

33-3-12/32

Results of photographic observations of Mars at the Kharkov  
Astronomical Observatory during 1954. (Cont.)

atmosphere is more transparent above the "seas" than above the  
"continents".

The polar cap is reddish compared with a white screen. This  
confirms previous results obtained at the Kh.A.O. The fact  
that the smoothness factor for the southern polar cap has a maxi-  
mum value in the blue region of the spectrum agrees well with  
N.P. Barabashov's hypothesis that the polar cap of Mars consists  
of a thin layer of hoar-frost or ice, above which float cloud-  
like formations.

There are 9 tables, 1 figure and 9 references, 8 of which are  
Slavic.

ASSOCIATION: Kharkov Astronomical Observatory (Kharkovskaya  
Astronomicheskaya Observatoriya)

SUBMITTED: November 9, 1956.

AVAILABLE: Library of Congress  
Card 3/3

KOVAL', I.K.

Integral brightness of the solar corona of June 30, 1954. Uch.-  
zap. KHGU 91:24-248 '57. (MIRA 15:3)  
(Sun--Corona)

KOVALEV, I.K.

✓ 3(1)

b.3

PHASE I BOOK EXPLOITATION

SOV/1391

Akademika nauk SSSR. Astronomicheskiy sovet.

Polnyye solnechnyye zatmeniya 25 fevralya 1952 i 30 iyunya 1954 g.  
Trudy ekspeditsiy po nablyudeniyu zatmeniy (Total Eclipse of the  
Sun, February 25, 1952 and June 30, 1954. Transactions of the  
Expedition to Observe Solar Eclipses) Moscow, Izd-vo AN SSSR, 1953.  
357 p. 1,200 copies printed.

Editorial Board: Pariyskiy, N.N., Candidate of Physical and Mathematical Sciences (Resp. Ed.); Kononovich, E.V. (Secretary); Kuz'min, A.D., Candidate of Technical Sciences; Mogilevskiy, E.I., Candidate of Physical and Mathematical Sciences (Deputy Resp. Ed.); Mustel', E.R., Corresponding Member, USSR Academy of Sciences; Ed. of Publishing House: Yegorova, N.B.; Tech. Ed.: Kashina, P.S.

PURPOSE: This book is intended for amateur and professional astronomers interested in eclipse phenomena.

COVERAGE: The present compendium is the fourth in a series published by the Academy of Sciences of the USSR on solar eclipses observed in the Soviet Union. The present collection reports on the results  
Card 1/8

**Total Eclipse (Cont.)**

SOV/1391

of observations obtained by scientific teams of 20 research institutions during the total solar eclipses of 1952 and 1954. The reports include studies of the sun's chromosphere, its total coronal brightness, monochromatic glow, structure, polarization photometry, and colorimetry. The results of studies on coronal radio emissions for various wavelengths and on the effect of the sun on the earth's atmosphere, based on the February 1952 and June 1954 eclipses, are presented. The individual articles are accompanied by tables, diagrams and bibliographic references.

**TABLE OF CONTENTS:**

Foreword	3
Krat, V.A. Gradients of Chromospheric Lines	5
Vyazanitsyn, V.P. Spectrophotometry of the Chromosphere, From Observations of the Total Solar Eclipse of 1952	7
Card 2/8	

## Total Eclipse (Cont.)

SOV/1391

Steshenko, N.V. Distribution of Chemical Elements and Electron Concentration in the Chromosphere (From Observations of the Total Solar Eclipse of February 25, 1952)

15

Steshenko, N.V. and Zemanek, Ye.Kh. Study of the Boundaries of the Chromospheric Lines of Hydrogen, Helium and Ionized Calcium

36

Koval', I.K. Total Coronal Brightness From Observations of Total Solar Eclipses of February 25, 1952 and June 30, 1954

49

Sharonov, V.V. Total Visual Photometry of the Solar Corona in 1952 and 1954

62

Sytinskaya, N.N. Photographic Evaluation of the Total Brightness and Color of the Solar Corona of 1954 in Yeysk

81

Kunsishvili, Ya.I. Radiometry of the Solar Corona During the June 30, 1954 Total Solar Eclipse

83

Card 3/8

Total Eclipse (Cont.)	SOV/1391
Senchuk, Yu.F. Generalized Photometry of the Solar Corona on February 25, 1952	159
Kapko, Ya.T. Photographic Photometry of the Solar Corona on February 25, 1952	173
Gindilis, L.M. Photometry of the Solar Corona on February 25, 1952	187
Aliyeva, G.K. Photometry of the Solar Corona on February 25, 1952	187
Sytinskaya, N.N. Distribution of Brightness and Color in the Solar Corona of June 30, 1954	189
Sharonov, V.V. Visual Colorimetry of the Solar Corona	199
Grigor'yev, P.V. and O.B. Vasil'yev. Photometric Observations of the Solar Corona With Automatic Aerial Cameras During the Total Solar Eclipse of June 30, 1954	207

Card 5/8

Total Eclipse (Cont.)	SOV/1391
Nesmyanovich, A.T. Photometry of the Corona of June 30, 1954	223
Konopleva, V.P. Multi-colored Photometry of the Solar Corona of June 30, 1954	233
Nadubovich, Yu.A. Photometry of the Solar Corona in Red Rays on June 30, 1954	247
Polupan, P.N. Photometry of the Solar Corona in the Green Line λ 5303Å	252
Pariyskiy, N.N and K.I. Petrova. Spectrophotometry of Coronal and Chromospheric Lines During the Eclipse of February 25, 1952	258
Vashakidze, M.A. Analysis of Radiation Polarization of the Solar Corona Based on Observations of Total Solar Eclipse of February 25, 1952	291

Card 6/8

Total Eclipse (Cont.)	SOV/1391
Fomenko, B.D. Variations in the Coefficient of Atmospheric Transparency During the Total Solar Eclipse of June 30, 1954	307
Gavrilov, I.V. and I.G. Kolchiniskiy. Computing Corrections of the Moon's Coordinates From Observations of the Eclipse of June 20, 1954 at the Main Astronomic Observatory of AS UkrSSR.	324
Vitkevich, V.V. and B.M. Chikhachev. Observation of Solar Radio Emissions in the Meter Wave Band During the Total Solar Eclipse of February 25, 1952	329
Troitskiy, V.S., M.P. Zelinskaya, V.L. Rakhlin, V.T. Bobrik. Results of Observation of Solar Radio Emissions in the 3.2 and 10 cm Wavelength During the Total Solar Eclipse of February 25, 1952 and June 30, 1954	330
Molchanov, A.P., E.M. Gyunninen, A.V. Mel'nikov, Al.P. Molchanov, L.L. Myasnikov, V.N. Rysakov, F.I. Skripov, M.M. Filippov.. Results of Solar Eclipse Observations of 1952 and 1954 in the 3.2 cm Wavelength	331
Card 7/8	

KOVAL', I. K. and BARASHOV, N. P.

"Preliminary Results of the Investigation of the Polarization of the Moon  
by Means of Light Filters."

Report presented at the Plenary Meeting of the Committee of Planetary Physics,  
Council of Astronomers, Khar'kov, 20-22 May 1958.  
(Vest. Akad. SSSR, 1958, No. 8, p. 113-114.)

3(1)

AUTHORS: Barabashov, N.P., and Koval', I.K. SOV/33-35-2-11/21  
TITLE: The Structure of the Southern Polar Cap of Mars in 1956 (K voprosu o stroyenii yuzhnay polyarnoy shapki Marsa v 1956 g)  
PERIODICAL: Astronomicheskiy zhurnal, 1958, Vol 35, Nr 2, pp 261-264 (USSR)  
ABSTRACT: Using the data of observations of Mars in 1956, made in visual light at various observatories of the world and also observations in the ultraviolet made at the Khar'kov Astronomical Observatory, the author considers one of the possible interpretations of the structure of caps of Mars. The simultaneous disappearance of the southern polar cap around September 1, 1956 in red and ultraviolet light (fig. 3,4) can be explained by the fact that at that time the caps lay wholly on the solid surface of the planet in accordance to Wright, [Ref 6]. The contrast in the brightness between the cap and the rest of the planet's surface observed in various parts of the spectrum is explained by the absorbing properties of atmosphere of Mars. An

Card 1/2

The Structure of the Southern Polar Cap of Mars  
in 1956

SOV/33-35-2-11/21

approximate estimation of the optical density of the planet's atmosphere in the ultraviolet ( $\tau \approx 0.3$ ) has been made. The author mentions especially observations of G.I. Tikhov [Ref 5] and Ye.L.Krinov [Ref 7]. There are 6 figures, 1 table, and 9 references, 6 of which are Soviet, and 3 American.

