that the planetary satellites as well must have formed under the same law. There are 1 table and 2 Soviet-bloc

ASSOCIATION: Moskovskiy gosudarstvennyy universitet im. M. V. Lomonosova (Moscow State University imeni M. V. Lomonosov)

Card 3/3





3,2400

AUTHORS:

J. 33/62/05 // U3/007/010

E032/E114 Lebedanskiy, A.I., and Salova, G.I.

The amount of free water on Mars TITLE:

PERIODICAL: Astronomicheskiy zhurnal, v.39, no.3, 1962, 494-505

In a previous paper it was shown that a considerable TEXT: amount of water may exist on Mars in the form of ice carried down to the surface by dust and screened by it. This idea was later developed by V.D. Davydov. In the present paper the authors consider the amount of water present on Mars in free state; i.e. in the atmosphere or the polar caps. The amount of water is determined by estimating the upper limit of the amount of ice crystals in the Martian atmosphere from its turbidity, and the rate of evaporation of ice from the polar caps. Elementary calculations indicate that the thickness of the layer of snow or cloud at the polar caps is 0.01 g/cm² and that the total amount of water in free state on Mars is 2 x 10^{15} g. It is pointed out that the calculations of Vaucouleurs and Janeau are subject to errors. The former assumed that a considerable part of the solar radiation absorbed by snow is used in evaporating it, whereas the Card 1/2

s/033/62/039/003/007/010 E032/E114

The amount of free water on Mars

present results indicate that the snow surface re-emits practically all the incident solar radiation and only a small fraction of it is used in evaporation. Jamesley on the other hand ignored the most important item in the radiation balance equation; namely the emission of the surface. The present low result for the thickness of the layer of snow, or more accurately hoar frost or clouds, is in agreement with the results obtained by A. Dollfus.

There are 3 figures and 4 tables.

ASSOCIATION: Mc. ovskiy gosudarstvennyy universitet im.

M.V. bemonosova (Moscow State University imeni

M.V. Lomonosov)

June 1, 1961. SUBMITTED:

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LEBEDINSKIY, A. I.

"Planetary Distribution of Aurorae."

report presented at the 13th Gen Assembly, IUGG, Berkeley, Calif, 19-31 Aug 63.

BAGARYATSKIY, B.A.; FEL'DSHTEYN, Ya.I.; LEBEDINSKIY, A.I., doktor fiz.-matem. nauk, otv. red.; MILYUTINA, Ye.N., red.

[Collection of articles] Sbornik statei. Moskva, Nauka. No.12. 1965. 56 p. (MIRA 18:4)

1. Akademiya nauk SSSR. Mezhduvedomstvennyy geofizicheskiy komitet. IV razdel programmy MGG. Polyarnyye siyaniya.

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	L 2176-66 FSS-2/EWT(1)/FS(s)/EWA(d) TT/GW UR/0026/65/000/009/P002/P004
-	ACCESSION NR: AP5025243
	Academician); Lebedinskiy, A. I. (Professor); Miles
	ACCESSION NR: AP5025243 AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Keldysh, M. V. (Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Khoderey, Yu. AUTHOR: Figure 1: Academician); Lebedinskiy, A. I. (Professor); Academician); Academician Aca
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AUTHOR: Lebedinskiy, A. I.; Glovatskiy, D. N.; Tulupov, V. I.; Khlopov, B. V.; Fomichev, A. A.; Shuster, G. I. TITLE: Infrared spectrophotometry of the Earth's thermal radiation SOURCE: Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. Moscow, 1965. Issledovaniya kosmicheskogo prostranstva (Space research); trudy konferentsii Moscow, Izd-vo Nauka, 1965, 65-77 TOPIC TAGS: spectrophotometer, IR spectrum, instrumentation satellite, thermal radiation, atmospheric radiation, radiation intensity, radiation spectometer/Cosmos 45 satellite ABSTRACT: Results and equipment used in an experimental study of the energy distribution of the Earth's thermal radiation are reported. A diffraction scanning spectrophotometer, mounted on Cosmos-45 comprised the basic equipment. The spectrophotometer was designed to measure thermal radiation in two bands, 7-20 µ and 14-38 µ. The spectral resolution for the first band ranged from 1.4 µ for the 7-µ wavelength to 1.1 µ for the 18-µ wavelength. For the second band, the range was from 2.8 µ for		L 2962-66 FSS-2/EWT(1)/FS(v)-3/EWA(d)	and the second s
TITLE: Infrared spectrophotometry of the Earth's thermal radiation SOURCE: Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. Moscow, 1965. Issledovaniya kosmicheskogo prostranstva (Space research); trudy konferentsii Moscow, Izd-vo Nauka, 1965, 65-77 TOPIC TAGS: spectrophotometer, IR spectrum, instrumentation satellite, thermal radiation, atmospheric radiation, radiation intensity, radiation spectometer/Cosmos 45 satellite ABSTRACT: Results and equipment used in an experimental study of the energy distribution of the Earth's thermal radiation are reported. A diffraction scanning spectrophotometer, mounted on Cosmos-45 comprised the basic equipment. The spectrophotometer was designed to measure thermal radiation in two bands, 7-20 µ and 14-38 µ. The spectral resolution for the first band ranged from 1.4 µ for the 7-µ wavelength		ACCESSION NR: AT5023566	UR/0000/65/000/000/0065/0077
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The spectral resolution for the first band ranged from 1.4 μ for the 7- μ wavelength	1		xperimental study of the energy distri
		bution of the Earth's thermal radiation are retrophotometer mounted on Cosmos-45 comprised to	the basic equipment. The spectrophoto

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the 14- μ wavelength to 2.1 μ for the 36- μ wavelength. The instantaneous field of vision of the optical system was 1°46' x 2°20', encompassing a radiating-surface area of 7.5 imes 10 km at the average altitude of 250 km. The instrument was capable of field of vision scanning within ±8°30°. Spectral intensity measurements were carried out at λ = 9.5 ± 0.6 μ for the first band and λ = 18.5 ± 1.35 μ for the second. Semiconductor bolometers with a sensitive area of 1 mm² were employed as radiation sensors. Radiation detected by the bolometers was converted into electrical signals with a frequency of 27 cps. The signals were amplified and converted into d-c voltages proportional to the radiation flux. To measure cloud cover below the satellite, a photometer operating at 6000-8000 Å with a resolution of about 30 km was used. From the data obtained during the flight of Cosmos 45, the following conclusions concerning the intensity of the Earth's thermal radiation were drawn: 1) The intensity at the minimum of the absorption band near 15 μ is almost constant. 2) A close correlation between the intensities at the other wavelengths was noted. This provides evidence that the effective radiation levels differ but slightly for various regions of the spectrum within 8-35 μ . 3) The lower layers of the troposphere are the basic source of the thermal radiation leaving the Earth's atmosphere. 4) There is a strong variable intensity of the ozone band with its center at 9.6 u. Orig. art. has: 14 fig-[GS] uresi

