

AUTHOR

GINZBURG, V.L., FAYN, V.M.

PA = 2080

TITLE

On the Quantum Effects occurring on Interactions of Electrons with High Frequency Fields in Resonant Cavities (O kvantovykh effektakh pri vzaimodeystvii elektronov c vysokochastochnymi polyami v polykh rezonatorakh).

PERIODICAL

Zhurnal Eksperimental'noi i Teoret. Fiziki, 1957, Vol 32, Nr 1,

pp 162-164 (U.S.S.R.)

Received 3/1957

Reviewed 4/1957

ABSTRACT

The authors investigated the following problem in classical manner: At the moment A = 0 with the kinetic energy $K_0 = mv_0^2/2$ a non-relativistical electron enters the resonator and leaves it at the moment t=7 with the energy $K_T = mv_T^2/2$. For reasons of simplicity the electric field E in the resonator on the path of the electron is assumed to be homogeneous and parallel to the velocity of the electron (such a case is absolutely real). If $E = E_1 \cos \omega t + (E_8 + E_0) \sin \omega t$ applies, m(dv/dt) = eE and $v_7 = v_0 + (e/m\omega) \left[E_1 \sin \omega t + (E_8 + E_0) (1 - \cos \omega r) \right]$ is obtained. Here E_1 and E_2 denote chance quantities and $E_1 = E_8 = 0$ and $E_1^2 = E_8^2 = V/d^8$ are assumed to apply, d denotes the path to be covered by the electron (thickness of the resonator) and \overline{V}^R denotes the mean square of the fluctuationvoltage. The averaging is carried out over the corresponding assemblies of the identical systems. The field in the resonator is assumed to influence the movement of the electrons only to a small extent so that the terms of the order of magnitude e^2 may be taken to be sufficient. Under these circumstances $(\Delta K_T)^2 = e^2 \overline{V}^2 \left[(\sin(\omega T/2)/(\omega T/2)) \right]^2$ applies. For

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APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R CIA-RDP86800513R00051 Frequency Fields in Resonant Cavities.

the dispersion of velocity then $(\overline{\Delta v_T})^2 = (\overline{\Delta K_T})^2 m^{-2}v_0^2$ applies. If $\omega_T \ll 1$, then $(\overline{\Delta K_T})^2 = e^{2\sqrt{2}}$ applies. If oscillations of different frequencies exist in the resonator, $(\Delta K_T)^2 = e^2 \int_0^\infty |\overline{V}_{\omega}|^2 [\sin(\omega v/2)/(\omega v/2)]^{-2} d\omega$, $\overline{V}^2 = \int_0^\infty |\overline{V}_{\omega}|^2 d\omega$ applies. For a slightly damping resonator with the frequency $\omega_0 = (IC)^{1/2}$ the following expression is found (proceeding from the general expression for $(\overline{\Delta K_T})^2$): $(\overline{\Delta K_T})^2 = (\overline{\Delta V_{\omega}})^2 (\overline{\Delta V_{\omega}})^2 = (\overline{\Delta V_{\omega}})^2 (\overline$ mechanical perturbation theory, their calculations, however, are more complicated and are suited only for the range of small damping. The entire quantum-like effect in the problem of the passage of an electron through a resonator is based on the consideration of the quantum-like fluctuations of radiation in the resonator and especially of the zero oscillations with the energy $\hbar \omega/2$. (Without images)

ASSOCIATION

Physical Institute "P.N.LEBEDEV" of the Academy of Sciences of the USSR and the State University GOR!KIY.

PRESENTED BY

SUBMITTED 21., 9. 1956

AVAILABLE Library of Congress

Card 2/2

AUTHOR:

GEJLIKMAN, B.T., GINZBURG, V.L.

PA - 2053

TITLE: PERIODICAL: In Memory of S.Z.BELEN KIJ.

Uspekhi Fizicheskikj Nauk, 1957, Vol 61, Nr 1, pp 129-132(U.S.S.R.)

Received: 3 / 1957

Reviewed: 3 / 1957

ABSTRACT:

On September 21st 1956 SEMEN ZACHAROVIC BELEN'KIJ, a well-known theoretical physicist whose reputation is mainly based on his works on cosmic rays, died at the age of 41. S.Z.BELEN'KIJ was born in Moscow on the 14. June, 1916; after his leaving examination (1931) he worked two years in an electric plant. From 1933 to 1938 he studied with great success at the physical faculty of Moscow university and worked from 1941 to 1943 at the Central Aero-Hydrodynamic Institute. He then undertook a dissertation at the Physical Institute of the Academy of Science of the USSR and became the head of one of the theoretical sectors of this institute in 1948. BELEN'KIJ wrote his first scientific work (on the scattering of X-rays) during his last university term. These works showed the author's aptitude for theoretical physical work. In 1938, as an aspirant at Moscow university, he was able to concentrate his whole attention on the problem I.E. TAMM had asked him to solve (theory of cascade showers in cosmic rays). BELEN'KIJ was able to determine the spectrum of cascade electrons and this work served as a basis for his candidates' dissertation written in 1941. (Reviewer's comment: In the USSR there are candidates - and doctors disserta-

Card 1/2

In Memory of S.Z.BELEN'KIJ.

PA - 2053

tions). In connection with the problems solved by experiments BELEN'KIJ undertook quite a series of further investigations, in which the development of the theory of cascade showers was practically completed. In 1948 his monography "Cascade processes in cosmic rays" was published. BELEN'KIJ's works on the cascade theory are of fundamental importance. Though it is true that the investigations of the cascade showers form the main part of BELEN'KIJ's entire activity, they were not his only domain of research; he also dealt with hydrodynamic problems as well as with the hydrodynamic and statistical theory of the multiple production of particles at high and superhigh energies. Recently BELEN'KIJ dealt with the nuclear cascade processes and with the phenomenological theory of the scattering of nucleons by nucleons at high energies. For his merits in solving applied problems he was awarded the Lenin order and the Stalin prize.

ASSOCIATION:

Not given.

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SUBMITTED:

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Card 2/2

CIA-RDP86-00513R00051673(

AUTHOR TITLE

J.NI OFF

GERSHMAN, B.N., GINZBURG, V.L., DENISOV, N.G.

53-4-4/7

The Propagation of Electromagnetic Waves in a Plasma (in the Ionosphere). (Resprestraneniye elektromagnitnykh voln v plasme (ionosfere) -Russian).

Uspekhi Fiz. Nauk, 1957, Vol 61, Nr 4, pp 561-612 (U.S.S.R.) Received 6/1957

Reviewed 7/1957

ABSTRACT

PERIODICAL

Starting out from the monograph by Ya.L.Al'pert, V.L.Ginsburg, El.Feynberg "The Propagation of Radio Waves" (Raspostraneniye rediovoln - Gostekhizdat, 1953, the paper under review deals with some problems of this field which have been clarified to a certain extent since the publication of the monograph. The consideration of the heat motion of electrons in a homogeneous medium in the magnetic field leads to the ocurrence of plasma waves, the consideration of the heat motion of ions, on the other hand, results in low-frequency magnetohydrodynamic and quasi-acoustic waves, both with dispersion. In inhomogeneous media it is possible that we have cases where the approximation of geometrical optics is no more permissible and where an interaction of waves takes place which would be independent in the homogeneous or quasi-homogeneous case. This is the case in the absence of a magnetic field at vertical incidence in the proximity of the reflection point and at oblique incidence in the proximity of the point (cJ) = 0, at the existence of a magnetic field at a small angle between the wave normal and the magnetic field (multiplication of the reflected radio signals), and at the beginning of the layer where the concentration of the electrons still is small. For the latter case the paper under review computes the boundary polarization of the short waves which leave the ionosphere for a certain model of the ionosphere, but it is unable to offer any new information about the ionosphere.