ASSOCIATION: Khar'kovskaya astronomicheskaya observatoriya (Khar'kov  
Astronomical Observatory)

SUBMITTED: August 21, 1957

Card 2/2

3(1)

AUTHOR: Koval', I.K.

SOV/33-35-5-14/20

TITLE: The Absolute Photometry of Venus in the Ultraviolet and Infrared  
(Absolyutnaya fotometriya Venery v ul'trafioletovykh i infra-  
krasnykh luchakh)

PERIODICAL: Astronomicheskiy zhurnal, 1958, Vol 35, Nr 5, pp 792-796 (USSR)

ABSTRACT: From August 23 to 26, 1956, the author obtained ca. 60 photos  
of Venus in the neighborhood of the quadrature. The 270 mm  
reflector of the Khar'kov Astronomical Observatory and ultra-  
violet as well as infrared light filters were used. The author  
describes his methods in detail and collects the results in three  
tables. For the evaluation a method of N.N.Sytenskaya [Ref 1]  
at the Observatory of the Leningrad State University was used.  
The author compares the results with those of I.A.Parshin [Ref 2],  
N.P.Barabashov [Ref 3] and with the model of V.V.Sharonov  
[Ref 5]. He thanks the students of the Khar'kov University  
M.Levertov and O.Bugayenko for aid during the observation.  
There are 5 tables, 1 figure, and 5 Soviet references.

ASSOCIATION: Khar'kovskaya astronomicheskaya observatoriya (Khar'kov  
Astronomical Observatory)

SUBMITTED: August 21, 1957  
Card 1/1

Koval, L.R.

PHASE I BOOK EXPLOITATION

SOV/4302

Akademiya nauk SSSR. Komissiya po fizike planet

Izvestiya, vyp. 1 (News of the Commission on the Physics of Planets, No. 1)  
Khar'kov, 1959. 108 p. 1,000 copies printed.

Editorial Board: N.P. Barabashov, Academician of the Academy of Sciences  
Ukrainskaya SSR (Resp. Ed.); V.I. Yezer'skiy, Candidate of Physics and  
Mathematics (Secretary); A.V. Markov, Professor; Yu. N. Lipskiy, Candidate of  
Physics and Mathematics; and A.T. Chekirda, Candidate of Physics and Mathematics;  
Ed.: D.A. Vaynsberg; Tech. Ed.: A.S. Trofimenco.

PURPOSE: This publication is intended for astrophysicists and astronomers.

COVERAGE: This collection of articles constitutes the first issue of a new journal  
on problems in planetary physics. The first six articles discuss the surface  
features, polarimetry, and spectrophotometry of the Moon. The remaining articles  
deal with the physics of Mars, Jupiter, and the asteroids. No personalities are  
mentioned. References accompany individual articles.

Card 1/3

## News of the Commission (Cont.)

SOV/4502

## TABLE OF CONTENTS:

Barabashov, N.P., and A.T. Chekirda. Types of Rocks Most Closely Corresponding to Those of the Lunar Surface	5
Kolchan, Ye. K. Investigation in Three Parts of the Spectrum of the Degree and Angle of the Position of the Plane of Polarization of Light Reflected From Lunar Features	41
Barabashov, N.P., and I.K. Koval'. Photographic Polarimetry of the Moon With Light Filters	55
Sergeyeva, A.N. Spectrophotometry of Lunar Formations	59
Barabashov, N.P., V.A. Yezerskaya, and V.I. Yezerskiy. The Problem of the Photometric Uniformity of the Moon's Surface	67
Bytinskaya, N.N. Probable Dimensions of the Ruggedness of the Microrelief of the Moon's Surface	81

Card 2/3

Koval, I.K.  
P2'

MASSIVE BOOK EXPLOITATION

SOV/3923

Akademiya nauk SSSR. Astronomicheskiy sovet. Komissiya po fizike planet

Rezul'taty nablyudenij Marса vo vremya velikogo protivostoyaniya 1956 g. v SSSR.  
(Results of the Observation of Mars in the USSR During the Great Opposition  
of 1956) Moscow, Izd-vo AN SSSR, 1959. 193 p. 1,300 copies printed.

Ed.: N.P. Barabashov, Academician, Academy of Sciences UkrSSR; Ed. of Publishing  
House: V.A. Bronshten; Tech. Ed.: V.V. Brugul'.

PURPOSE: The book is intended for astronomers and astrophysicists, particularly  
those interested in the study of the planet Mars.

COVERAGE: This is a collection of 8 articles on the results of observations of  
Mars during the opposition of 1956. The observations were organized by the  
Committee for the Investigation of Physical Conditions on the Moon and Planets  
of the Astronomic Council of the Academy of Sciences USSR, and were conducted  
mainly at the southern observatories of the Soviet Union during the period from  
August to December, 1956. The preliminary results of those observations are

Card 1/3

## Results of the Observation of Mars (Cont.)

SOW/3923

summarized at a conference of all participants held in Moscow, December 1956. The present publication covers the investigations of Mars' surface and atmosphere by photometry as well as by visual observations, problems of the flora of the planet and the photometric investigation of its luminous Argyre zone. References accompany most of the articles.

## TABLE OF CONTENTS:

## Observations of Mars in the USSR in 1956

Tikhov, G.A. Brief Results of the Observations of Mars by the  
Astrobotany Section During the Period of the Great Opposition of 1956

Barabashov, N.P., and I.K. Noval'. Photographic Photometry of  
Mars With Light Filters

Sytinskaya, N.N. Photometric Investigation of Optical Properties of  
the Atmosphere of the Planet Mars

Sharonov, V.V. Surface and Atmosphere of Mars According to Photographic,  
Photometric, and Colorimetric Observations Conducted in Tashkent in 1956

Card 2/3

Koval', I.K.

PHASE I BOOK EXPLOITATION

SOV/4093

Barabashov, Nikolay Pavlovich, and Ivan Kirillovich Koval'

Fotograficheskaya fotometriya Marsa so svetofil'trami vo vremya velikogo protivostoyaniya v 1956 g. (Photographic Photometry of Mars Using Light Filters During the Great Opposition of 1956) Khar'kov, Izd-vo Khar'kovskogo univ-ta, 1959. 529 p. 1,000 copies printed.

Resp. Ed.: A.T. Chakirda; Eds.: A.N. Tret'yakova, and D.A. Vaynberg;  
Tech. Ed.: A.S. Trofimenko.

PURPOSE: This book is intended for scientists interested in physical conditions on planets, and may be used by students and aspirants concerned with problems of planetary physics.

COVERAGE: The book discusses the methodology and results of observations of Mars during the great opposition of that planet in 1956, and contains tabulated data compiled at the Kharkov Astronomical Observatory on the absolute photometry of Mars. The book also contains some results of visual observations of the Martian surface. The major part of the text is a tabulation (Table 4) of spectral values of the brightness factor  $\rho$  for different points on the continent, bright areas,  
Card 14

Photographic Photometry of Mars (Cont.)

sov/4093

and polar cap of Mars: The main series of observations were not made simultaneously, but lasted from 30 to 40 minutes each, and hence, for each individual light filter, the length of the central meridian of the planet, and the values  $\zeta$  (incident angle of light ray),  $\epsilon$  (angle of reflection), and  $\varphi$  and  $\lambda$  (width and length of the planetocentric coordinates of the point source of light, respectively) are given for the period from June 15/16 to October 21/22. The value of the phase angle  $\alpha$  is also shown for each date and each filter. The supplement contains graphs which plot the distribution of brightness along the equator (in a band -  $10^\circ$ ,  $-30^\circ$  in width) and over the central meridian of the planet. The abscissas give values of  $\epsilon$ , and the ordinates, values of  $\rho \cdot 10^2$  (for simplicity). There are 9 figures, 14 tables, and 15 references: 13 Soviet and 2 English. No personalities are mentioned.

## TABLE OF CONTENTS:

- |   |   |
|---|---|
| 1. Purpose and methodology of observations of Mars during<br>the great opposition of 1956 | 1 |
| 2. Processing observational data  | 1 |
| 3. Visibility of detail on the surface of Mars in 1956                                    | 1 |
| 4. Clouds, fogs, and precipitations observed on Mars in 1956                              | 1 |
| 5. Some details observed on the Martian disc  | 2 |
| 6. Description of the appearance and changes in the southern polar cap                    | 2 |

Card 2/4

BABABASHOV, N.P., prof.akad.; KOVAL', I.K.