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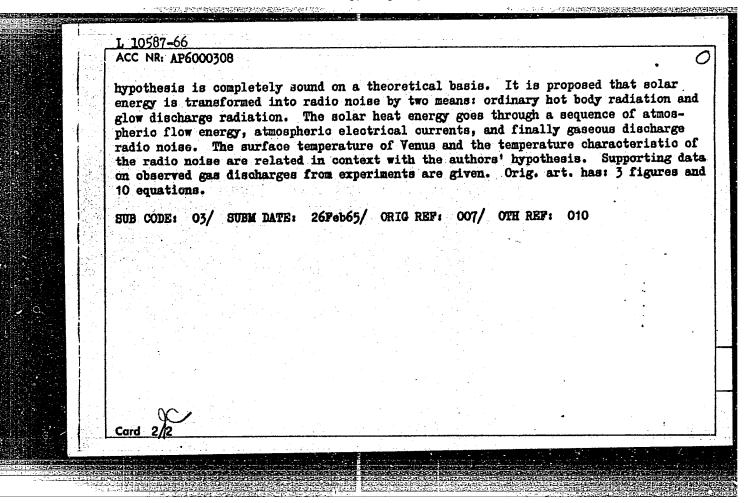
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L 10587-66 FBD/EWT(1) GW/WS=2 SOURCE CODE: UR/0293/65/003/0	006/0917/0926
AUTHORS: Vakhnin, V. W.; Lebedinskiy, A. I.	47
ORG: none TITLE: On the nature of radio noise radiation from the surface of Venus	2,55 B
SOURCE: Kosmicheskiye issledovaniya, v. 3, no. 6, 1965, 917-926	<i>\Chi</i>
TOPIC TAGS: Venus planet, cosmic radiation, cosmic radiation energy, gas radio astronomy	discharge,
ABSTRACT: The increased level of radiowave radiation from Venus (600-70 explained as "quiescent" or "glow" discharges in the upper atmosphere, or in radiated noise 200-300K above normal thermal radiation. Two hypothes advanced in explanation of the phenomenon: 1) the radiation comes from the planet which is heated by means of the "hotbed" effect in the atmosphere the planet which is heated by means of the "hotbed" effect in the atmosphere sagan, Science, 133, No. 3456, 1961), and 2) the radiation is created by Sagan, Science, 133, No. 3456, 1961), and 2) the radiation is created by Caraged particles in heated and extremely rarefied layers of the Venusian charged particles in heated and extremely rarefied layers of the Venusian (C. Sagan, op. cit. and D. E. Jones. "Planet", Space Science 5, No. 2, 19 review of some of the literature pertaining to the study of the same profile authors present and discuss some of the data obtained during the operation of the same profile authors present and discuss some of the Mariner data are insufficient several reasons are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed nor the same profile authors are given in demonstrating that neither the hotbed neither the hotb	ses are ne surface of here (see C. the motion of n ionosphere 961). A blem is given. ration of the



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ACCUSSION NRI AT5023567
AUTHOR: Lebedinskiy, A. I.; Krasnopol'skiy, V. A.; Kuznetsov, A. P.; Tozenss, V. A
TITLE: Investigation of terrestrial atmospheric radiation in the visible and visible regions
SOURCE: Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. Moscow, 1965. Issledovaniya kosmicheskogo prostranstva (Space research); truay konferentsi.
Moscow, Ird-vo Nauka, 1965, 77-88 TOPIC TAGS: atmospheric radiation, visible radiation, IR radiation, UV radiation, Topic TAGS: atmospheric radiation, visible radiation, IR radiation measurement, airglow instrumentation satellite, satellite data analysis, radiation measurement, airglow
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only on the day side of the Earth. Airglow was measured by a colorimeter (field only on the day side of the Earth. Airglow was measured by a colorimeter (field only on the day side. For switching the inswision, 120 km in width) operating only on the underlying surface, a sensor which measured course and fixing on the underlying surface, a sensor which measured four the course of the course o
illumination at 0.6 to 0.85 u was used. The colorimeter on a common axis mounted along a disk. One filter
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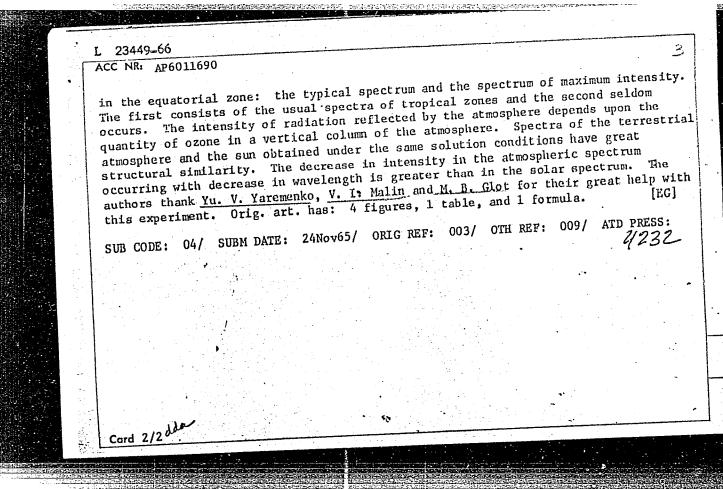
screened out UV radiation; its long-wave boundary was near 6000 Å. The second was used to investigate wavelengths at 2500 to 4000 Å, and two narrow-band filters with passbands of 100 Å filtered emission at 5577 Å and 3914 Å, respectively. An IR spectrophotometer recorded thermal radiation concurrently with the operation of the other two instruments. A correlation was found to exist between readings of the UV and IR spectrophotometers in the 9.65-µ ozone band. A correlation of intensities was also disclosed near the long-wave boundary of the UV spectrum at λ > 3000 Å and in the readings of the illumination sensor. These readings depended strongly on cloudiness because the albedo of clouds in the red zone is substantially greater than the albedo of the Earth's surface and of the clear atmosphere. The correlation con-Firms that at $\lambda > 3000$ Å, the noticeable part of atmospheric radiation is due to tropospheric dispersion and reflection occurring below the basic mass of the ozone layer. Conclusions were also reached on local, diurnal, and latitudinal variations of airglow. A difficulty arose in the evaluation because of the dependence of the readings on cloud cover. In making the measurements in space, it was necessary to include reflections of airglow from the atmosphere and glow of astronomical origin in addition to airglow itself. Consequently, results varied with atmospheric conditions by as much as a factor of two, with the minimum occurring during cloudless weather and the maximum during total cloudiness. The correlation of readings of one light filter (5577 Å) with the others indicated that the share of illumination from the stars and