APPROVED FOR RELEASE: Thursday, July 27, 2000

GINZBURG, VL.

· 53-2-2/9

AUTHOR: TITLE:

The Origin of Cosmic Radiation. (Proiskhozhdeniye kosmicheskikh

Uspekhi Fis. Nauk. 1957, Vol 62, Nr 2, pp 37-98 (U.S.S.R.)

ABSTRACT:

PERIODICAL:

The theory of the origin of cosmic radiation as expounded in the present paper is based upon radioastronomic and, of course, also on other experimental data. However, the present paper by no means claims to be complete either in a historical respect or with respect to experimental data. The paper is arranged as follows:

1.) The primary cosmic rays near the earth.

2.) The nature of the bremsstrahlung of cosmic radiation and the distribution of cosmic radiation in the galaxy.

3.) The motion of cosmic particles in the interstellar medium. 4.) The supernovae and the novae as sources of cosmic radiation.

Summarizing it may be said that the assumption concerning the acceleration of cosmic particles by explosions of supernovae is permitted within the framework of existing knowledge and sufficient in order to explain all known facts. It stands to reason, however, that there is as yet no guarantee that the supernovae and novae furnish the entire energy. There is further no proof against the existence of any other sources of energy or against an additional acceleration of particles in the inter-

Card 1/2

APPROVED FOR RETEASES HURSDAY, JULY 27, 2000 CIA RDP86-00515R0005107

AUTHOR

GINZBURG, V.L.

53-1a-7/18

TITLE

The Use of Artificial Satellites for the Purpose of the Verification

of the General Relativity

(Iepol'zovaniye iskusstvennykh sputnikov zemli dlya proverki obshchey te-

orii otnositel'nosti. Russian)

PERIODICAL

Uspekhi Fiz. Nauk, 1957, Vol 63, Nr la, pp 119 - 122 (U.S.S.R.)

ABSTRACT

The author at first gives a short survey of the present stage of the problem of the experimental verification of the general relativity. The perihelion displacement, the gravitational displacement of the spectral lines, and the deflection of light rays by a field of gravitation are discussed in short. The effects predicted by the theory were observed, but in spite of this additional confirmations of the theory would be desirable. The astronomical methods used up to now have not yet been fully exploited. New ways, however, have to be found which permit a more rapid and more exact verification. Some possibilities in this direction are offered by making use of artificial satellites.

The perigee of an artificial satellite will be displaced just like the perihelion of the planets. The displacement of the perigee is, however, considerably greater than even is the case of the planet Mercury, and attains values of about 1500 " per century for satellites near the earth. This effect increases the nearer the satellite is to the earth.

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The Use of Artificial Satellites for the Purpose of the Verification of the General Relativity

The orbit of such earth-near satellites can be determined far more accurately than the orbit elements of the planet Mercury and the relativistic effect can apparently be observed without difficulties. On the other hand, however, the orbit of the satellite, even if the relativistic effects are neglected entirely, is not exactly elliptical because the air resistance in the ionosphere, the nonspherical distribution of the masses on the earth, and perturbation by other celestial bodies, especially the moon, act upon the motion of the satellite. Perturbation by the moon can easily be taken into account, but not the other influences mentioned here. It is not yet possible to say whether the orbit of an artificial satellite may, for the purpose of verifying the relativistic effects, be predetermined with sufficient exactitude. V.L. GINZBURG, however, has no doubt as to the existence of such a possibility and also points out an effect of the general relativity which, in principle, may be observed by studying the orbit of the satellite. An additional perihelion displacement of the satellite and the motion of the node of the satellite's orbit is concerned, which may be caused by the revolution of the earth. This very interesting effect of the general relativity attains 50 " during a century for artificial satellites and

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The Use of Artificial Satellite for the Purpose of the Verfication of the General Relativity

is therefore of the same order of magnitude as the entire relativistic effect in the case of the planet Mercury. This revolution effect, with a sufficiently high accuracy of measurement, could be separated from the total effect. The artificial satellite may also be used for measuring the shifting of the frequencies within the radar domain, which are due to gravitation. Within the domain of visible frequencies such observations are not yet possible, and in the case of satellites near the earth (h 800 km) not even in the radar domain. For satellites near the earth even the quadratic Doppler effect is greater than the shifting of frequency due to gravitation, and this, of course, applies to a higher extent to the linear Doppler effect. Therefore satellites are necessary for the purpose of the verification of frequency shifts due to gravitation, which must be sufficiently far from the earth (h $\approx r_{+}$). Perhaps, however, this will also be possible for satellites which are nearer to the earth.

Thus, the use of artificial satellites offer very attractive possibilities for a further examination of the general relativity. (No illustrations).

Card 3/4

GINZBURG, Y.L.

53**-1a-**9/18

AUTHOR

VERNOV, S.N., GINZBURG, V.L., KURNOSOVA, L.V., RAZORENOV, L.A.,

FRADKIN, M.I.

TITLE

The Investigation of the Composition of Primary Cosmic Radiation (Issledovaniye sostava pervichnogo kosmicheskogo izlucheniya. Russian)

PERIODICAL

Uspekhi Fiz. Nauk, 1957, Vol 63, Nr la, pp 131 - Nr lb ựp 148 (U.S.S.R.)

ABSTRACT

According to the data available at present, cosmic radiation consists of protons, a-particles and, to a far less extent, of heavy nuclei. The distribution of the nuclei with Z > 2 has as yet not been investigated sufficiently well and also other problems are still to be solved. Rockets are not suited for such measurements because their time of flight outside the atmosphere is too short. By means of artificial earth satellites, however, the necessary statistical material for the investigation of rarely occurring heavy nuclei can be obtained. One of the most important problems concerns the numerical ratio between the currents of the light nuclei Li, Be, B and the nuclei C, N, O, F. By experimental determination of this ratio the various theories concerning the creation of cosmic radiation can be confirmed or rejected, If the particles of the cosmic radiation in the clouds of the supernovae are accelerated, a value \geq 0,1 is obtained for the ratio (Li, Be, B) / (C, N, O, F). In the case of this theory the ratio can also be somewhat higher, but never .lower than O,1. The data at present obtained for this ratio contradict each other. The problem whether or not nuclei with Z > 30 exist in cos-

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The Investigation of the Composition of Primary Cosmic Radiation
mic radiation can also be solved by means of artificial earth satellites.
The existence of such nuclei in cosmic radiation would, on account of
its large interaction cross section and the short range in the inter-

stellar space, indicate an exceptionally large amount of heavy elements existing in the sources of cosmic radiation.

The experimental data on the composition of primary radiation. The results of the experiments carried out in 1952 - 1953 have already been published in form of a collection of articles. The respective results obtained within the last years have been compiled in two tables. The importance of the geographical location of the place of observation in the case of equal geomagnetic latitude is pointed out. From the point of view of determining the anergy spectrum of the various nuclear groups in primary cosmic radiation, with the help of artificial earth satellites afford great possibilities, because in this way the intensity of the fluxes of the particles with various energies (even at different widths) can be determined by means of the same devices. This, naturally, will considerably increase the reliability of the data obtained concerning the energy spectrum of the primary nuclei. One of the most interesting problems of primary cosmic radiation is the determination of

Card 2/4

The Investigation of the Composition of Primary Cosmic Radiation the amount of the nuclei of the group Li, Be, B. The experimental method for the study of the charge spectrum of nuclei in primary cosmic radiation. Such methods are of advantage as do not . discriminate the particles with respect to their charge and mass. The use of particle counters in the case of which, on the occasion of the passage of a particle, the produced pulse depends upon the charge of the particle, forms part of this method. The application of such devices to an artificial earth satellite is, besides, of advantage in-so-far as the measured data can be telegraphed to the earth. The disadvantages of methods which are based upon the ionization of a medium by rapidly charged particles, are enumerated. The CHEREKOV counter is free from such disadvantages. The conditions to be fulfilled when measuring by this method, are enumerated. The apparatus is disdussed on the basis of a drawing. During the time of observation of one week about 1000 nuclei with $Z \ge 6$ cm, 7000 c-particles and a corresponding number of Li-, Beand B-nuclei can be registered. For the experiments it is intended to register the differential spectrum of the nuclei with respect to Z in the interval from the a-particle up to oxygen. Such a method is realizeable only if the device is able to solve every peak belonging to the various values of Z. The use of artificial satellites offers new possi-

Card 3/4

The Investigation of the Composition of Primary Cosmic Radiation bilities for the investigation of the primary cosmic radiation. viz. measuring of the primary proton flux, explaining of the part played by the "albedo" of the atmosphere of the earth, the determination of the lower limit of the electron-positron components, the study of the interaction of the primary particles with matter and the variations with respect to time of intensity. (With 7 illustrations and 4 tables).