Photographic polarimetry of the moon with light filters. Izv.  
Kom.po fiz.plan. no.1:55-58 '59. (MIRA 13:7)

1. AN USSR (for Barabashov).  
(Polariscope) (Moon—Observations)

KOVAL', I.K.

Degree of the evenness of continents and seas on Mars. Izv.  
Kom. po fiz. plan. no. 1:85-92 '59. (MIRA 13:7)  
(Mars(Planet))

BARABASHOV, N.P. [Barabashov, M.P.], akademik; KOVAL', I.K.

Distribution of brightness in Mars' seas. Dop.AN UESR no.2:  
153-155 '59.  
(MIRA 12:5)

1. AN USSR (for Barabash). 2. Khar'kovskaya astronomicheskaya  
observatoriya.  
(Mars (Planet))

BARABASHOV, M.P., akademik; KOVAL', I.K., kand.fiz.-mat.nauk

Main results of observations of Mars during 1956 at the Kharkov  
Astronomical Observatory. Visnyk AN URSR 30 no.1:15-23 Ja '59.

(MIRA 12:4)

1. AN USSR (for Barabashov)  
(Mars (Planet))

30273

3,1550 (1041,1057)

S/035/61/000/010/027/34  
A001/A101

AUTHORS: Barabashov, N.P., Koval', I.K., Chekirda, I.T.

TITLE: Some results of photometry of cloudy formations on Mars

PERIODICAL: Referativnyy zhurnal. Astronomiya i Geodeziya, no. 10, 1961, 66, abstract 10A#57 ("Izv. Komis. po fiz. planet", 1960, no. 2, 36-40)

TEXT: Blue clouds observed mainly in the equatorial zone of the planet are described on the basis of photographs taken in 1958 by means of a 10" reflector of the Khar'kov Observatory with the equivalent focus equal to 15.3 m in ultraviolet ( $\lambda$  3600) and blue ( $\lambda$  4200) portions of spectrum. The existence of a photometric scale makes it possible to obtain the curves of brightness distribution along the equator of intensity, when the clouds are present and at their absence. The analysis of materials has shown that the presence of clouds on the disk and terminator does not depend on the albedo of the underlying solid surface of Mars. The value of cloud-background contrasts, and the albedo of the clouds on the terminator are higher than on the disk. On an average, the bright-

Card 1/2

Some results of photometry . . .

30273  
S/035/61/000/010/027/034  
A001/A101

ness of clouds exceeds the background brightness by 14%. On assumption that the clouds studied are analogous, in their nature, to cirrus clouds of the Earth's atmosphere, their thickness was estimated to be 3 - 6 m.

I. Lebedeva

[Abstracter's note: Complete translation]

Card 2/2

S/033/60/037/02/009/013  
E032/E914

AUTHORS: Barabashov, N. P., Koval', I. K.

TITLE: Some Results of Studies of Contrasts on Mars ✓

PERIODICAL: Astronomicheskiy zhurnal, 1960, Vol 37, Nr 2, pp 301-303  
(USSR)

AESTRACT: It is well-known that the continent-sea contrast  $K$  for Mars has a maximum value in red and a minimum in blue. According to measurements carried out by the present authors in 1956,  $K = 0.286$  for  $\lambda_{\max} = 647 \text{ m}\mu$ , and  $K = 0.07$  for  $\lambda_{\max} = 420 \text{ m}\mu$ . However, occasionally, the contrast in blue light reaches 80% of the contrast in red light. Usually, this is explained by increased transparency of the blue atmosphere on Mars. If one accepts the explanation that the true contrast is the same for all wavelengths and the apparent contrast depends on the state of the Martian atmosphere, then one can try and estimate the optical thickness of the Martian atmosphere in blue light from the corresponding contrast. For this purpose it is assumed

Card 1/3

✓

S/033/60/037/02/009/013  
E032/E914

Some Results of Studies of Contrasts on Mars

that  $K_{\lambda} = \text{const}$  and the red contrast is taken to be approximately equal to the true contrast, since the optical thickness of the atmosphere in red light may be taken to be approximately zero. The present paper is concerned with a critical examination of the possibility of an explanation of the variation in  $K_{\text{blue}}$  by variations in the optical thickness. Fig 1 gives the values of the contrast for red, green and blue rays (curves 1, 2 and 3 respectively) for June, July, August and September, 1956. These data refer to the central regions of the planet. Fig 2 gives the corresponding plot for the ratio  $K_{\text{blue}}/K_{\text{red}}$  for the same observational period. Another characteristic considered is  $B = \rho_1(1 - K_{\text{red}}/K_{\text{blue}})$  where  $\rho_1$  is the albedo of the

Card 2/3

VC

S/751/61/000/008/001/005

AUTHORS: Barabashov N. P., Koval', I. K., Chekirda, A. T.

TITLE: Photometric Investigations of Mars in 1958.

SOURCE: Akademicheskaya nauk SSSR. Komissiya po fizike planet. Izvestiya. no. 3. Kharkov, 1961. 3 - 5.

TEXT: Data are presented on the distribution of brightness along the intensity equator and central meridian of Mars, obtained on the basis of photographic observations of the planet using optical filters. The emulsions and filters employed are listed, and the exposures stated. No tie-in was made with the sun, since the experiment was not aimed at obtaining a detailed list of absolute brightnesses. All negatives were calibrated with a tubular photometer and measured with a MF-2 microphotometer. Data on the contrast between the bright and dark regions, and on the variation of the brightness of the northern and southern polar regions of Mars, are tabulated. Except for one longitude region, satisfactory correlation is observed between the "red" and "blue" curves, and the lack of correlation in the particular region is attributed to the specific distribution of the energy in the spectrum of the corresponding seas

(Card 1/2

Photometric Investigations of Mars....

S/751/61/000/008/001/005

of Mars, with the blue atmospheric haze playing a secondary role. There are 5 tables.

ASSOCIATION: Astronomicheskaya observatoriya Khar'kovskogo universiteta  
(Astronomical Observatory of the Kharkov University)

Card 2/2

S/751/61/000/008/005/005

AUTHCR: Koval', I. K.

TITLE: Results of an investigation of contrasts on Mars

SOURCE: Akademiya nauk SSSR. Kaniissiya po fizike planet. Izvestiya, no.3. Kharkov, 1961. 76 - 85

TEXT: A more detailed interpretation is attempted of results previously obtained by the author (in conjunction with N. P. Parabashov (Astron. zhur. v.37, no.2, 1960) and based on photographic observations of Mars made in June-August 1956, that the contrasts between continents and oceans on Mars exhibit fluctuations from day to day when viewed in all parts of the spectrum except the ultraviolet. These fluctuations cannot be ascribed to variations in the transparency of the Mars atmosphere, for it would lead to an unreasonably large variation in the transparency coefficient in blue light. Approximate calculations are presented of the true values of the contrasts in green and blue light, and it is concluded that the contrasts do not remain constant over the spectrum, their variations being due to differences in the optical properties of the oceans of

Card 1/2

Results of an investigation...

S/751/61/000/008/005/005

Mars themselves. Calculations are presented in favor of this assumption. The incomplete nature of the data is pointed out. There are 7 figures and 6 tables.

ASSOCIATION: Glavnaya astronomicheskaya observatoriya AN SSSR (Main Astronomical Observatory, Academy of Sciences SSSR)

Card 1/2

3.1550 (1041,1057)

33423

S/033/62/039/001/007/013  
E032/E514

AUTHORS: Koval', I.K. and Morozhenko, A.V.