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	small. A comparison of these	readings (averaged) w
zodiacal light was relatively the results of ground observat sky radiation at 2500—3000 Å	ions at 3200—4000 Å lead to	does not exceed star (
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tion exists in this regioned	the absence in the night sky	of high-energy excit
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	기가 있는 것이 되는 경기를 받는다.	
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SOURCE CODE: UR/0203/66/006/002/0185/0189 TT/GW FSS_2/EWT(1) 23449-66 ACC NR: AP6011690 AUTHOR: Krasnopol'skiy, V. A.; Kuznetsov, A. P.; Lebedinskiy, A. I. ORG: Moscow State University, Institute of Nuclear Physics (Moskovskiy gosudarstvennyy universitet Institut yadernoy fiziki) TITLE: Measurements of the ultraviolet spectrum of the earth made by the satellite "Kosmos-65" SOURCE: Geomagnetizm i aeronomiya, v. 6, no. 2, 1966, 185-189 TOPIC TAGS: solar ultraviolet radiation, ultraviolet spectrophotometer, diffractional monochromator, ozone concentration, spectral brightness, radiation intensity ABSTRACT: Solar ultraviolet radiation reflected from the terrestrial atmosphere was measured with an <u>ultraviolet spectrophotometer</u> mounted on the artificial satellite "Kosmos-65." The spectrophotometer used was a double diffractional monochromator operating in the spectral range of 2250—3070 Å. 2500 spectra were obtained during the flights. Spectra were recorded on motion-picture films. The distribution of energy in the violet spectrum changes with the zenithal distance of the sun and the geographical latitude. Local peculiarities caused by the ozone concentration appear. Longer waves in the ultraviolet range penetrate deeper into the atmosphere and increase the albedo intensity. Two kinds of spectra were obtained on cloudless days 2 523.72:629.192.2 Card 1/2



SOURCE CODE: UR/0384/66/000/001/0079/0081 (N)ACC NR: AP6019461 AUTHOR: Vakhnin, V. M. (Candidate of physico-mathematical sciences); Lebedinskiy, A. I. (Professor) ORG: none TITLE: Radio noise and the temperature on Venus SOURCE: Zemlya i vselennaya, no. 1, 1966, 79-81 TOPIC TAGS: radio noise, Venus probe, space temperature, glow discharge, rarefied gas ABSTRACT: The use of radio signals emitted by Venus to study its surface temperature is discussed. A theoretical explanation of Venus' apparently high temperature surface (considering the "hot house effect", the ionospheric hypothesis and the contradiction of this appothesis by the peculiarities of radio signals emitted from Venus) is presented. The electric glow discharge in rarefied gases in relation to the very slow speed with which the planet Venus rotates around its axis and the possible existence of high velocity global breezes which do not create disturbances are considered. It is proposed that the atmospheric current going through the upper layers of the atmosphere of the planet Venus creates a continuous glow discharge resulting in powerful radio noise and a low degree of luminescence. The proposed explanation is substantiated by experiment al data and can be explained theoretically. If this interpretation is correct, then Card 1/2

ACC NR: AP6019461

the surface temperature of Venus would be between 50-60°C. Orig. art. has: 3 figures.

SUB CODE: 03_17/

SUBM DATE: none

"APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110

L 45257-66 FEBT(1)/EEC(k)-2/FCC JKT/TT/DD/GW ACC NR: AP6020934 SOURCE CODE: UR/0029/66/000/006/0008/0011

AUTHOR: Lidov, M. L., (Doctor of Physical and Mathematical Sciences); B. Lebedinskiy, A. I., (Doctor of Physical and Mathematical Sciences, Professor); Vernov, S. N., (Corresponding Member of the Academy of Sciences SSSR)

ORG: none

TITLE: The battle for the Moon continues

SOURCE: Tekhnika-molodezhi, no.6, 1966, 8-11

TOPIC TAGS: moon, space, lunar surface, lunar radiation, lunar landing /Geiger counter, Luna 9, Luna 10

ABSTRACT: The interviewer reviews briefly the history of lunar research, presents a table of chronology and facts and repeats questions and answers. Dr. M. L. Lidov stated that one of the problems solved by Luna 9 and Luna 10 was that of landing at the most favorable time, i.e., lunar daybreak. Another problem was the selection of the most "economical" trajectory along which to send the heaviest apparatus. He stressed the importance of human initiative in the

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guidance and control of space vehicles and this spoke of a third major problem ACC NR: AP6020934 involving slowing down the space station. The braking process was controlled from Earth and was started 2 1/2 hr before the actual landing. The flight problems faced by Luna 9 and Luna 10 were similar enough but owing to a correction, Luna 10 moved to a flight trajectory along which it approached the Moon to within a minimum distance of 1000 km. Dr. A. I. Lebedinskiy discussed the special features of the lunar landscape as transmitted on television, explained the structure of the space station and its operation when taking pictures, also the topography of the moon. The objectives of Luna 10 were research on lunar space and lunar radiation. Academician S. N. Vernov, a specialist in cosmic rays, stated that the Geiger counter installed on Luna 9 recorded protons, electrons, and gamma-quanta with a strength sufficient to penetrate the hermetically scaled space station. Lunar radiation was found to be 10 times stronger than that of the Earth; it was found to be harmless to astronauts clad in space suits. Orig. art. has: 5 figures and

SUB CODEL 14, 03, 18, 22/ SUBM DATE: none/

Card 2/2

1 table.

SOURCE CODE: UR/0293/66/004/006/0838/0841

AUTHOR: Lebedinskiy, A. I.; Lozhnikov, A. A.; Tulupov, V. I.

ORG: none

TITLE: Measurements of lunar radiation flux in the infrared and visible

regions of the spectrum by the Luna-10 satellite

SOURCE: Kosmicheskiye issledovaniya, v. 4, no. 6, 1966, 838-841

TOPIC TAGS: lunar radiation, lunar satellite, lunar surface / Luna-10 lunar

satellite ABSTRACT:

The lunar radiation flux in the infrared (IR) and visible spectral bands was measured by the Luna-10 satellite from a lunar orbit. The results have not yet been fully analyzed, and the data presented by the authors are only preliminary.