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"Investigation of the Composition in Primary Cosmic Rays,"

"Investigation of the Composition in Primary Cosmic Rays,"

Fizicheskikh Nauk, Vol. 63, No. 1-2, p. 190, September 1957.

SO: JPRS Report No. 187

GINZBURG, V.L.

"The Use of Artificial Earth Satellites for the Purpose of Proving General Relativity Theory,"
1-2, p. 175, September 1957.

Uspekhi Fizicheskikh Nauk, Vol. 63, No.

SO: JPRS Report No. 187

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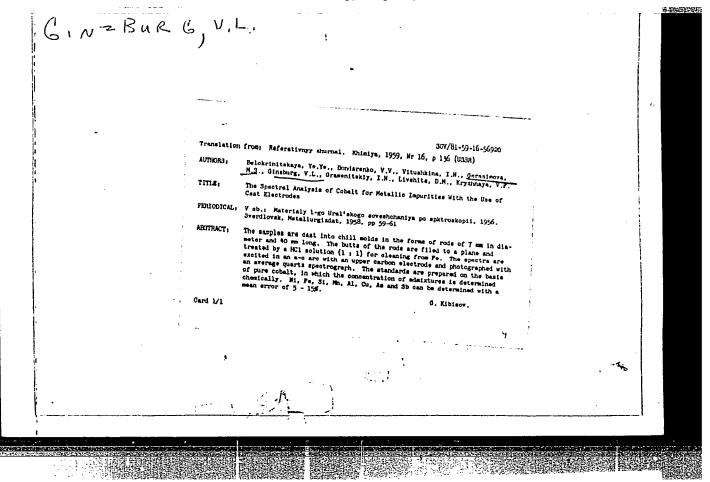
CINZBURG, V. L.

"The Role of Surface Energies in Superconductivity,"

"The Theory of Superfluidity," with L. P. Pitayevskiy

reports submitted but not presented at the Kamerlingh Onnes Conference, Leiden, Conf. on Low Temperature Physics, Leiden, 23-28 Jun 58.

Lebedev Physical Inst, AS USSR



Absorption and radiation of electromagnetic waves by a magnetically active plasma. Isv.vys.ucheb.sav.; radiofiz. 1 no.2:59-65 '58. (MIRA 11:11)

1. Issledovatel'skiy radiofizicheskiy institut pri Gor'kovskom universitete. (Radio waves)

06457

AUTHOR:

Ginzburg. V.L.

SOV/141-1-5-6-1/28

TITLE:

Radio-astronomy and the Origin of Cosmic Rays

PERIODICAL: Izvestiya vysshikh uchebnykh zaveděniy, Radiofizika, 1958

Yoll, Nr 5-6, pp 3 - 8 (USSR)

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ABSTRACT: This is a review paper which was first read during the symposium on radio-astronomy at the conference of the International Astronomical Union, which took place in

Paris, France, in August, 1958.

It is suggested that the main part of non-thermal cosmic radio emission has a synchrotron nature and is not generated in stellar atmospheres. The second problem considered is that of the mechanism of acceleration in the envelopes of supernovae and novae and certain other regions of interstellar space. There seems to be no doubt of the effectiveness of the statistical mechanism in envelopes. However, a number of important details must still be filled in. Korchak et al (Ref 15) have suggested that the statistical acceleration of nuclei with Z > 2 may become very effective compared with the acceleration of protons (as far as conditions of injection are concerned). This is important

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Radio-astronomy and the Origin of Cosmic Rays

in connection with the problem of the chemical composition of cosmic rays (Ref 7). The third problem discussed is that of the origin of electrons giving the general galactic radio emission. It is suggested that the material available at present cannot lead to a definite conclusion as to whether the electrons are primary or secondary. There are 20 references, of which 10 are Soviet, 8 English and 2 German.

Fizicheskiy institut im. P.N. Lebedeva AN SSSR ASSOCIATION: (Physics Institute imeni P.N. Lebedev of the Ac.Sc., USSR) and Gor'kovskiy gosudarstvennyy universitet (Gor'kiy State University)

SUBMITTED:

June 26, 1958

Card 2/2

06458 sov/141-1-5-6-2/28

AUTHORS:

Ginzburg, V.L. and Zheleznyakov, V.V.

TITLE:

On the Mechanisms of Sporadic Solar Radio Emission

PERIODICAL:

Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika,

1958, Vol 1, Nr 5-6, pp 9 - 16 (USSR)

ABSTRACT:

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This paper was read at the symposium on radio-astronomy during the conference of the International Astronomical Union, which took place in August, 1958 in Moscow. Possible coherent and incoherent mechanisms of sporadic solar radio emission in an isotropic and magneto-active coronal plasma are considered. The problem has been considered by the present authors in Refs 1-3 and the present paper is a summary of the results obtained. types II and III bursts, which are an important part of sporadic solar radio emission, are unpolarized or only weakly polarized. It is suggested that the magnetic field in the region where these bursts are produced is very low (possibly less than 1 0e). Under these conditions, the plasma may be considered as isotropic in the first approximation. The presence of frequency drift and other properties of types II and III bursts suggests that they are due to particle streams. In an

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On the Mechanisms of Sporadic Solar Radio Emission

isotropic plasma these streams excite only longitudinal waves. The existence in the plasma wave of a longitudinal electric field leads to an instability of the particle stream in the plasma and, as a result, coherent emission of plasma waves takes place. Incoherent and coherent emission of plasma waves takes place simultaneously but they have different frequency and angular spectra and depend on the parameters of the problem in a different way. It is argued that noncoherent emission of plasma waves by particle streams can, in principle, explain the appearance of type III bursts. It is, however, possible that when reabsorption is taken into account in detail, this mechanism may turn out to be unsuitable. Moreover, type II bursts cannot be connected with incoherent emission by particle streams since the particle velocity is not suitable. Coherent emission of plasma waves by particle streams can explain the properties of type III bursts and very probably also type II bursts. Since type I bursts are polarized, the analysis can only be

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On the Mechanisms of Sporadic Solar Radio Emission

carried out by taking the magnetic field into account. In this case, the incoherent emission by particle streams may be divided into Cherenkov radiation and synchrotron radiation. If reabsorption is taken into account it turns out that types I, II and III bursts cannot be associated with synchrotron radiation of electrons. Cherenkov effect cannot explain these bursts either. A charged particle stream moving in a magneto-active plasma is in general unstable and this leads to the coherent emission of ordinary and extraordinary waves. If the magnetic field is weak this coherent emission is practically identical with the coherent emission of plasma waves. In a stronger field (greater than 1 0e), the coherent radiation leaves the corona predominantly in the form of ordinary waves and hence it can be associated with type I bursts. In order to produce the observed type I bursts, the oscillations in the corona must have an amplitude of about 10 V/cm. How such oscillations are excited is not clear.