TITLE: On some properties of the yellow haze observed on Mars in 1956

PERIODICAL: Astronomicheskiy zhurnal, v.39, no.1, 1962, 65-72

TEXT: It is stated that all the observers of Mars noted a relatively sudden decrease in contrast between the bright and dark regions on the surface of the planet in the middle of September, 1956. Most observers, including the present authors, are said to ascribe this phenomenon to the appearance of a fine dust cloud in the Martian atmosphere. The aim of the present paper was to determine the optical thickness of this cloud in red and infrared light. The analysis is based on the experimental results reported by N. P. Barabashov and I. K. Koval' (Ref. 5: "Photographic photometry of Mars using light filters", Izd-vo KhGU, Khar'kov, 1960). A detailed analysis of these data, including corrections for the seasonal variations in the surface contrasts, leads the present authors to the conclusion that the dust cloud had mainly scattering properties and that its optical

Card 1/3

33423

On some properties of the yellow ... S/033/62/039/001/007/013  
E032/E514

thickness increased with increasing wavelength. It is estimated that the optical thickness was as follows:

Table 5

$\lambda, \text{m}\mu$	840	750	647	530	430	360
$\tau_\lambda$	0.34	0.26	0.19	(0.12)	(0.08)	(0.01)

Next, using the results reported by V. V. Shuleykin (Ref.11: Physics of the sea, Izd-vo AN SSSR, M., 1953, p.589), the authors attempt to deduce the radius of the particles of which the cloud was made up. It is estimated that this radius was  $1.45 \mu$ . Finally, a calculation is made of the sedimentation rate of such particles. It is assumed in this calculation that the particles do not collide with each other and that convective motion in the atmosphere may be neglected. Taking the mean free path of gas molecules in the Martian atmosphere to be  $10^{-7} \text{ cm}$ , and assuming

Card 2/3

KOVIL, I.K.

Distribution of Brightness in the Edge Zone of Mars

Report to be submitted for the 4th International Space Science Symposium  
(COSPAR) Warsaw, 2-12 June 63

KOVAL', I.K.; MOROZHENKO, A.V.

Degree of smoothness of the Martian deserts. Izv. Kom. po fiz. plan.  
no.4:38-39 Ag '63. (MIRA 18:5)

1. Glavnaya astronomicheskaya observatoriya AN UkrSSR.

ACCESSION NR: AT3007874

S/2617/63/005/001/0047/0067

AUTHOR: Didychenko, Ye. I.; Koval', I. K.; Morzhenko, A. V.

TITLE: Results of spectrophotometric observations of Mars in 1960-1961

SOURCE: AN UkrRSR. Holovna astronomichna observatoriya. Izvestiya, v. 5,  
no. 1, 1963, 47-67

TOPIC TAGS: spectrophotometric observation, intensity distribution, disc radius  
brightness, polar cap

ABSTRACT: In view of the almost total lack of spectrophotometric observations of Mars, the three authors undertook to obtain as many photometrically standardized spectrograms of it as possible with different orientations of the spectrograph aperture, to study: 1) the distribution of intensity in the spectrum of the continents and seas at different distances from the center, the polar caps and the Right clouds; 2) the distribution of brightness along the equator of intensity, the equator of the planet and the central meridian in monochromatic rays, including a study of the spectral course of the continent-sea contrast, establishment of the boundary of its disappearance, and also possible shifts of this boundary

Card 1/3

ACCESSION NR: AT3007871

depending upon the state of the atmosphere, characterized by possible changes in the distribution of energy in the spectrum of the continents. Observations were made in the Cassegrain focus of the 70 cm reflector telescope of the State Astronomical Observatory of the Ukrainian SSR Academy of Sciences from 1 October, 1960 to 6 February, 1961, with spectrograph ASP-5, aperture 0.05 mm, 70 spectrograms being obtained on "Kodak OAF", but only 23 selected as best. The authors conclude: 1) The reflecting power of Mars (center, edges, polar caps, Right clouds) according to the spectrum varies smoothly from 0.330 ( $\lambda = 6550 \text{ \AA}$ ) to 0.053 ( $\lambda = 4030 \text{ \AA}$ ), so that no dispersion waves were discovered during the 1960-1961 opposition; 2) the distribution of brightness along the radius of the disc (along the equator of intensity) at  $\lambda = 6550 \text{ \AA}$  near opposition is well represented by  $\rho \propto \cos i$ ; but with growth of the phase angle the curve rises above the sinusoid branch; 3) the curvature of the brightness curve along the equator of intensity diminishes smoothly from the red to the violet end of the spectrum; 4) the brightness curves along the equator of intensity and the central meridian coincide in the red rays, leading to the conclusion that both polar caps were atmospheric formations in the period of observations; 5) the spectral course of the continent-sea contrast could not be

Card 2/3

ACQUISITION NR: AT3007874

studied because areas with low contrast were located near the center of the disc when the atmospheric conditions were favorable (on the dates of the observations). Photoelectric observations led to a mean contrast value of 0.33 with lambda = 620 A. Original has 11 graphs and 6 tables (forming 2/3 of the article).

ASSOCIATION: none

SUBMITTED: 00

DATE ACQ: 28Jun63

ENCL: 00

SUB CODE: AS

NO REF SOV: C04

OTHER: 000

Card 3/3

VOLOTKOVSKIY, S.A., doktor tekhn.nauk; FURSOV, V.D., inzh.; KOVAL', I.K.  
inzh.; RUD', V.I., inzh.

Operating characteristics of electric charging devices with semi-conductor rectifiers for use in mines. Vest. elektroprom. 34 no.8:  
62-64 Ag '63. (MIRA 16:9)  
(Electric current rectifiers) (Electricity in mining)

KOVAL', I.K., otv. red.; FEDOROV, Ye.P., red.; GORYNYA, A.A., red.;  
MOLCHINSKIY, I.G., red.; LUKATSKAYA, F.I., red.;  
BEREZINETS, L.P., red.

[Physics of the moon and planets] Fizika Luny i planet.  
Kiev, Naukova dumka, 1964. 137 p. (MIRA 17:10)

I. Akademiya nauk URSR, Kiev. Holovna astronomichna ob-  
servatoriya.

J. 248 3-65 EWC(7) EWT(1)/E O(t)  
ACCESSION NR: AT 10 1983

Pe-5/Pap-4 GW/MILK  
1/0000/64/000/000/0046/0053

2  
2  
B7

AUTHOR: Koval', I.K.

TITLE: A study of the optical properties of the Martian surface and atmosphere

SOURCE: AN UkrSSR Glavnaya astronomicheskaya observatoriya. Fizika Luny i planet (Physics of the moon and planets) Kiev, Naukova dumka, 1984, 46-53

TOPIC TAGS: Mars, brightness distribution, Martian surface, Mars albedo, Martian atmosphere

ABSTRACT: The paper develops probable fundamental characteristics of the Martian surface and atmosphere, based upon a review of a series of previously published investigations. After a review of the inconsistencies in the determination of the brightness ratio  $m = \frac{I_0}{I_{90}}$  where  $I_0$  and  $I_{90}$  are brightnesses at light impingement angles of  $60^\circ$  and  $0^\circ$ , respectively, the author offers some possible explanations and concludes that various technical errors and intrinsic drawbacks of the photometric methods used may account for most of the inconsistencies. Turning to the behavior of "m" as a function of wavelength, he notes the dependence of "m" upon the dust content of the Martian atmosphere. He then discusses two variants of atmospheric characteristics on the basis of a list of observati

cord 1/2

L 4303-65  
ACCESSION NR: AT40 9983

drawn from several studies. The most important phenomenon relates to the dependence of contrast between continental and "maritime" portions of the Martian surface upon dust content. Summarizing, the author concludes that the Martian atmosphere has, basically, scattering characteristics. Furthermore, the color of Mars is fundamentally explainable as that of its own surface. The Dollase's supposition of a limonite powder surface seems quite plausible. The true brightness contrasts of Martian continents and seas become completely washed-out by the atmospheric scattering at 4000 Å. In the bandwidth 4000 - 6500 Å, the Martian atmosphere exhibits a Raleigh behavior which slows down the rate of brightness decrease toward the visible margin at short wavelengths. The Martian atmosphere contains quantities of floating coarse particles, which behave in an anti-Raleigh manner with regard to light scattering. Thus, both to the right and to the left of 6500 Å, in the investigated bandwidth of 4000-6400 Å, the influence of the Martian atmosphere increases toward the edges. Observations aimed at the study of the surface alone should therefore use wavelengths in the 6500 Å region. Orig. has: 3 tables and 1 formula.

ASSOCIATION: Glavnaya astronomicheskaya observatoriya AN UkrSSR, Kiev (Main Astronomical Observatory, AN Ukr. SSR)

SUBMITTED: 07 May 84

ENCL: 00

SUB CODE: AA

NO (E) SOV: 011

OTHER: 001

Card 2/2

L 25 MIG-65 EMT(1) 2395(v)/EMI(t) Pe-5/Pac-2 GW/MLK

ACCESSION NR: A74047884

S/0000/64/000/000/0054/0057

AUTHOR: Glugayenko, I. A.; Bokarenko, O. I.; Koval, I. K.; Morozhenko, A. V.