Radiation in the two spectral bands was of different origin. The visible band was used to measure radiation caused by heat from the Sun being reflected by the Moon's surface; the IR band measured the Moon — radiated heat which was emitted by the Moon's surface, the temperature of which varies from -150 to +120°C. Infrared and visible radiation coming directly from the Sun also contributed to the total measured in both bands.

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UDC: 629.195.3:523.37

A detector carrying two sensors (15 x 30-mm thin plates) mounted outside the Luna-10 body were utilized to separate different heat radiation contributions (see Fig. 1). The detector action was based on the variable resistance principle (i.e., changes in heat radiation varied the detector resistance). Detector resistance was measured by a circuit (such as a resistance bridge) and the results of these measurements were telemetered back to Earth. One of the sensors was covered with enamel, which absorbed 85-95% of the incident IR radiation and reflected 70-75% of the radiation in the visible band. The other sensor was covered with thin gold foil, which reflected 97-99% of the IR radiation and passed most of the visible radiation. Data taken simultaneously from both sensors will be used to isolate that portion

of the total heat radiation which was contributed by the Moon. The satellite was rotating around its own axis with a period which was shorter than the detectors' thermal time constants, and the readings therefore represent the average values of thermal radiation. The sensors were sampled simultaneously every two minutes.

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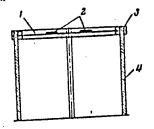
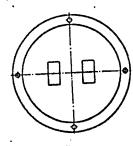


Fig. 1. Thermal radiation detector

1 - Heat insulator; 2 - sensors; 3 - clamping ring; 4 - stand.



Temperatures measured by the sensors (A, with enamel cover; B, with gold foil cover) during the 8th, 31st, and 49th measurement sessions are given in Fig. 2. In the 8th session the satellite crossed

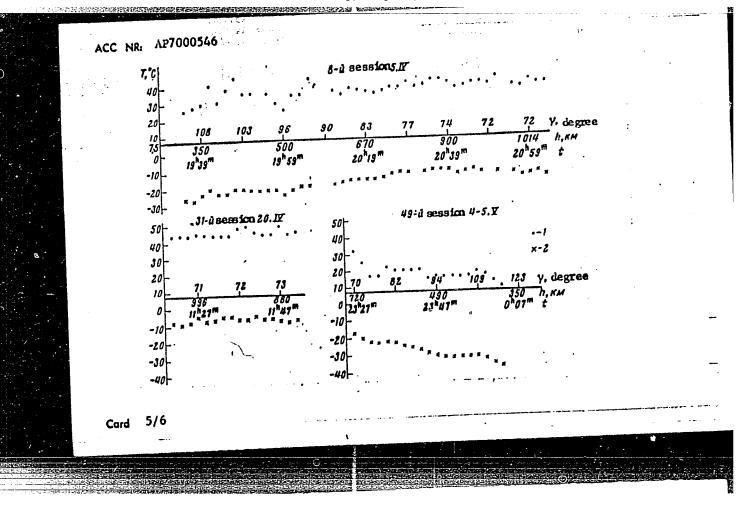
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from the dark to the sunlit portion of the Moon. During the 49th session, a crossing from the light to the dark portion was made. In the 31st measurement session the satellite was on the sunlit side of the Moon. The time marks in Fig. 2 indicate Moscow time. As expected, the temperatures on the light side of the Moon are higher than on the dark side. During crossings the sensors registered corresponding changes in temperature after a time delay associated with their thermal inertia.

The influence of thermal detector inertia was investigated by taking temperature measurements at different instants during a single session. The instants at which these measurements were taken varied from session to session. These readings are plotted and compared in the figure, illustrating the effect of thermal time delay on the sensor output.

From 2-3 May through 13 May the temperatures of both sensors declined noticeably. The A-sensor temperature decreased by 22°C and that of the B-sensor, by 50°C. After 13 May the temperatures again began to rise. The speculation that this phenomenon is associated with the changing orientation of the satellite with respect to the Sun is confirmed by the fact that the B-sensor, the more sensitive of the two to the Sun's rays, was affected more than the A-sensor.

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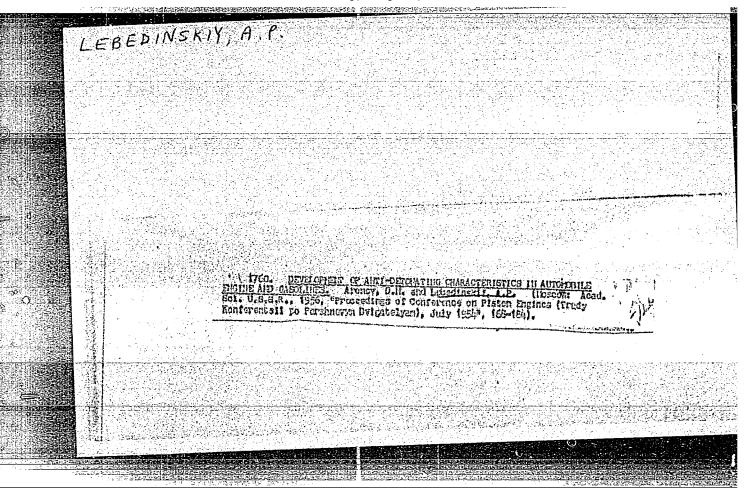
Fig. 2. Detector readings during sessions 8, 31, and 49

1 - B-sensor; 2 - A-sensor; Y - angle between a line extending from the Moon's center to the Sun and a line between the Moon's center and the Luna-10; h - satellite height.

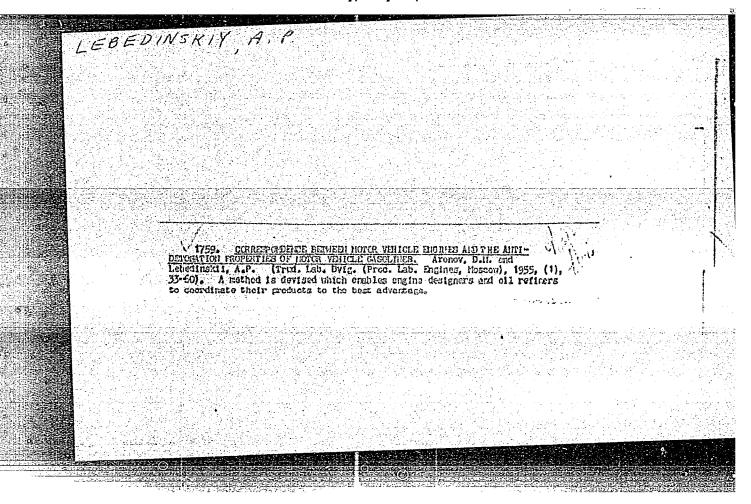
The authors state that analysis and evaluation of the results will be published in the future. The authors thank A. D. Levchenko and V. V. Vernigor for assistance in the carrying out of the experiments. Orig. art. has: 5 figures. FSB: v. 3, no. 1/SUB CODE: 22,03 / SUBM DATE: 22Aug66