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On the Mechanisms of Sporadic Solar Radio Emission

There are 2 figures and 18 references, of which 4 are English and 14 Soviet.

ASSOCIATIONS: Fizicheskiy institut im. P.N. Lebedeva AN SSSR (Physics Institute im. P.N. Lebedev of the Ac.Sc., Gor'kovskiy gosudarstvennyy universitet (Gor'kiy State University)

SUBMITTED:

June 7, 1958

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GINZBURG, V.L.

25-2-2/43

DESCRIPTION OF THE PROPERTY OF

·AUTHOR:

Ginzburg, V.L., Corresponding Member, Academy of

Sciences, USSR

TITLE:

Artificial Satellites and the Theory of Relativity (Iskusstven-

nyye sputniki i teoriya otnositel'nosti)

PERIODICAL:

Nauka i Zhizn', 1958, # 2, p 7-12 (USSR)

ABSTRACT:

Vitaliy Lazarevich Ginzburg, Corresponding Member of the USSR Academy of Sciences, gave a lecture recently on the use of artificial satellites for the checking of Einstein's general theory of relativity. This, along with other problems concerning cosmic space, the atmosphere of the earth, etc., will be solved with the help of artificial satellites.

The author deals with three effects of the general theory of relativity which can be observed in our solar system:

1. Deviations in the movements of planets;

2. the deflection of rays of light, and

3. gravitational change in the frequency of spectral lines.

In general, the above aspects of the theory of relativity are borne out by the usual methods of astronomical observation,

however, there are still inaccuracies in measurement.

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"APPROVED FOR RELEASE: Thursday, July 27, 2000

CIA-RDP86-00513R00051673

Artificial Satellites and the Theory of Relativity

25-2**-**2/43

There are five sketches.

ASSOCIATION: Akademiya nauk SSSR (Academy of Sciences of the USSR)

AVAILABLE: Library of Congress

Card 2/2

AUTHOR: Ginzburg, V. L.

SOV/126-6-6-4/25

TITLE: Role of Surface Energy in Superconductivity (Rol' poverkhnostnoy energii v yavlenii sverkhprovodimosti)

PERIODICAL: Fizika metallov i metallovedeniye, 1958, Vol 6, Nr 6, pp 994-998 (USSR)

ABSTRACT: The author deals with properties of massive samples. At H = 0 and $H \ll H_{\text{ML}}$, $\phi^2 = \phi_{\text{CO}}^2 = \text{const.}$, and Londons' equation is obeyed. Here, H is the magnetic field, H_{ML} is the critical field for a massive sample, ϕ 's are electron wave-functions. At higher fields spatial non-uniformity of ϕ becomes important and this leads to the appearance of additional energy with density:

$$\frac{h^2}{2m} (\gamma \psi)^2 = \frac{H^2 \kappa \pi \delta \kappa^2}{4 \pi \kappa^2} (\gamma \psi_0)^2$$

where $\psi_0=\psi/\psi_{e\phi}$, h is Planck's constant, m is the Card 1/3

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SOV/126-6-6-4/25

Role of Surface Energy in Superconductivity

electron mass, δ_o is the depth of penetration of a weak magnetic field and $\kappa=\sqrt{2}~{\rm eH_{km}}~\delta_o^2/{\rm hc}$, e is the electron charge, c is the velocity of light. This energy is called surface energy in superconductivity. Its appearance is always connected with a boundary of a superconducting phase with either vacuum or normal phase. No other surface energy need be considered in the theory of superconductivity. For a massive metal, this surface energy density is necessary to find the value of σ_{nS} , which is the surface energy of a boundary between superconducting and normal phases. The $\sigma_{
m nS}$ is required in determination of the limits value of of supercooling and superheating (i.e. fields $H_{\mu 1}$ and $H_{\mu 2}$) and in determination of the dependence of the depth of penetration $\delta_{\rm H}$ on the magnetic field intensity. The author discusses in detail calculation of Hul and Hu2 He compares the calculated values with experimental data and concludes that the theory does not contradict the experiment. Further studies of the effect of surface energy in superconductivity require fuller empirical data, The

Card 2/3

SOV/126-6-6-4/25

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Role of Surface Energy in Superconductivity

paper is entirely theoretical. There are 2 tables and 18 references, 9 of which are Soviet, 8 English and 1 German.

ASSOCIATION: Fizicheskiy institut imeni P.N. Lebedeva AN SSSR (Physics Institute imeni P. N. Lebedev, Academy of Sciences USSR)

SUBMITTED: June 28, 1958.

Card 3/3

SOV-25-58-7-53/56 Ginzburg, V.L., Corresponding Member Sciences AUTHOR: of the USSR Academy of

Answers to Questions (Otvety na voprosy). Is "Time Travel" TITLE:

Possible (Vozmozhno li puteshestviye vo vremeni)

PERIODICAL: Nauka i zhizn', 1958, Nr 7, pp 77 - 78 (USSR)

Referring to a preceding article by the author on the Ein-ABSTRACT: stein theory of relativity, one reader asks whether "time travels" are possible. The author admits the theoretical

possibility, but positively denies its practical realiza-

tion.

1. Time--Theory

Card 1/1

SOV-26-58-8-1/51

AUTHORS:

Ginzburg. V.L., Associate Member of the USSR Academy of Sciences; Fradkin, M.I., Candidate of Physico-Mathematical Sciences

TITLE:

The Origin of Cosmic Rays (Proiskhozhdeniye kosmicheskikh

luchey)

PERIODICAL:

Priroda, 1958, Nr 8, pp 3-12 (USSR)

ABSTRACT:

Cosmic rays were discovered more than 40 years ago, but radioastronomical data permitted conclusions on their origin only
in 1950-1953. Primary cosmic rays have an energy of 109-1010 ev
with a small percentage reaching 1015-1018 ev. The primary
rays collide with the molecules of the atmosphere and form
the secondary rays which consist of mesons, electrons, and
photons. The primary rays can be observed at altitudes of
20 - 30 km by means of balloons, at 100 km by rockets, and
at higher altitudes by artificial satellites. The intensity
of the rays depends on the geomagnetic latitude. This latitude effect indicates that the cosmic rays consist of charged
particles. The distribution of the particles according to
energies (the energy spectrum) shows that the higher the
energy, the lower the number of particles (Figure 1). The
principal part of the primary rays is made up of protons.

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The Urigin of Cosmic Rays

507-26-58-8-1/51

Their intensity for particles with an energy higher than 1.4 · 109 ev is equal to 1 proton per om2/sec. In the primary cosmic rays are also heavier particles, like the nuclei of helium, carbon, oxygen, silicon, iron, etc. The relative composition of the primary rays is given in Table 1. Electrons, positrons, and photons, could not yet be detected in the primary rays. The place of origin of the cosmic radiation is investigated by means of radioastronomy. The radiation of the Galaxy in the radio wavelengths is a general radiation and radiation of single sources. These cosmic radic waves are due to the radiation of relativistic electrons which move in interstellar magnetic fields. In interstellar space, magnetic fields are present with 10-15-10-6 oersted. Electrons which move with an energy of 108-109 ev in this field emit a radiation in the radio wave length. The power of the magnetic field changes with the activity of the sun spots. The emitted radiation decreases the energy of the electrons. The energy of the particles with high energy changes to a greater degree than that of particles with low energy. The Galaxy is surrounded by a "corona" which emits radio waves. This may be observed in the nebula M31 in the Andromeda constellation which is in many respects similar to our own Ga-