TITLE: Brightness distribution in the marginal zone of Mars

SOURCE: AN UkrSSR. Glavnaya astronomicheskaya observatoriya. Fizika Lunny i planeta (Physics of the moon and planets). Kiev, Naukova dumka, 1984, 54-57

TOPIC TAGS: light scattering, Mars opposition, brightness distribution, Martian atmosphere, light absorption, turbulence, vibration, photoelectric observation

ABSTRACT: The purpose of this work was to determine the optical characteristics of the Martian atmosphere by a study of brightness as a function of the angle of incident light. To obtain this information, a study of the marginal zone is imperative, but photographic methods are found to be deficient for this purpose. The method used involved a photoelectric sensor, coupled with a very small diaphragm opening subtending only 0".35. The device was placed at the Cai's main focus of a 70-cm reflecting telescope. Photomultipliers were used, with filters covering a spectral range of 1550-5000 Å. During the Mars opposition of Feb. 4, 1983, the conditions were perfect and 40 to 50 diameter transits were made for each light filter, with the zenith distance never exceeding 35°. The effective amplitude of

Card 1/3

I 15046-65

ACQUISITION NR: AT 10 1984

turbulent image vibration was only 0.1%. The authors found the true brightness distribution along the diameter of Mars, which requires correction for washout and image vibration, in the following way. An integral equation was set up by writing:

$$F(x) = \frac{1}{A} \int_{-\infty}^{\infty} f(x-y) \cdot S(y) dy \quad (1)$$

where  $F(x)$  and  $f(x)$  are respectively the observed and true brightness distribution along the diameter of Mars. The kernel,  $S(x-y)$ , is the brightness distribution for a finite, "orinally" vibrating point light source. This can be determined from the expression

$$S(x-y) = A \int_{-\infty}^{\infty} R_d \cdot (x-y - y)^2 \cdot e^{-y^2/2\epsilon^2} dy \quad (2)$$

Here,  $A$  is the normalization constant,  $R_d$  is the diaphragm diameter and  $\epsilon$  is the amplitude of image vibration. The true brightness distribution was determined by first solving expression (2) for the kernel  $S$ , and then solving the integral equation (1) by an iterative

Class 2/3

L 21016-15

ACCESSION NR: A1049984

method which converges rapidly. Results of brightness determination for  $\lambda = 4200 - 6000$  are given in a table. They point to the prevalence of scattering in the visible region. Results for  $\lambda = 3550 - 4200$  Å are discussed quantitatively without giving the details in table form. The conclusion is drawn that the Martian atmosphere has significant true absorption at around 3550 Å. Orig. art. has: 2 tables and 6 formulas.

ASSOCIATION: None

SUBMITTED: 07 May 51

ENCL: 00

SUB CODE: AA

NO REF Sov: 00

CARRIER: 000

CPLS: 3/3

YAKOVKIN, A.A., otyv. red.; FEDOROV, Ye.P., red.; AKSENT'YEVA, Z.N., red.; BARABASHOV, N.P., red.; BOGORODSKIY, A.F., red.; GORYN'YA, A.A., red.; KOVAL', I.K., red.; KOLCHINSKIY, I.G., red.; TSESEVICH, V.P., red.; KOVALENKO, L.D., red.

[Figure and motion of the moon] Figura i dvizhenie Luny.  
Kiev, Naukova dumka, 1965. 135 p. (MIRA 18:7)

1. Akademiya nauk URSR, Kiev.

KOVAL', I.K., otv. red.; KOVALENKO, L.D., red.

[Problems in astrophysics; study of the atmospheres of Venus and Mars. Abstracts of the reports] Voprosy astrofiziki; issledovanie atmosfer Venery i Marsa. [Tezisy dokladov] Kiev, Naukova dumka, 1965. 164 p. (MIRA 18:9)

1. Soveshchaniye rabochey gruppy po issledovaniyu planet zemnogo tipa, Kiev, 1964.

L 16988-66 EMT(1) GW

ACC NR: AP6001511

SOURCE CODE: UR/03&4/65/000/006/0053/005;

AUTHOR: Koval', I. X. (Candidate of physico-mathematical sciences)

28

22

B

ORG: none

TITLE: The study of earth-like planets

SOURCE: Zemlya i Vselennaya, no. 6, 1965, 53-55

TOPIC TAGS: Venus planet, Mars planet, planetary atmosphere

ABSTRACT: This is a survey of investigations of the planets Mars and Venus, particularly in view of setting up stations there. The Commission on Planetary Physics of the Astronomical Council of the AN SSSR organized a group in 1963 to study these planets, to gather all available information, and to plan future investigation. A conference was held in 1964 in Kiev and another in 1965 at the Crimean Astrophysical Observatory. Atmospheric pressure on Mars, first computed by optical thickness, was determined to be 60 mm Hg, but it was concluded that this value must be in error because of interference of aerosol particles. A spectroscopically determined value is 15-20 mm Hg. The photoelectric method gives a maximum of 15 mm. Data from Mariner-4 suggest a value of

Cord 1/2

L 16988-66

ACC NR: AP6001541

about 10 mm Hg. V. I. Moroz considers the pressure on Venus to be 10 atm, but observational values are available only for the upper boundary of the cloud layer (20 mm Hg). Infrared spectroscopy reveals carbon dioxide, oxygen, and water vapor in the atmosphere of Venus, and carbon dioxide and water vapor in the atmosphere of Mars. Transparent particles (ice?) have been discovered in the upper atmosphere of Venus. A number of workers spoke at the last conference (Crimea): M. V. Morozhenko and M. M. Fospergelis on polarimetric investigations, I. A. Mitrofanova on molecular absorption spectra of different gases, I. N. Minin on theoretical work at the Main Astrophysical Observatory on interpretation of polarimetric data, and N. A. Kozyrev on observations of the dark side of Venus. Orig. art. has 2 figures.

SUB CODE: 03/ SUBM DATE: none

Card 2/2 7/25

L 11520-66	EWT(1)	(W)
ACQ N:	AR6001134	SOURCE CODE: UR/0269/65/005/009/0055/0055
SOURCE:	Ref. zh. Astron. tsirkulyar, No. 9.51.470	
AUTHOR:	Koval', I. N.	
TITLE:	Photoelectric measurements of the brightness distribution over the disk of Mars	
REFERENCE SOURCE: Astron. tsirkulyar, no. 319, marta 15, 1965, 1-4		
TOPIC THOS: Mars planet, optic brightness, Gaussian distribution, atmosphere, photoelectric method, telescope, solar radiation		
TRANSLATION: An attempt is made to correct observed curves of the brightness distribution over the disk of Mars for the entire set of distortions caused chiefly by the earth's atmosphere. For these purposes, parallel observations of Mars and stars by the method of photovoltaic image scanning were begun in 1965 at the Main Astronomical Observatory, Academy of Sciences, Ukrainian SSR, with the 70-cm reflector. The stellar profiles obtained from the observations are well approximated by a Gaussian distribution. This result was used for an approximate correction of the photoelectric profile of Mars obtained in the opposition of 1963. For a filter with $\lambda_{eff} 600 \text{ m}\mu$ , the corrected curve of the decrease in brightness toward the limb has a shape that is close to $\sqrt{\cos i}$ ( $i$ is the angle of incidence of the solar rays), instead of the		
Card 1/4		UDC: 523.43

L 11528.66

ACC NR: AR6001134

observed Lambert curve. It is concluded that the surface of Mars differs from the  
Lambertian in that it is rougher. I. K.

SUB (DKR, O)

Card 1/2

PLAKSENKO, N.A.; KOVAL', I.K.

Characteristics of the distribution of impurity elements in  
the rocks of siliceous-ferruginous formation of the Kursk  
Magnetic Anomaly and their genetic significance. Dokl. AN SSSR  
161 no.1:210-213 Mr '65.  
(MIRA 18:3)

I. Voronezhskij gosudarstvennyy universitet. Submitted June 8,  
1964.

ACC NR: AT6033321

SOURCE CODE: UR/0000/66/000/000/0018/0030

AUTHOR: Bugayenko, L. A. -- Bugayenko, O. I. ; Koval', I. K. ; Morozhenko, A. V.