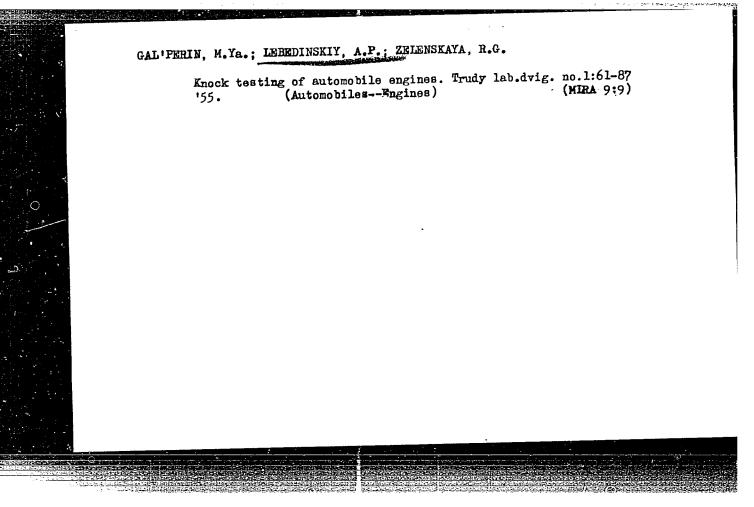
Card 6/6

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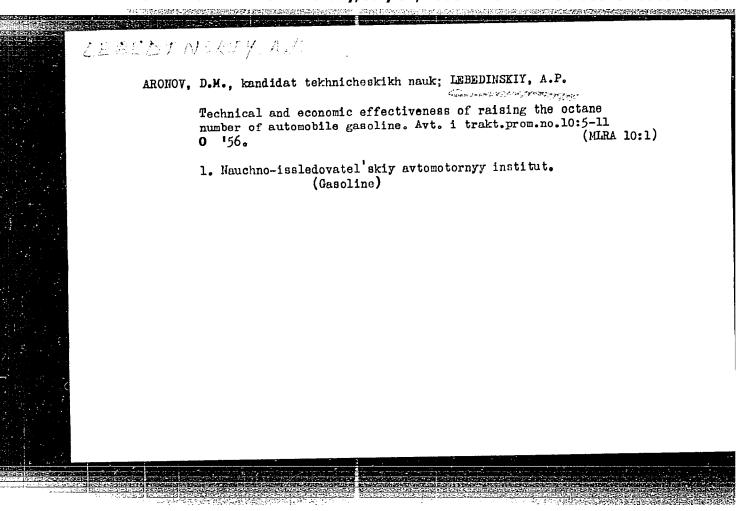


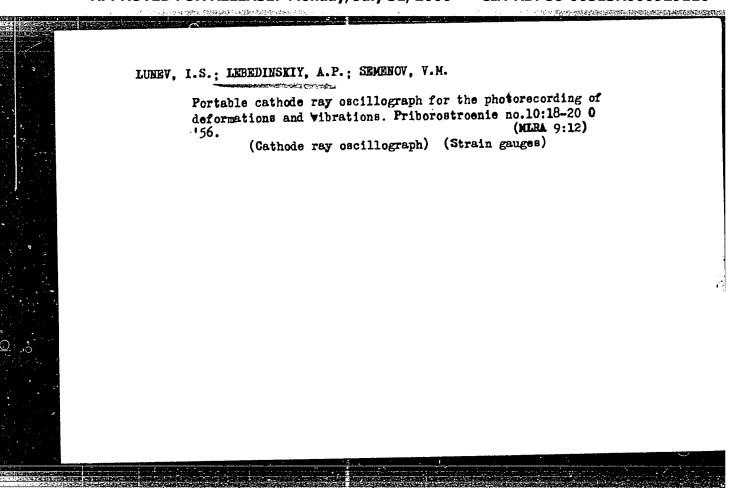
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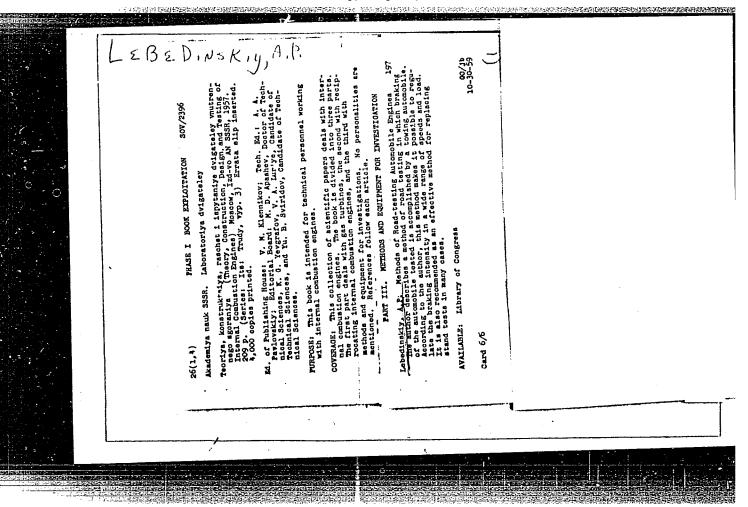


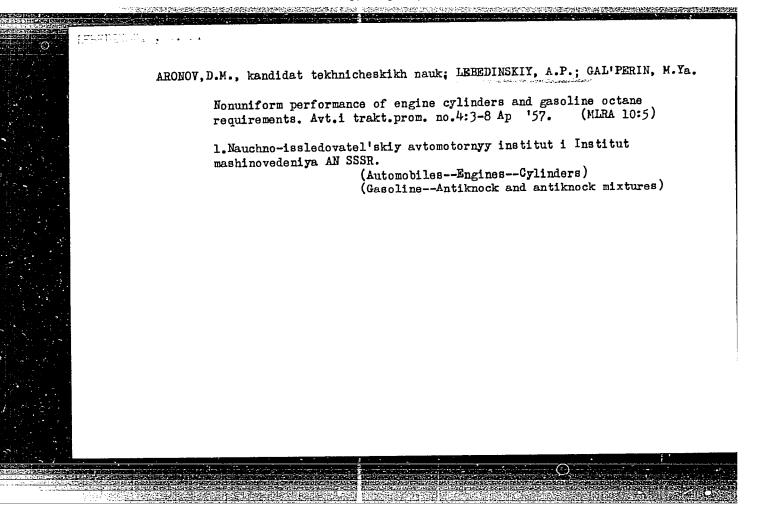


"APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513R000929110









KONEY, Boris Fedorovich; ARONOV, David Matveyevich; KUROV, Boris Alekseyevich; LEBEDINSKIY, Aleksandr Pavlovich; NILOV, N.A., inzh., retsenzent; YEGORKINA, L.I., red.; NAKHIMSON, V.A., red.; TIKHA-NOV, A.Ya., tekhn.red.; UVAROVA, A.F., tekhn.red.