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The Origin of Cosmic Rays

307-26-58-8-1/51

laxy. The radio waves are emitted by an area which is considerably greater than the visible area (Figure 3) and has the form of a spheroid, whereas the nebula is optically a flat disc. In our Galaxy, the area covered by cosmic rays has a radius of 50,000 light years. The space is filled by interstellar gas with a concentration of 0.01 - 0.03 particles per cm3. In some "clouds" it reaches a concentration of 10 particles per cm3. In collisions with the gas, the protons lose energy and form mesons. The fission of heavy nuclei leads to the formation of Li, Be, and B nuclei, the concentration of which near the earth's surface supplies information on the number of collisions at higher altitudes. The high energy of the cosmic particles is explained by the statistic mechanism. If heavy particles with high speed collide with light particles, the latter are accelerated by a transition of the energy from the heavy particles to the light ones. The single radio sources in the universe could be identified by galaxies or accumulations of galaxies. The radiations of these sources are caused by relativistic charged particles moving in the magnetic fields of the nebulae. The rebulae are the residues of super-novae. The energy of the electrons in their magnetic fields is 1045-1048 erg. Every 30 years, a

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The Origin of Cosmic Rays

30V-26-58-8-1/51

super-nova arises in the Galaxy. The power of the electrons generated then reaches 10^{36} - 10^{39} erg/sec. The energy lost by the electrons for the emission of radiowaves is 10^{39} erg/sec. It can be seen that the energy for radio emission is supplied by the super-nova. It is shown that the super-nova is also the source of protons and nuclei in the cosmic rays. The novae, one hundred of which arise every year, must also be considered as a source for cosmic rays. The novae and super-novae are accumulated principally near the center of the Galaxy. The cosmic rays are scattered by the chaotically distributed magnetic fields, so that they reach the earth from all sides. It is possible that a part of the cosmic rays, especially with an energy of more than 10^{15} ev, is of metagalactic origin.

There are 2 graphs, 2 tables, 1 photo and 1 Soviet reference.

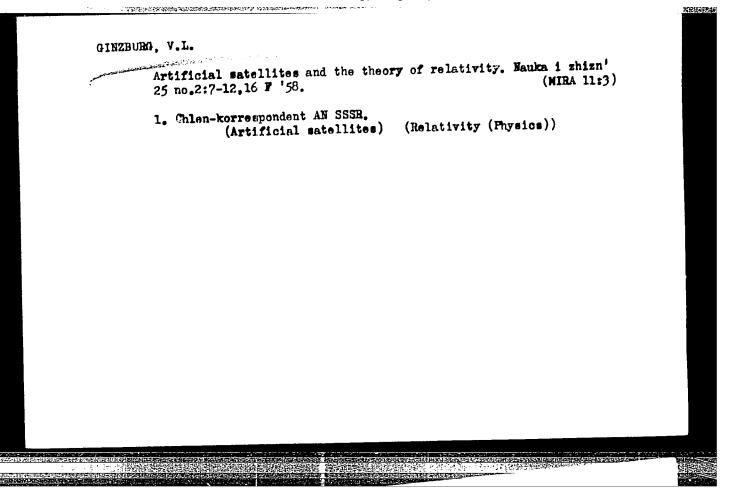
1. Cosmic rays—Sources 2. Cosmic rays—Analysis 3. Cosmic rays—Properties 4. Radio astronomy—Applications

Card 4/4

GINSBURG, V. L. (Acad. Sci. USSR)

Industrial

"Is Flight into Time Practicable?" Journal of Scientific and Research,
Vol. 17 A. P. 352, 1958 Council of Scientific and Industrial Research, India.



SINZBURG, V. L.

56-1-18/56

AUTHOR:

Ginzburg, V. L.

TITLE:

On the Annihilation and the Occurrence of Superconductivity in a Magnetic Field (O razruzhenii i vozniknovenii sverkhprovodimosti v magnitnom pole).

PERIODICAL:

Zhurnal Eksperimental noy i Teoretisheshor Piziki, 1950, Vol. 34, Nr 1, nr. 113-125 (USSR).

ABSTRACT:

The present paper investigates the transitions from the state of superconductivity into the normally conducting state and vice versa in the presence of an external magnetic field. At the outset the author deals with the general thermodynemic relations of superconductors. Among other formulae are given for the density of the free energy of the superconductor and a condition is given for the minimum of free energy. In a few pure metals the relation K (for aluminum for example K=0.05 at $T\rightarrow 0$). If such conditions revail, it is possible to let K = O for the sake of simplicity, which is just the case in the investigation of the annihilation of the superconductivity of samples with small dimensions. The author here investigates this case at full length, tracing the computations step ty step. The general investigation of the problem of

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on the Annihilation and the Occurrence of Superconductivity in a 56-1-18/56

APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-UUS. De very complicated and for this mostly to the range of K L/Oc ~lproves to CIA-RDP86-00513R0005

be very complicated and for this reason the author limits himself to the boundary case of $KL/\sqrt[4]{2}$ l. In actual experiment the overheating of the superconducting phase can be observed only with difficulty in massive samples, whereas with samples of medium dimensions this may be different. The assumption of an overheating does not contradict theory. On the strength of the results of this paper a further investigation of the problem of the boundaries of the domains of overheating and undercooling should be of interest, in particular in samples of medium dimensions. There are 6 figures and 16 references, 9 of which are Slavic.

ASSOCIATION:

Physical Institute imeni P. N. Lebedev AN USSR (Fizicheskiy institut imeni P. N. Lebedeva Akademii nauk SSSR).

SUBMITTED:

July 11, 1957

AVAILABLE:

Library of Congress

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... 6.4NZ Baky

56-1-46/56 On the Theory of the Rayleigh Dispersion of Light in Liquids

models for the range of radiofrequencies. As example the author investigates solid hollow, nonmetallic spherules having dispersing dipoles in their centers. The totality of these spherules will at any density of them disperse the radiowaves due to the fluctuations of the orientation of the dipoles. This dispersion even takes place in the complete absence or under neglect of the dispersion connected with the inhomogeneous spatial distribution of the spherules and with other factors. Under real conditions the antisymmetrical part of the dispersion in the case of weak' absorption is very small. In Rytov's papers (references 1, 4) factually only part of the dispersion is investigated and this still more restricts the applicability of the formulae obtained by Rytov. There are 5 references, all of which are

Physical Institute imeni P. N. Lebedev AN USSR (Fizicheskiy institut imeni P. N. Lebedeva Akademii nauk SSSR) ASSOCIATION:

October 19, 1957 SUBMITTED:

Library of Congress AVAILABLE:

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CIA-RDP86-00513R00051673(**APPROVED FOR RELEASE: Thursday, July 27, 2000**

"APPROVED FOR RELEASE: Thursday, July 27, 2000

CIA-RDP86-00513R00051673

AUTHORS:

Ginzburg, V. L., hitayerokiy, L. F. SOV/56-34-5-26/61

TITLE:

On the Theory of Superfluidity (F teeris sverkhtekuchesti)

CHRIODICAL:

Zhurnal *kaperimental noy i teoreticheskoy fiziki, 1959,

Vol. 34, Nr 5, pp. 1240-1245 (USSE)

ALOTHACT:

This paper deals with the properties of helium near the s-point. In the problem investigated in this paper the expansion parameter must be correlated to the density of the superfluent part φ of the liquid. φ is different from zero in He II and equal to zero in He I. Taking into account the quantum character of the phenomena in liquid helium, it is natural to choose as such a parameter the complex function $\psi(x,y,z) = \eta e^{i\varphi}$ which plays the rôle of an "effective wave function" of the superfluent part of the liquid. This paper deals only with those stable problems the normal part of which is assumed to be at rest. The velocity φ of the superfluent part is zero: v = 0. For this case the thermodynamic potential φ is given explicitly. The total thermodynamic potential is given as

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APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R00051673(

On the Theory of Superfluidity

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FdV with $F = (h^2/2m^2) |\nabla \psi|^2 + F_0(p,T/\psi|^2)$. By variation

with respect to ψ^* and ψ (considering the boundary condition $\psi=0$) an equation that is an analogon of the equation used in the phenomenological theory of superconduction is obtained. Finally the equation $-(h^2/2m)\Delta\psi + (eF/\hbar)\psi^2\psi = 0$ is obtained. To this equation belongs also the boundary condition ψ -0 which is to be used also for the free surface of helium. The thermodynamic potential F is expanded (as in the ordinary theory of phase transitions) into powers of

 $|\Psi|^2:F_0 = F_T(p,T)-\alpha|\Psi|^2 + (\beta/2)|\Psi|^4$.