ORG: none

TITLE: Electrophotometric sections of the Mars planet disk in the spectral range of the 355-600 m $\mu$  interval

SOURCE: AN UkrSSR. Fizika i planet (Physics of the Moon and the planets) Kiev, Naukova dumka, 1966, 18-30

TOPIC TAGS: Mars planet, star, Mars, brightness distribution

ABSTRACT: Electrophotometric sections of the images of Mars and of some stars situated at a small angular distance from the planet were obtained with a 70-cm reflector at the Main Astronomical Observatory of USSR in 1956. The information now being published represents the experimental part of an investigation aimed at correcting the brightness distribution curve along the disk of Mars by calculating the influence of factors in the earth atmosphere. The authors thank Z. Merkulova

Card 1/2

ACC NR: AT6033321

APPROVED FOR RELEASE: 06/14/2000 CIA-RDP86-00513R000825510014-5"  
and V. Pipko for their assistance in calculations. Orig. art. has: 5 figures and 3 formulas. [Based on authors' abstract]

SUB CODE: 03/SUBM DATE: 19Mar66/ORIG REF: 017/OTH REF: 003/

Card 2/2

ACC NR: AT6033322 SOURCE CODE: UR/0000/66/000/000/0031/0044

AUTHOR: Koval', I. K.

ORG: none

TITLE: Position of the maximum brightness on the disk of Mars

SOURCE: AN UkrSSR. Fizika Luny i planet (Physics of the Moon and the planets)  
Kiev, Naukova dumka, 1966, 31-44

TOPIC TAGS: optic brightness, planetary atmosphere, Mars planet, Mars  
atmosphere thickness

ABSTRACT: The results of the photometric processing of several photographs  
of the planet Mars, obtained in 1956 for several phases along the equator of  
intensity, are given. It is shown that with a decrease in the wavelength the optical-  
brightness maximum shifts from the subsolar point to the limb. In the case of pure  
scattering, the optical thickness of Mars' atmosphere and the surface albedo for  
blue and green rays ( $429 \text{ m}\mu$  and  $530 \text{ m}\mu$ ) are estimated for spherical indi-  
catries on Lambert's ground. In the case of pure UV diffusion, there is no agree-  
ment between the theoretical curves; this indicates the existence of true absorption

Card 1/2

ACC NR: AT6033322

in the spectral region. Orig. art. has: 16 figures and 2 tables. [Based on author's  
**APPROVED FOR RELEASE: 06/14/2000 CIA-RDP86-00513R000825510014-5"**  
abstract]

SUB CODE: 03/SUBM DATE: 19Mar66/ORIG REF: 003/OTH REF: 002/

Card 2/2

KOV.L', I. P.

Dissertation defended for the degree of Candidate of Agricultural Sciences  
at the Institute of Forest and Wood; Siberian Branch

"Role of the Pine Tree in Increasing the Productivity of Low-Bonitetic Oak  
Forests of the Northwest Caucasus."

Vestnik Akad. Nauk, No. 4, 1963, pp 119-145

KOVAL', I.P.

Silvicultural economic evaluation of some types of oak and  
oak-pine forests in the northwestern Caucasus. Trudy Inst.  
lesa AN Gruz. SSR 10:205-214 '62. (MIRA 17:3)

BELEN'KIY, Dmitriy Mikhalevich; KOVAL', I.V., ved. red.

[Main conveyors] Magistral'nye konveiery. Moskva, Nedra,  
1965. 220 p. (MIRA 18:5)

SUSTER, M., prof. dr., DrSc.; CTSARIK, A.; HAVRILA, L.; KOVAL, J.;  
DEMKOY, J.; JABLONICKY S.; STOLINA, J.; SVATY, I.; VRZAL, J.;  
ZRUBEC, P.

Incidence of scleroma in eastern Slovakia. Cesk. otolaryng. 14  
no.1:10-13 F'65.

1. Otoryngologicka klinika Lekarskej fakulty University P.J.  
Safarika v Kosiciach (prednosta: prof. dr. M. Suster, DrSc.)  
a ORL oddelenia Obvodniho ustavu narodniho zdravi, Presov, Kosice,  
Humenne, Spisska Nova Ves, Michalovce, Poprad a Roznava.

KOVAL, J.

Unusual case of foreign body (raw potato) in the pharynx. Cesk.  
otolar. 6 no.3:144-145 June 57.

1. Otolaryngologicka klinika UK v Kosiciach, prednosta doc. MUDr.  
Michal Suster.

(PHARYNX, foreign bodies  
raw potato, surg. (Cz))

SUSTER, M.; HAVRILA, L.; KOVAL, J.

Course of severe esophageal wounds in current combined therapy!

Cesk. otolaryng. 11 no. 6: 370-371 D '62.

(ESOPHAGUS) (ESOPHAGITIS) (MEDIASTINITIS)

KOVAL, Jan, inz.; SLAVIK, Ivan, inz.

Possibility of using acetic acid for wood delignification.  
Papir a celulosa 18 no. 11:215-216 N°63.

1. Ceskoslovenska akademie ved, Chemicky ustav Slovenskej  
akademie vied, Bratislava.

KOVAL', L.A.

Selection of formulas for computing on calculating machines  
errors caused by the effect of relief in gravity prospecting.  
Izv. AN Kazakh. SSR Ser.geol. no.2:120-124 '62. (MIRA 15:6)  
(Calculating machines) (Gravity prospecting)

KAZHINA, A.O.; KOVAL', L.A.; LYAPICHEV, B.I.

"Electron mathematics" in the service of geophysics. Izv. AN  
Kazakh. SSR. Ser. geol. no.3:100-101 '62. (MIRA 15:7)  
(Electronic calculating machines) (Geophysics)

L 15 59-63	EW (1)/RDS/ES(v)	AJ/FTC	Pg -4/Po-4/Pe-4/Pq-4 S/0169/63/000/005/DO11/D018
ACCESSION NR: AR3012962			68
SOURCE: RZh. Geofizika, nos. 5098			✓
AUTHOR: Koval', L. A.			✓
TITLE: On the calculation of the effect of relief during a gravimetric survey			
CITED SOURCE: Tr. Kazakh. politekhn. in-ta, sb. 22, 1962, 173-177			
TOPIC TAGS: relief, Bouguer correction, intermediate layer, spherical layer, gravimetric survey			
ABSTRACT: Corrections for relief can be calculated on computers; hence the question is posed concerning the determination of the necessary radius for calculating the effect of relief and the compilation of maps of the effect of relief which correspond to the requirements of various surveys. Correction for the effect of relief is a supplement (heterovalent) to the correction for the attraction of the intermediate layer (Bouguer correction). It is proposed to calculate the effect of relief within the same limits in which the attraction of the intermediate layer on the near-to-the-spherical surface of the earth must be taken into consideration. The need for alteration of Bouguer's correction under mountain conditions is			
Card : /2			

1 155-9-63

ACCESSION NR: AR3002062

indicated. An exact formula is given which permits calculation of the amount of Borgier's correction for various values  $r = R_{sub 0} \gamma$ , where gamma is the central angle which determines the thickness of the layer;  $R_{sub 0}$  is the radius of the earth;  $r$  is the radius of the spherical layer. I. Yesakov

DATE ACQ: 12Jun63

SUI CODE: PH

ENCL: 00

Card 2/2

GRODINA, Yu. V.; SOFROSHENKOV, A. F.; KOVAL', L. A.

Wear resistance of combined coatings under the effect of hydro-abrasive friction. Izv. vys. ucheb. zav.; chern. met. 7 no. 4: 124-128 '64. (MIRA 17:5)

1. Sibirskiy metallurgicheskiy institut.

ACCESSION NR: AP4033704

S/0148/64/000/004/0124/0128

AUTHOR: Grdina, Yu. V.; Sofroshenkov, A. F.; Koval', L. A.

TITLE: Resistance of Combined Coatings During Hydroabrasive Wear

SOURCE: IVUZ. Chernaya metallurgiya, no. 4, 1964, 124-128

TOPIC TAGS: diffusion layer, heat treatment, hydroabrasive wear, calorization, titanization, chrome plating, siliconizing

ABSTRACT: In an earlier paper the authors investigated the properties of diffusion layers produced by combining chemical treatment with heat treatment, and they continue their research by reporting additional test results. Sleeves, checkers and segments were exposed to hydroabrasive wear. The treatment consisted of calorizing (950-10000 C) for 12 hrs., titanizing (1080 C) for 10 hrs., siliconizing (1080-1100 C) for 11 hrs. and chromizing (1150 C) for 8 hrs. All parts were ground, degreased and nitrided at 500-550-520 C for 70 hrs. The authors found that wear resistance depended not only on microhardness but also on microstructure, brittleness of the layer and test conditions in which pulp-on coal lines as well as hot steel runners were simulated. Although the method

Card 1/2

ACCESSION NR: AP4033704

appears somewhat complicated, it is recommended for many parts exposed to hydroabrasive wear. Orig. art. has: 4 figures and 2 tables.