[Automobile carburetor engines; characteristics and methods for their determination] Avtomobil nye karbiuratornye dvigateli; kharakteristiki i metody ikh opredeleniia. Moskva, Gos.nauchnotekhn.izd-vo mashinostroit.lit-ry, 1960. 229 p. (MIRA 13:4) (Automobiles-Engines)

LEBEDINSKIY, A.P.; GERSHMAN, I.I., kand.tekhn.nauk

Automobile engines using various fuels. Avt.prom. no.4:18-23 '60.

(MIBA 13:6)

1. Gosudarstvennyy soyuznyy ordena Trudovogo Krasnogo Znameni nauchnoissledovatel'skiy avtomobil'nyy i avtomotornyy institut.

(Automobiles--Engines)

LEBEDINSKIY, A. P.

Effect of antiknock qualities of fuel on the power and economic indices of engines. Avt. prom. 29 no.5:5-10 My '63. (MIRA 16:4)

1. Gosudarstvennyy soyuznyy ordena Trudovogo Krasnogo Znameni nauchno-issledovatel'skiy avtomobil'nyy i avtomotornyy institut.

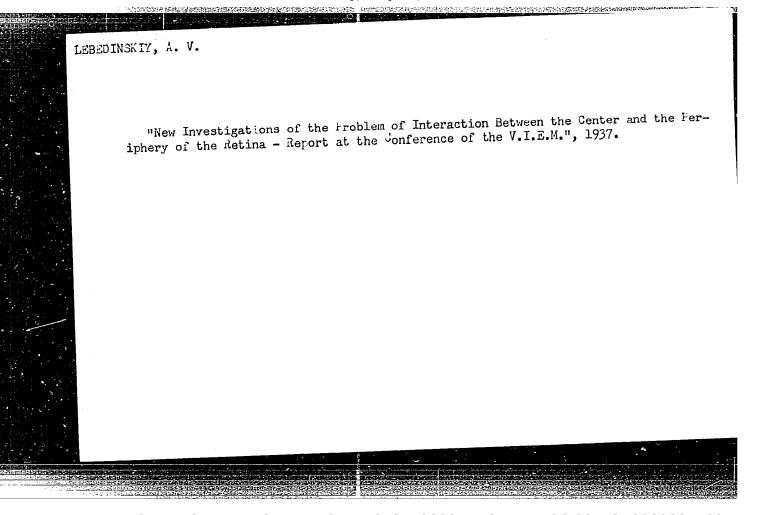
(Motor fuels) (Motor vehicled—Engines)

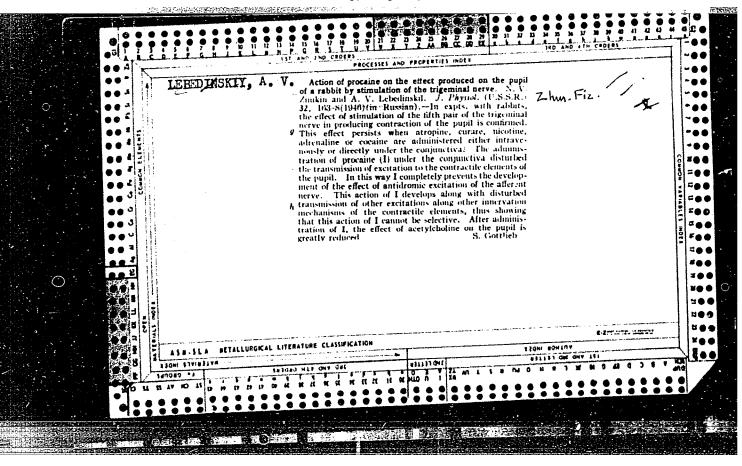
LEBEDINSKIY, A. V., DIONEZOV, S. M., ZAGURULKO, L. T. and TURTSAYEV, Zh. F.

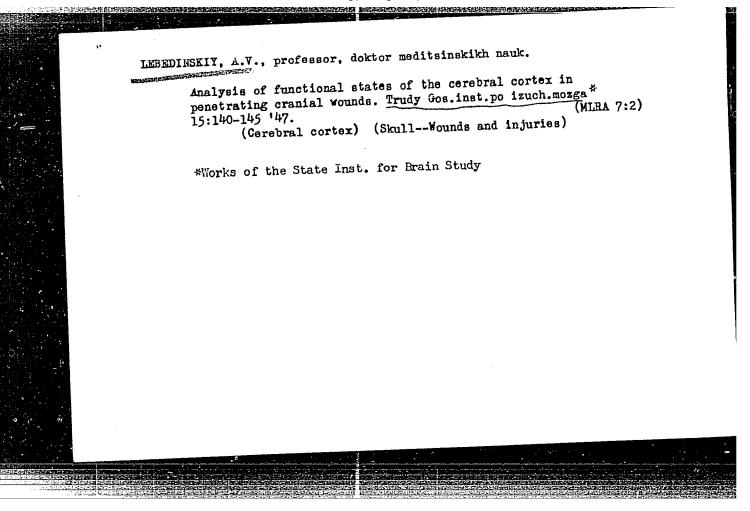
"The Effect of Physical effort on the Dark-Adaptation of the Eye", Fiziolog.
Zhurnal SSSR, Vol. 16, 5th Ed., 1933.

LEBEDIMSKIY, A. V., DIONESOV, S. M. and TURTSAYEV, Zn. F.

"The Effect of Reflex Excitations (Cold) on the Light Sensibility of the Dark-Adapted Eye", Fiziolog. Zhurnal SSSR, Vol. 17, 1st. ed., 1934.







LEBEDINSKIY, Andrey Vladimirovich (Dr.), and MIKHEL'SON, N. I.