The theory used in this paper can be used only in the immediate neighborhood of the \(\lambda\)-point. The second section of this paper deals with some special problems. First the authora investigate the properties of helium near a solid wall. In this case an additional surface energy appears. Then a helium film, i.e. a helium layer with the density d, is investigated. The temperature of the \(\lambda\)-transition in a helium film is lower than in great masses of helium. Finally a vortex in Re II is investigated. There are 1 figure and in references, 7 of which are Soviet.

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AND CLATTON: Pizicheskiy inctitut im. P. N. Lebedeva Akademii nauk SASS (Physics Institute imeni P. N. Lebedeva Akademii nauk SASS (Physics Institute imeni P. N. Lebedeva, AC DASS)

Institut fizicheskith problem in Redenii nauk MASSS (Institute for Problems on Physics, AS USSS)

December 10, 1957

1. Helium (Liquid)—Properties 2, Helium (Liquid)—Mathematical analysis 3, Low temperature research

SOV/56-34-6-28/51 AUTHOR: Ginzburg, V. L.

On the Electromagnetic Waves in Isotropic and Crystalline TITLE: Media With Respect to the Spatial Dispersion of the Dielectric Permeability (Ob elektromagnitnykh volnakh v izotropnykh i

kristallicheskikh sredakh pri uchete prostranstvennoy dis-

persii dielektricheskoy pronitsayemosti)

Zhurnal eksperimental'noy i teoreticheskoy fiziki, 1958, PERIODICAL:

Vol. 34, Nr 6, pp. 1593-1604 (USSR)

The influence of the spatial dispersion may be taken into ABSTRACT:

account by the following expression for the relation between D and E:

 $D_{i} = \varepsilon_{ik}(\omega) E_{k} + \gamma_{ik}(\omega) \frac{\partial E_{k}}{\partial x_{f}} + \delta_{ik\ell m}(\omega) \frac{\partial^{2} E_{k}}{\partial x_{f} \partial x_{m}}$

The term with $\gamma_{ik\ell}$, which corresponds to the optical activity, is neglected. The author investigates plane waves and writes the above given equation in the form $D_i = \hat{\mathcal{E}}_{ik} E_k$, $\hat{\mathcal{E}}_{ik} E_i$

 $-\alpha_{ik\ell m} s_{\ell} s_{m} h^{2}$, $\alpha_{ik\ell m} = (\omega/c)^{2} \delta_{ik\ell m}$. s denotes the unit Card 1/3

SOV/56-34-6-28/51 On the Electromagnetic Waves in Isotropic and Crystalline Media With Respect to the Spatial Dispersion of the Dielectric Permeability

vector of the normal to the wave, $\hat{n} = n - ix - the complex$ refraction index. The above given expansions not always are sufficient, sometimes a different expression has to be used. In an isotropic medium the tensors $\hat{\epsilon}_{ik}$, ϵ_{ik} , $\alpha_{ik}\ell_m$, and β ik/m can be regarded as scalars and the above mentioned equations tion is written down in the form $\vec{D} = \hat{\xi} \vec{E}$, $\hat{\xi} = \xi - \alpha \hat{n}^2$. This equation and its solution is discussed for transverse and for longitudinal waves. The possibility of the observation of the new wave (which corresponds to a new radical for \hat{n}^2 caused by the account of the spatial dispersion) depends in a remarkable degree on the intensity of the absorption. In the following part of this paper the tensor a is specialized for media of various crystalline types. The phenomena in the media with various crystal symmetry are discussed in detail. The last part of this paper gives some remarks on the collective energy losses and on the Vavilov-Cherenkov effect. The new waves caused by the account of the spatial dispersion can be excited easily if the Vavilov-Cherenkov effect is used.

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sov/56-34-6-28/51

On the Electromagnetic Waves in Isotropic and Crystalline Media With Respect to the Spatial Dispersion of the Dielectric Permeability

There are 3 figures and 10 references, 8 of which are Soviet.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva Akademii nauk SSSR

(Physics Institute imeni P. N. Lebedev, AS USSR)

SUBMITTED: January 16, 1958

Card 3/3

"APPROVED FOR RELEASE: Thursday, July 27, 2000

CIA-RDP86-00513R00051673

21(8) . AUTHOR:

Ginzburg, V. L., Fayn, V. M.

sov/56-35-3-54/61

TITLE:

On the Radiation of Systems With Many Levels Which Move in a Medium With Super-Light-Velocity (Ob izluchenii sistem s mnogimi urovnyami, dvizhushchikhsya v srede so sverkhsvetovoy

skorost'yu)

PERIODICAL:

Zhurnal eksperimental noy i teoreticheskoy fiziki, 1958,

Vol 35, Nr 3, pp 817 - 818 (USSR)

ABSTRACT:

The present paper deals with interesting possibilities offered in connection with systems of many levels moving with a velocity greater than that of light. If, initially, the system was on a single level (e.g. the lowest energy level) it will be possible, in the course of time, to observe it in all those states into which it may pass over by direct or cascade-like radiation transition. Formulae are given for the degree of occupation of the levels and for the energy emitted into the unit solid angle in the unit of time. To the systems which have many levels there also belong the bunches of atoms or molecules with two suitable levels. The radiation of such bunches (which have dimensions smaller than the wave length)

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On the Radiation of Systems With Many Levels Which Move in a Medium With Super-Light-Velocity

507/56-35-3-54/61

is coherent with and similar to the radiation of a system in a magnetic field. However, the radiation (with a velocity greater than that of light) of such bunches as well as of single atoms and molecules or of para- and ferromagnetic particles is of hardly any practical importance. However, the radiation (with a velocity greater than that of light) of electrons moving along a magnetic field is perfectly real. In this connection, a metallic slowing-down system but also a dielectric or a plasma located near the bundle can play the part of this medium. Next, some details connected with this phenomena are given. A more detailed report on this Doppler-radiation of electrons moving with a velocity greater than that of light is intended to be given at a later date. There are 9 references, 8 of which are Soviet.

ASSOCIATION:

Gor'kovskiy gosudarstvennyy universitet (Gor'kiy State

University)

SUBMITTED:

June 30, 1958

Card 2/3

3(1)

AUTHORS:

Ginzburg, V. L., and Zheleznyakov, V. V. SOV/33-35-5-3/20

TITLE:

On the Possible Mechanisms of Sporadic Solar Radio Emission (Radiation in Isotropic Plasma) (O vozmozhnykh mekhanizmakh sporadicheskogo radioizlucheniya solntsa (izlucheniye v izotropnoy plazme))

PERIODICAL: Astronomicheskiy zhurnal, 1958, Vol 35, Nr 5, pp 694-712 (USSR)

ABSTRACT:

The authors discuss the coherent and incoherent mechanisms of sporadic solar radio emission in isotropic coronal plasma. They show that it is impossible or improbable to combine type II bursts and type III bursts with an incoherent plasma mechanism of radio emission, while the description by coherent plasma mechanisms leads to no contradiction. Because of polarization the consideration of type I bursts related to sunspots by isotropic plasma only is senseless. In a following note the case of magnetoactive plasma shall be considered. About the contents of both notes it was partly reported on November 27, 1957 at the Radioastronomical Committee of the Astronomical Assembly of the Academy of Sciences of the USSR. It is mentioned in a footnote that, according to a remark of D.A. Frank-Kamentskiy, the question

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On the Possible Mechanisms of Sporadic Solar Radio Emission (Radiation in Isotropic Plasma)

SOV/33-35-5-3/20

whether the transition of plasma waves into electromagnetic waves is essential for the dispersion of plasma waves at coronal electrons is investigated by A.A.Vedenov and R.Z.Sagdeyev. There is 1 figure, and 17 references, 13 of which are Soviet, 2 American, 1 Australian, and 1 German.