ASSOCIATION: Sibirskiy metallurgicheskiy institut (Siberian Metallurgical Institute)

SUBMITTED: 08Jun63 DATE ACQ: 07May64 ENCL: 00

SUB CODE: MM NO REF Sov: 001 OTHER: 000

Card 2/2

BOBACH, P.G.; KOVAL', L.A.

Effect of prolonged administration of qualitatively different foods  
on periodical motor function of the small intestine. Vop. pit. 15  
(MLRA 9:7)  
no.2:5-11 Mr-Ap '56.

1. Iz Instituta fisiologii Kiyevskogo gosudarstvennogo universiteta  
imeni T.G. Shevchenko.

(INTESTINE, SMALL, physiology,  
eff. of food of various types on motor funct. (Rus))

(FOOD), affects,  
on small intestinal motor funct., eff. of prolonged  
admin. of various foods (Rus))

KDVAL', L.A.

Role of acetylcholine and adrenaline in the motor regulation of the  
small intestine. Nauk zap. Kyiv. un. 16 no.17;103-111 '57.  
(MIRA 13:2)

(ACETYLCHOLINE) (ADRENALINE) (INTESTINES)

KOVAL', L.O.; DOBROVOL'SKAYA, Z.O. [Dobrovols'ka, Z.O.]

Comparative characteristics of the effect of acetylcholine and carbocholine on the motor function of isolated sections of the small intestine. Nauk zap. Kyiv. un. 16 no.18:83-88 '57.

(MIRA 13:2)

(INTESTINES) (CHOLINE)

BOGACH, P.G. [Bohach, P.H.]; KOVAL', L.A. [Koval', L.O.]

Mechanism of reflex influences from the rectum on the motor activity of the small intestine. Fiziol. zhur. [Ukr.] 5 no.3:329-336 My-Je '59. (MIRA 12:10)

I. Naukovo-doslidnyi institut fiziologii Kiiv's'kogo derzhuniversitetu im. T.G. Shevchenka, viddil fiziologii travlennya i krovoobigu.  
(INTESTINES--INNERVATION)

17(3,4)  
AUTHOR:Koval', L. O. (Koval', L. A.)

SOV/21-59-7-25/25

TITLE:

Action of Adrenalin and Acetylcholin of the Blood on  
the Motor Function of the Small Intestine Through  
the Central Nervous System

PERIODICAL:

Dopovidi Akademii Nauk Ukrains'koi RSR, 1959, Nr 7,  
pp 806-810 (UkrSSR)

ABSTRACT:

On introducing adrenalin into the isolated blood stream of a dog's head connected with the body by nerve connections only, the motor activity of the intestine is inhibited (latent period of reaction 20 -45 seconds). After the inhibitory phase a second, excitation phase sometimes sets in. On introducing acetylcholin into the blood stream of the head, inhibition of the intestinal motor function also sets in after 10-120 seconds. Proceeding from these data and the results of experiments on stimulation of the vagus nerves on a background of the inhibitory phase after introducing acetylcholin before and after adrenalectomy of animals with a crushed spinal cord and severed

Card 1/2

SOV/21-59-7-25/25

Action of Adrenalin and Acetylcholin of the Blood on the Motor Function of the Small Intestine Through the Central Nervous System

vagus nerves (with undisturbed circulation) the motor function during the second phase of acetylcholin action is effected as a consequence of the action of the latter through the central nervous system and the influence of the adrenalin of the adrenals. There are 3 diagrams and 4 references, 3 of which are Soviet and 1 French

ASSOCIATION: Naukovo-doslidchyy instytut fiziologiyi tvaryn pry Kyivskomu derzhavnому universyteti, (Scientific-Research Institute of Animal Physiology at Kiev State University)

PRESENTED: D.S. Vorontsov, Member AS UkrSSR

SUBMITTED: February 5, 1959  
Card 2/2

KOVAL', L.A. [Koval', L.O.]

Effect of acetylcholine and adrenaline on the motor activity  
of the small intestine as related to functional state and  
age. Fiziologiya zhivotnykh [Ukr.] 6 no.2:213-220 Mr.-Ap '60.

(MIRA 13:7)

1. Institut fiziologii zhivotnykh Kyivskogo gosudarstvennogo  
universiteta im. T.G. Shevchenko, otdel fiziologii pishcheva-  
reniya i krovobrazheniya.

(INTESTINES) (CHOLINE) (ADRENALINE)

KOVAL', L. A.

Cand Biol Sci - (diss) "Role of adrenalin and acetylcholine in the control of the motor function of the small intestine." Kiev, 1961. 18 pp; (Academy of Sciences Ukrainian SSR, Division of Biological Sciences); 110 copies; price not given; list of author's works at end of text (12 entries); (KL, 6-61 sup, 207)

KOVAL', J.A.

Use of electronic computers for calculating gravimetric terrain.  
Izv. AN Kazakh.SSR. Ser. geol. nauk no.4:96-107 '63. (MIRA 16:9)

1. Kazakhskiy politekhnicheskiy institut, Alma-Ata.

KOVAL', L.G. [Koval', L.H.]

Zooplankton of the northwestern part of the Black Sea. Nauk.  
zap. Od.biol. sta no.3:27-44'61. (MIRA 16:6)  
(BLACK SEA—ZOOPLANKTON)

KOVAL', L.G. [Koval', L.E.]

Zooplankton in estuary areas of the northwestern part of the Black  
Sea in 1954-1957. Nauk.zap.Od.biol.sta. no.1:34-51 '59. (MIRA 14:7)  
(Black Sea—Zooplankton)

KOVAL', L.G.

Ecologic characteristics of the development and distribution of zooplankton in the northwestern part of the Black Sea. Vop. ekol. (MIRA 16:6)  
5:98-100 '62.

1. Biologicheskaya stantsiya AN UkrSSR, Odessa.  
(Black Sea—Zooplankton)

"APPROVED FOR RELEASE: 06/14/2000

CIA-RDP86-00513R000825510014-5

KOVAL', L.G. [Koval', L.H.]

Zooplankton of Eastern Sivash. Pratsi Inst. hidrobiol. AN URSS  
no.35:41-49 '60. (MIRA 14:4)  
(Sivash--Zooplankton)

APPROVED FOR RELEASE: 06/14/2000

CIA-RDP86-00513R000825510014-5"

"APPROVED FOR RELEASE: 06/14/2000

CIA-RDP86-00513R000825510014-5

KOVAL', L.G. [Koval', L.H.]

Zooplankton of Lake Molochnoye. Pratsi Inst. gidrobiol. AN UkrSSR  
no. 35:138-142 '60. (MIRA 14:4)  
(Molochnoye, Lake--Zooplankton)

APPROVED FOR RELEASE: 06/14/2000

CIA-RDP86-00513R000825510014-5"

KOVAL', Lev Mikhaylovich; YAROSLAVTSEV, Boris Alekseyevich; GUROV, S.,  
red.; KUZNETSOVA, A., tekhn. red.

[From small-scale to over-all mechanization] Ot maloi mekha-  
nizatsii - k kompleksnoi. Moskva, Mosk. rabochii, 1961. 102 p.  
(MIRA 15:2)

(Assembly-line methods)  
(Moscow--Electric machinery industry)

KOVAL', L.M.

Modernizing the clutch of the hydraulic drive of a press.  
Mashinostroitel' no.8:14 Ag '62. (MIRA 15:8)  
(Power presses--Hydraulic drive)

KOVAL', L.P.

Scientific and technical conference at the Volgodonsk Combine of  
Synthetic Fat Substitutes. Masl.-zhir. prom. 27 no.9:48 S '61.  
(MIRA 14:11)

1. Vneshstatnyy korrespondent zhurnala "Masloboyno-zhirovaya promy-  
shlennost'".  
(Oils and fats) (Acids, Fatty)

KOVAL', M. A.

KOVAL', M. A. — "Roentgeno-Vasographic Method in the Evaluation of the Changes of the Permeability of the Blood-Carrying Vessels in the Presence of Inflammation, (Experimental-Clinical Investigation)." Crimean State Med Inst imeni I. V. Stalin, Simferopol', 1955. (Dissertation For the Degree of Candidate in Medical Sciences).

SG: Knizhnaya letopis', No. 37, 3 September 1955

KOVAL', M.A. kand. med. nauk.