"Reselient-Viscous Properties of the Skeletal Muscle and its Changes under the Influence of Sympathetic Enervation." Zef. Zhur., Vol 33, No h, 1947, p 505.

Influence of Physiology, Military Hed Acad imeni Kirov.

So: W-h396

LEBEDINSKIY, Andrey Vludimirovich and Savvin, N. G.

"On Relations of Sympathetic Enervation to the Reaction of Contracting Formations on Efferent Influences of Various types." Zef. Zhur., Vol33, No 6, 1917, p 719.

Chair of Physiology, Military Med Acad S. M. Korov.

So: U-1:396

LEBEDINSKIY, A. V.

USSR/Medicine-Publications Medicine-Water Supply Hov 48

nliew Books" 2 P

Lists new books, giving authors, publishing dates and prices. Includes V.F. Kozhin's "Water Supply" and A. G. Ginetsinskiy and A. V. Lebedinskiy's "Fundamentals of Human and Indiana Charles Animal Physiology."

PA 49/49T77

CIA-RDP86-00513R000929110(APPROVED FOR RELEASE: Monday, July 31, 2000

PA 3/5ULDO TEBEDINSKIY, A. V. Optical Adaptation to Darkness," A. V. "Contempory Representations of the Mechanism of Lebedinskly, Leningrad, 22 pp increase in the sensitivity to light which occurs "Uspekhi Sovrem. Biol." Vol XXVI, No. 3 (6) in weak illumination remains valid, but important details have been added and whole process of adapta. Theory that ocular adaptation is a process of Compares and analyzes author's experiments and conclusions of many other authors. tion has been examined from various viewpoints. Eye, Adaptation of 3/50155 APPROVED FOR RELEASE: Monday, July 31, 2000 CIA-RDP86-00513P0



LEBEDINSKIY, A. V.

Assoc. Member, USSR Academy of Medical Science.

"I.P. Pavlov on the Trophic Role of the Nervous System." Izvestia, 1949, Lecture

Current Digest of the Soviet Press, Vol. 1 No. 36, 1949, page 55, (In Library)

LEBEDINSKIY, A. V.

33468. K Stoletiyu So Dnya Rozhdeniya I. P. Pavlova. Voprosy Neyrokhirurgii,
1949, No. 5, c. 12-16, c. Portr.

SO. Letopis' Zhurnal'nykh Statey, Vol. 45, Moskva, 1949

LEBEDINSKIY, A. V.

USSR/Medicine - Intestines, Physiology Medicine - Intestines, Sensitivity Jan 49

"Problem of the Sensitivity of the Internal Organs" A. I. Bronshteyn, A. V. Lebedinskiy, V. M. Simenko, Chair of Physiol, Chair of Faculty Surg, Mil Med Acad imeni S. M. Kirov, 12 pp

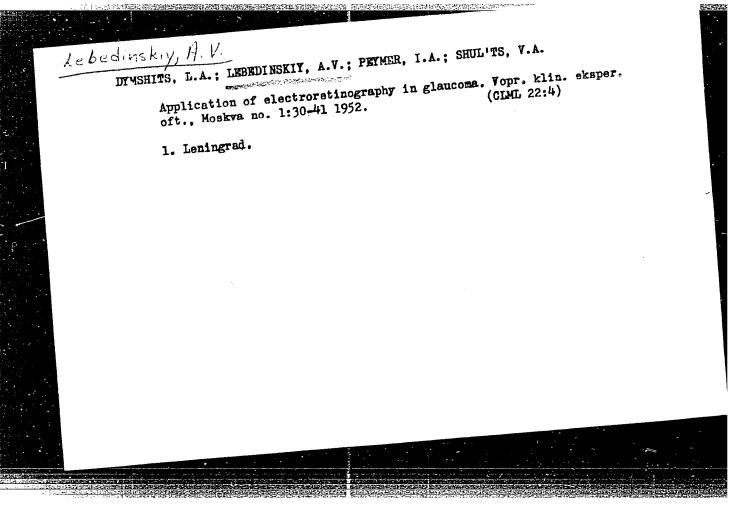
iol. Zhur. SSER" Vol XXXV, No l

Conducted tests on four animals with abnormally small intestines to determine degree of expansion before pain became noticeable. Determined that the animal sensory reactions could be observed not only when intestine was moved, but also when an electric charge acted on interior walls of the intestine. Charge produces reactions by: (1) stimulating peristalsis of the intestine (direct current), and (2) irritation of the sensory nerve in the caecum (pulsed current).

PA 47/49T62

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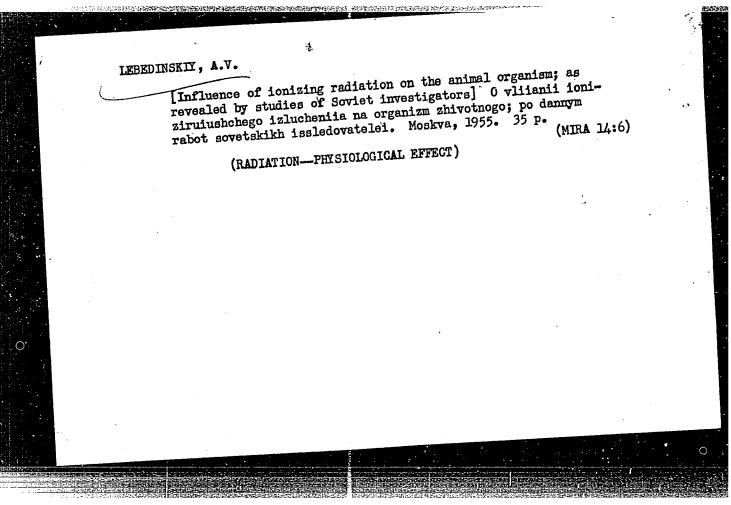
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1	LEBEDINSKIY,	A.	V.;	MOZZHUKHIN, A. S.

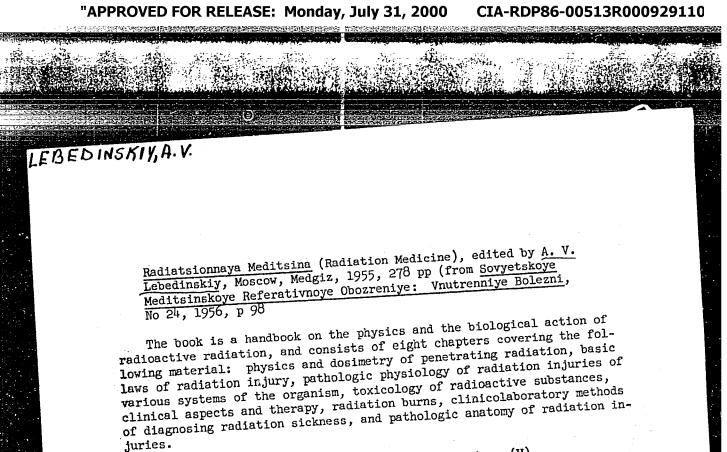
- USSR (600)
- 7. I. P. Pavlov on the work of V. Yu. Chagovets. Fizien. Zhur. 39, No. 2, 1953.