ASSOCIATION: Fizicheskiy institut imeni P.N. Lebedeva Akademii nauk SSSR (Physical Institute **imeni** P.N. Lebedev of the AS USSR) Radiofizicheskiy institut pri Gor'kovskom universitete imeni N.I. Lobachevskogo (Radiophysical Institute at the Gor'kiy University imeni N.I. Lobachevskiy)

SUBMITTED: April 23, 1958

Card 2/2

24(5) AUTHORS:

Ginzburg, V. L., Eydman, V. Ya.

TITLE:

On the Cherenkov Radiation of Dipole Moments (O Cherenkovskom

SOV/56-35-6-28/44

izluchenii dipolnykh momentov)

PERIODICAL:

Zhurnal eksperimental'noy i steofeticheskoy fiziki, 1958,

Vol 35, Nr 6, pp 1508-1512 (USSR)

ABSTRACT:

Bunches of particles with dimensions sufficiently small with respect to the wave length in the medium givet the same Cherenkov radiation as point particles with a corresponding charge and multipole moments. Therefore, the investigation of the Cherenkov radiation of magnetic and electric dipoles is of interest irrespective of the fact that it is only very weak for separated particles (electrons, neutrons). With respect to the question of the Cherenkov radiation of the magnetic moment, contradictory opinions are, however, found to be expressed in publications (Refs 1-6). In this connection the authors developed a calculation method, which differs somewhat from that used in earlier papers (Refs 2-4). It is first developed for the Cherenkov radiation of electric and magnetic dipoles moving in a continuous medium, and further for that of dipoles moving in channels or gaps ($\varepsilon = \mu = 1$).

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On the Cherenkov Radiation of Dipole Moments SOV/56-35-6-28/44

The case in which ϵ and μ are different from 1 is finally discussed. The authors thank L. S. Bogdankevich, A. V. Gaponov, M. A. Miller and I. M. Frank for discussions. There are 12 references, 11 of which are Soviet.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva Akademii nauk SSSR

(Physics Institute imeni P. N. Lebedev of the Academy of Sciences, USSR) Gor'kovskiy gosudarstvennyy universitet

(Gor'kiy State University)

SUBMITTED: June 27, 1958

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SOV/56-35-6-37/44

24(3) AUTHOR:

Ginzburg, V. L.

TITLE:

On the Nonlinear Interaction of Radiowaves Propagated in a Plasma (O nelineynom vzaimodeystvii radiovoln,

rasprostranyayushchikhsya v plazme)

PERIODICAL:

Zhurnal eksperimental'noy i teoreticheskoy fiziki, 1958,

Vol 35, Nr 6, pp 1573-1575 (USSR)

ABSTRACT:

With the propagation of sufficiently strong radio waves in a plasma, especially in the terrestrial ionosphere, nonlinear phenomena (cross modulation, interaction, and also "self-action" of non-modulated radio waves) occur (Refs 1-3). As far as the author knows, only nonlinearity in dependence on the effective number $\nu_{\rm eff}$ of collisions having a field strength E_1 of a strong radio wave 1 has hitherto been investigated. In the most simple case of a non-modulated wave 1 (when it holds that $E_1 = E_0 \cos(\omega t - k r) = E_0 \cos \phi$), the condition $\omega^2 \gg \nu_{\rm eff}^2$ is satisfied and the influence exercised by the constant magnetic field is neglected. The nonlinear effect is in that case

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APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R000516730

connected in the most simple manner with the modification of

20v/56-35-6-37/44

On the Nonlinear Interaction of Radiowaves Propagated in a Plasma

electron velocity, and an expression is also written down for the corresponding variation of electric conductivity. These phenomena become a little more complicated if electron velocity distribution, modulation, etc. are taken into account, which is, however, not of essential importance for the linear effect under investigation. In an inhomogeneous isotropic plasma (and also in a homogeneous magnetically active plasma) electron concentration (unlike what is the case with an isotropic and homogeneous plasma) depends on radio wave field strength. An expression for the thereby caused variation of the radio wave 1 and of a plane wave in a magnetically active medium is written down. Variation of the electron concentration AN leads to a proportional variation of the dielectric constant E' of the magnetically active plasma. The nonlinear effect investigated is linear with respect to field 1, and the combination frequencies occurring are equal to $\omega^+ \pm \omega$. The effect investigated here is of the same kind as in the case of the scattering of transversal radio waves on plasma waves in an isotropic medium. The concrete influence exercised by this effect

Card 2/3

sov/56-35-6-37/44

On the Nonlinear Interaction of Radiowaves Propagated in a Plasma

upon the propagation of radio waves in the terrestrial

atmosphere and in the solar corona remains to be investigated.

There are 5 Soviet references.

ASSOCIATION:

Fizicheskiy institut im. P. N. Lebedeva Akademii nauk SSSR

(Physics Institute imeni P. N. Lebedev of the Academy of

Sciences USSR)

SUBMITTED:

August 23, 1958

Card 3/3

507/53-66-2-1/9

. AUTHORS:

Cetmantsev, G. G., Ginzburg, V. L., Shklovskiy, I. S.

TITLE:

Radioastronomical Investigations With the Aid of Artificial Earth Satellites (Radioastronomicheskiye issledovaniya s

pomoshch'yu iskusstvennykh sputnikov Zemli)

PERIODICAL:

Uspekhi fizicheskikh nauk, 1958, Vol 66, Nr 2, pp 157-161

(USSR)

ABSTRACT:

Artificial satellites are of great importance for opticalas well as for radio-astronomy; they may serve as receiving stations for near- and far ultraviolet-, X-ray- and far infrared radiation which, because of absorption in the atmosphere, does not reach the surface of the earth, as well as for the r.f.-range where absorption in the troposphere and refraction and absorption in the ionosphere act upon radiation. The authors first discuss absorption in the troposphere (especially in the $\lambda < 2$ cm range), connection with the effective temperature of the radiation source, solar and lunar radiation, the influence exercised by the ionosphere, and several problems of a general nature; dis-

Card 1/3

cussion is based upon scientific publications mentioned

Radioastronomical Investigations With the Aid of Artificial Earth Satellites

(Refs 1-8). The conditions for a receiving apparatus for the range 10 cm $< \lambda <$ 10 m are then discussed ($T_{\rm eff} = a\lambda^2$ °, intensity I, $= \frac{2kT_{\rm eff}}{2} \sim \lambda^{0.8}$; with $\lambda \sim 3$ m, $T_{\rm eff}$ is of the order of 10³ degrees, at 30 cm $< \lambda <$ 100 m $T_{\rm eff} \sim 10^6$ to 10⁷ degrees, I, \cong const; $\lambda > 100$ m: $T_{\rm eff} \approx 10^7$ degrees). The authors further discuss radio-receiving apparatus. For $\lambda > 100$ m very low limiting values of the noise factor(F $_{\rm n} \sim 2)$ are obtained for coincidence superheterodyne receiving sets. For large λ wire antennae of several 10 m length would be necessary; as this is impossible in a Sputnik, frame antennae with ferrite core are used, which can be of very small dimensions ($\ell \sim$ 10 cm, weight 300 g). The axis of the frame is parallel to the metal surface of the Sputnik. Because of a Sputnik's own rotary motion also the position of the frame is modified which causes fluctuations of the intensity of reception. It is therefore necessary to know the orientation of the frame at every instant. The antenna will not receive a radiation for which it holds that $\mathcal{E}(f,\mathbb{N})$ = 0 at the place of reception. If the magnetic terrestrial field is

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Radioastronomical Investigations With the Aid of Artificial Earth Satellites

neglected, it holds that

$$\varepsilon(f) = 1 - \frac{4\pi e^2 N}{m(2\pi f)^2} = 1 - 8.10^7 \frac{N}{f^2}$$

Here N is the electron concentration, f - the frequency of the radiation received. In interplanetary space $11 \sim 1$ to 5.10²,

 $\mathcal{E}(f) > 0$, $f > f_o = 9.10^4 - 2.10^5$ or $\lambda = c/f < \lambda = 1.5$ to 3 km. When measuring f_o it is possible to calculate N according

to the aforementioned formula. The influence exercised by the terrestrial field complicates investigation, but this influence is not very considerable for relatively fast Sputniks. There are 11 references, 4 of which are Soviet.