Roentgen vasographic date of vascular changes in experimental tuberculous inflammation. Vest. rent. i rad. 33 no.6:22-28 N-0 '58. (MIRA 12:1)

1. Iz Nauchno-issledovatel'skogo instituta travmatologii, ortopedii i protezirovaniya g. Stalino (dir. - kand. med. nauk. N.V. Novikov).  
(TUBERKULOZIS, exper.)

vasc. changes, angiography in dogs (Rus))

(ANGIOGRAFIY  
in tuberc., determ. of vasc. changes in dogs (Rus))

KOVAL', M.A., kand.med.nauk (Yalta, ul. Voykova, d.10); VLASOV, I.V.;  
Dobritskaya, A.A.

Use of vasography in inflammatory diseases of the extremities. Nov.  
khir.arkh. no.5:68-74 S-0 '59. (MIRA 13:3)

1. Khirurgicheskiye etdeleniya Alushtinskoy gorodskoy bol'nitsy i  
Yaltinskoy gorodskoy bol'nitsy. Nauchnyy rukovoditel' - zaveduyu-  
shchiy kafedroy fakultetskoy khirurgii Krymskogo meditsinskogo  
instituta prof. P.P. Tsvarenko.  
(EXTREMITIES, LOWER) (ANGLOGRAPHY)

KOFAL', M.M., inzh.-mekhanik

Universal tractor-mounted loader. Mekh. sil', hosp. 9 no. 8:3 of cover  
Ag '58.

(Loading and unloading)  
(Agricultural machinery)

BONDARENKO, M.G. [Bondarenko, M.H.]; VORONEZHSKIY, V.I. [Voronezhs'kyi, V.I.]; KITAYTSEVA, Z.P.; KOVAL', M.M.; KOLODA, V.D.; KORSAKOV, O.O.; KREMINSKAYA, Ye.D. [Kremins'ka, Ye.D.]; KUKTA, G.M. [Kukta, H.M.], inzh.-mekhan.; PIVOVAR, S.G. [Pivovar, S.H.]; SOLOVEY, V.I.; OLEFIRENKO, G.A. [Olefirenko, H.A.], red.; GULENKO, O.I. [Hulenko, O.I.], tekhn.red.

[New agricultural machines] Novi sil's'kohospodars'ki mashyni. Kyiv, Derzh.vyd-vo sil's'kohospodars'koi lit-ry URSR, 1959. 231 p.  
(MIRA 13:4)

(Agricultural machinery)

KOVAL', M. N., Engineer

"New Principles for Designing the General Layout of Heavy Machine Building Plants (With a Centralized Storeshouse of Basic Raw Materials)." Sub 25 Jun 51, Moscow Order of the Labor Red Banner Construction Engineering Inst imeni V. V. Kuybyshev

Dissertations presented for science and engineering degrees in Moscow during 1951.

SO: Sum. No. 480, 9 May 55

KOVAL', M.N., kand. tekhn. nauk.

Joint operational technology of industrial, water, and railroad  
transport. Zhel. dor. transp. 40 no.2:38-42 F '58. (MIRA 11:3)  
(Transportation) (Railroads--Freight)

KOVAL', M.N., kand.tekhn.nauk

Drawing up general plans for heavy industrial plants in  
connection with water and rail supply of raw materials.  
Trudy MIIT no.105:243-287 '58. (MIRA 11:9)  
(Material handling)

Koval', M. P.: Master Vet Sci (diss) -- "The fauna and cellulosolytic activity of the microflora of scars on cows". Leningrad, 1959. 17 pp (Leningrad Vet Inst of the Min Agric USSR), 150 copies (KL, No 13, 1959, 110)

KOVAI, M.P., absent.

Use of a nutrient salt solution in dyapepsia of calves.  
Veterinariia 43 no. 5:83. My '65. (MIRA 18:6)

1. Grodzenskiy sel'skokhozyaystvennyy institut.

KOVAL', M.P., kand. veter. nauk; VAS'KO, I.V.

Case of carbamide poisoning in cows. Veterinariia 41 no. 10:49-50  
0 '64.  
(MIRA 1871)

1. Grodzenskiy sel'skokhozyaystvennyy institut (for Koval'),
2. Glavnyy veterinarnyy vrach stantsii po bor'be s bolaznymi zhivotnykh Grodzenskogo proizvodstvennogo upravleniya (for Vas'ko).

STASHEVSKIY, Dmitriy Nikolayevich [Stashev's'kiy, D.N.]; KOVAL', M.Y.,  
red.; RAKHLENKA, N.P., tekhn.red.

[International significance of the seven-year plan for the  
development of the national economy of the U.S.S.R.] Mishna-  
rodne znachenija siedmichnogo planu rozvitiyu narodnogo hospo-  
darstva SRSSR. Kyiv, Vyd-vo Akad.nauk UkrSSR, 1959. 72 p.

(Russia--Economic policy)

(MIRA 13:3)

GANZHA, Ivan Fomich [Hanzha, I.Kh.]; KOVAL', M.V., red.; BUNIY, R.O.,  
tekhn.red.

[First collective farms in the Ukraine, 1917-1920] Pershi  
kolektyvni hospodarstva na Ukrayini, 1917-1920 rr. Kyiv, Vyd-vo  
Akad.nauk URSR, 1960, 154 p. (MIRA 13:9)  
(Ukraine--Agriculture, Cooperative)

(1) L 12087-56 EWT(m)/EWA(1)/T/EWP(1)/EWP(k)/EWP(r)/EWP(b)/EWA(c) IWP(c)  
 ACC NR: AP6000610 SOURCE CODE: UR/0129/65/000/012/0040/0043  
 MJW/JD/HW  
 AUTHOR: V. I. Yams, O. S.; Pol'shov, N. M.; Koval', M. Ya.  
 99,55 99,55 99,55  
 ORG: Nikopol' Southern Tube Plant (Nikopol'skiy yuzhnотрубный завод)  
 99,55  
 TITLE: Effect of temperature and rate of heating on the grain size of Kh18N12T steel  
 44,55  
 SOURCE: Metallovedeniya i termicheskaya obrabotka metallov, no. 12, 1965, 40-43  
 TOPIC TAGS: grain size, steel, metal tube, titanium, heat treat furnace/  
 Kh18N12T steel  
 ABSTRACT: Annular specimens of Kh18N12T steel (0.08% C, 1.41% Mn, 0.66% Si, 0.017% P,  
 0.007% S, 17.85% Cr, 11.39% Ni, 0.5% Ti), cut from cold-rolled boiler superheater  
 tubes, were heated at 800-1200°C, or increasing the temperature by 50-100°C at a time,  
 for 30 min, and air-cooled, with the object of determining the conditions under  
 which grain homogeneity can be maximized. Experiments with the use of different fur-  
 naces (muffle, induction, box, continuous roller) showed that the most suitable fur-  
 nace for this purpose is the continuous roller furnace, (furnace length 10 m, travel-  
 ing rate of tubes 0.4 m/min, temperature 1200-1230°C). The higher the heating rate,  
 the higher is the temperature needed to obtain a roughly identical grain size. In  
 addition, the effect of Ti on grain size was investigated on specimens of work-  
 hardened tubes from two melts and compared with specimens of Ti-free Kh18N10 steel;

Card 1/2

UDC: 621.785.16:620.186.5:669.14.018.84

L 12087-66

ACC NR. AF 6000610

2

It was found that in Ti-containing steel the grain size decreases with decreasing temperature and increases with increasing temperature at a faster rate than in Ti-free steel. In addition, given the same heating temperatures, the microstructure of Kh18Ni2T steel containing 0.44% Ti becomes more coarse-grained than the microstructure of the same steel containing 0.65% Ti. This is attributable to the retarding effect of titanium carbides on grain growth. As the titanium carbides become dissolved at elevated temperatures, an intensive grain growth sets in so that then, at temperatures above 1150°C, the grain size in Ti-containing steel (Kh18Ni2T) becomes much larger than in Ti-free steel (Kh18Ni10).  
Orig. art. has: 2 tables, 5 figures.

SUB CODE: 11, 13, 20/ SUM DATE: none/ ORIG REF: 003/ OTH REF: 000

Card 2/2

Koval', N.

Moldavia's products on the international market. Vnesh. torg. 42  
no.8:44-45 '62. (MIRA 15:9)

1. Predsedatel' Gosudarstvennogo planovogo komiteta Soveta  
Ministrov Moldavskoy SSR.  
(Moldavia—Commerce)