April 1953, Uncl. 9. Monthly List of Russian Accessions, Library of Congress,

CIA-RDP86-00513R0009291100 **APPROVED FOR RELEASE: Monday, July 31, 2000**



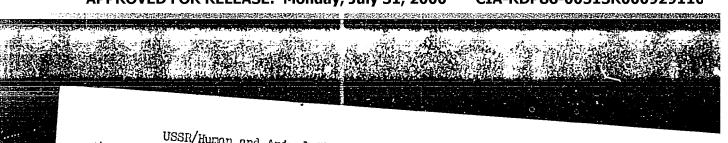
CIA-RDP86-00513R0009291



The book is intended for physicians and students. (U)

APPROVED FOR RELEASE: Monday, July 31, 2000

SUM. 1345



USSR/Human and Animal Physiology (Normal and Pathological). Effects of Physical Factors. Ionizing Emissions.

Abs Jour: Rer Zhur-Biol., No 17, 1958, 80118.

Author : Lebedinskiy, A.V.

Title

: On the Influence of Ionizing Radiation on the Organs of an Animal (According to Data of the Work of Soviet Investigators).

Orig Pub: V.sb.: Deystviye oblucheniya na organism. M., Izd-vo

Abstract: Review. Problems are discriminated on the influence of ionizing radiation on the functional condition of the CNS, the morphological changes of the nerve cells, skeletal muscles and skin coverings, on the change of reflex and conditioned reflex reaction, the condition

Card :1/2

114

EOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R0009291100

USSR/Human and Animal Physiclogy (Normal and Pathological).

Effects of Physical Factors. Ionizing Emissions.

Ats Jour: Ref Zhur-Biol., No 17, 1958, 80118.

of the interoreceptors, as well as on the nucleic acid metabolism changes in the blood picture and the blood-forming organs and the permeability of the blood-brain barrier. Bib. 76 titles.

Card : 2/2

EXCERPTA MEDICA Sec.12 Vol.10/12 Ophthalmology Dec 56

1808. LEBEDINSKY A. V. and PEYMEV I. A. Dept. of Physiol., Milit. Med. Acad., Leningrad. *Polarization phenomena in the retina of the

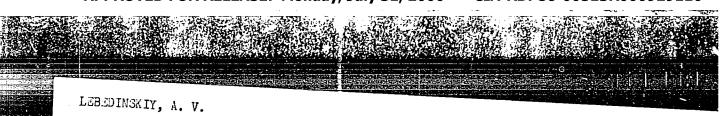
eye (Russian text) PROBL. FIZIOL. OPT. (Moscow) 1955, 11 (107-112) Tables 2 Illus. 3

The experiments were performed on enucleated eyes of the spring frog. The eye was put on a small gauze pad, moistened with Ringer's solution and placed in a small dark chamber. One chlorinated silver electrode was connected to the margin of the cornea, the other to the back-pole of the eye near the outlet of the optic nerve. The currents in the retina were registered by the loop oscillograph. The indifferent polarizing electrode was connected to the gauze pad, the active polarizing electrode was placed on the surface of the eye bulb between the electrodes connected to the oscillograph. The strength of the polarizing current was 20-100 μα. The electroretinogram taken during the illumination of the eye had been registered before the switching on of the polarizing current, during its action and after the switching off. The following results were obtained: during the passing of the direct current and polarization by the anode a distinct lowering of all the 3 peaks, 'A', 'B' and 'C', was observed; the lowering was, to a certain extent, proportional to the strength of the polarizing current. By polarization with the cathode a marked heightening of all the peaks was noticed, particularly of the peak 'B'. After the switching off of the polarizing current, while polarizing with the anode, a distinct increase of the waves was noticed, while, when using the cathode as polarizing electrode, a distinct decrease was noticed after switching off. The changes in the ERG, induced by the polarizing current, vanished only 3 min. after the switching off. The authors regard the results obtained as proof of the fact that the waves 'B' and 'C' of the ERG are the electrical expression of the activity of the ganglionic cells of the retina. The changes observed in the ERG during the polarization of the retina are considered to be changes in the functional state of the nerve cells.

Tron - Leningrad (XII, 2)

APPROVED FOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R000929110



"Effects of Ionizing Radiation of the Nervous System of Animals," Meditsinskiy Ragotnik, Vol. 18, No. 66, 1955.

Comment: This paper was presented by the author at the Geneva Conference on the Peaceful uses of Atomic Energy.

Translation W-31506, Oct 1955.

CIA-RDP86-00513R0009291100

"APPROVED FOR RELEASE: Monday, July 31, 2000

CIA-RDP86-00513R000929110

LEBEDINSKIY, A.V.

USSR/General Division - Congresses. Sessions. Conferences.

A-4

Abs Jour : Ref Zhur - Biologiya, No 1, 1957, 88 K.

: A.V. Lebedinskiy, Editor. Author

Inst Title

: Theses of the Plenary Reports Read at the All-Union

Conference on Medical Radiology, Moscow, 1956.

Orig Pub : M., Medgis, 1956, 31 pp.

: No abstract. Abst

Card 1/1

LEBEDINSHIY, P.V.

KROTKOV, F.G., redaktor; LEHEDINSKIY, A.V., redaktor; IGHAT YEV, A.I., redaktor; LAHDAU, S.P., redaktor; KHANOVA, T.M., redaktor; BELLCHIKOVA, Yu.S., tekhnicheskiy redaktor

[Abstracts of reports at the conference on late sequelae of affections caused by ionizing radiation] Referaty dokladov na konferentsii po otdalennym posledstviiam porazhenii, vyzvannykh vozdeistviem ioniziruiushchei radiatsii. Pod red. vyzvannykh vozdeistviem ioniziruiushchei radiatsii. Pod red. F. G. Krotkova, A.V. Lebedinskogo, A.I. Ignat'eva. Moskva, (MIRA 10:4) Gos. izd-vo med. lit-ry, 1956. 82 p.

1. Russia (1923- U.S.S.R.) Komitet meditsinskoy radiologii. (RADIATION--PHYSIOLOGICAL EFFECT)