Card 3/3

GINZBURG VL

20-3-13/59

AUTHOR:

Ginzburg, V. L., Corresponding Hember AN USSR

TITLE:

The Critical Current for Superconducting Films (Kriticheskiy tok dlya sverkhprovodyashchikh plenok)

PERIODICAL:

Doklady AN SSSR, 1958, Vol. 118, Nr 3, pp. 464 - 467 (USSR)

ABSTRACT:

If films are used which are laid upon a cylindrical surface, the determination of the critical current obviously is more reliable than the determination of the critical field strength. Therefore the author here discusses the computation of the critical current somewhat more exactly than in one of his previous works (reference 1). The most interest deserve thin films with a thickness of 1 ~ 10-5 to 10-6 cm. In case of application of cylindrical supports with ~1 mm diameter such films can be supposed to be plane and the cylindrical configuration of the film must be considered only in case of the boundary conditions for the field. First the equations for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the determination of the function \(\psi \) and of the vector for the field is given. The complete current flows in

Card 1/2

20-3-13/59

The Critical Current for Superconducting Films

the direction of the cylinder axis and in the same direction an outside field can be applied as well. In the inside surface of the cylindrical film the field strength of the field, which is caused by the current, is equal to zero. For the field strength h in the film and for the corresponding potential a terms are written down. The critical field strength of the field induced by the current can be ascertained from an equation given here. The transition into the normal state has under certain conditions the character of a second order transition. There are 2 figures, and 4 references, all of which are Slavic.

ASSOCIATION:

Physical Institute imeni P. N. Lebedev AN USSR

(Fizioheskiy institut im. P. N. Lebedeva Akademii nauk SSSR)

SUBMITTED:

October 31, 1957

AVAILABLE:

Library of Congress

Card 2/2

AFRIKYAN, Levon Melkonovich; GINZBURG, V.L., red.; GARIBYAN, G.M., kand.fiz.-mat.nauk, red.; AZIZBEKYAN, L.A., tekhn.red.

[Works on theoretical physics] Raboty po teoreticheskoi fizike.
Pod red. V.L.Ginzburga i G.M.Garibiana. Erevan, Izd-vo Akad.
nauk Armianskoi SSR, 1959. 74 p. (MIRA 12:12)

1. Chlen-korrespondent AN SSSR (for Ginzburg).
(Physics)

"APPROVED FOR RELEASE: Thursday, July 27, 2000

CIA-RDP86-00513R00051673

GINSBURG, V.L.

"CERTAIN ASPECTS OF COSMIC RAY ORIGIN THEORY" V.L. Ginsburg

Certain aspects of the theory of cosmic ray origin are discussed in the light of the works which appeared since the Varenna conference.

report presented at the International Cosmic Ray Conference, Moscow 6-11 July 1959

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06330 SOV 141-21-1-2/19

AUTHORS: Gershman, B. N. and Ginzburg, V. L.

TITLE: On the Formation of Ionospheric Irregularities

PERIODICAL: Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika, 1959, Vol 2, Nr 1, pp 8-13 (USSR)

ABSTRACT: In the case of the lower layers of the icrosphere (in particular, the E-layer), there is no doubt that the formation of irregularities is due to the turbulization of gas currents and both the turbulization and the irregularities are produced by ionospheric winds (Ref 1). Therefore, the only controversial problem is the mechanism of the formation of irregularities in the F-layer and one is mainly concerned with irregularities responsible for the twinkling of radio stars and the spread of the F-echo. The present paper is mainly concerned with the discussion of the motion of the ionized component of the gas in the ionosphere. The quasihydrodynamic equations, given by Eqs (1)-(3), are employed. In these equations the subscripts e, i and m refer

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On the Formation of Ionospheric Irregularities

to electrons, ions and molecules, respectively. u are the velocities, $\rho_e = mN_e$, $\rho_i = MN_i$ and $\rho_m = MN_m$ are the N_e , N_i , N_m are the concentrations of densities, electrons, ions and molecules, -e , m are the charge and the mass of an electron, M is the mass of the ions and molecules (assumed equal, the charge of the ions is taken as equal to e), Ho is the intensity of the terrestrial magnetic field (the difference between the magnetic field and H_{O} is neglected), E is the intensity of the electric field, η_{e} , η_{i} and η_{m} are the viscosity coefficients, ν_{ei} , $\gamma_{\rm em}$ and $\gamma_{\rm im}$ are the numbers of collisions of electrons with ions and molecules, and ions with molecules, g is the acceleration due to gravity and p is the pressure. If the mean density p = 9e + qi is introduced and $\underline{\underline{u}}_p = (\rho_e \underline{\underline{u}}_e + \rho_i \underline{\underline{u}}_i)/(\rho_e + \rho_i)$ then Eqs (1) and (2) give Eq (4), where $N = N_e = \bar{N}_i$ and $N \leqslant N_m$. In this equation terms Card2/5 involving pressure, viscosity and the force of gravity are

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06330 SOV/141-2-1-2/19

On the Formation of Ionospheric Irregularities included and

 $j = eN(\underline{u}_i = \underline{u}_e)$, $\underline{u} = (\rho_m \underline{u}_m + \rho_p \underline{u}_p)/(\rho_p + \rho_m)$, $\nu_{em} \gg \nu_{im}$, $m\nu_{im} \gg m\nu_{em}$ and $m\nu_{im} \gg m\nu_{ei}$. The system of equations (1)-(2) also leads to Eq (5), in which the unimportant terms have been neglected and $\nu_e = \nu_{ei} + \nu_{em}$. It is clear from Eq (4) that in the absence of the field \underline{H}_0 and the current \underline{i} , the velocities \underline{u}_p and \underline{u} become equal during a time of the order of ν_{im} . The current \underline{i} for $\underline{H}_0 = 0$ is damped out during a time of the order of ν_{im} . If one confines one's attention to quasi-static processes then the time derivative in Eqs (4) and (5) may be neglected and one obtains Eqs (6), where σ_{ii} , σ_{i} and σ_{ii} are the conductivities parallel to μ_{i} , perpendicular

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SOV/141-2-1-2/19 On the Formation of Ionospheric Irregularities ${\rm H}_{\rm o}$ and the Hall conductivity respectively. are the parallel and perpendicular to \underline{H}_{0} components \underline{E}^{*} . Also, ω_{H} = eH_{O}/mc and Ω_{H} = eH_{O}/Mc . If the z-axis is chosen along the \underline{H}_0 and condition (7) is satisfied, then one obtains Eq (8), where the y-axis is chosen to be perpendicular to \underline{u} . Condition (7) is satisfied for altitudes $\geq 90-100$ km. If condition (9) is satisfied, then Eqs (8) assume the form given by the first three equations at the top of p 11. If condition (10) is satisfied then one obtains Eq (11). It is shown that an ionospheric wind can be set up in the F-layer only in the presence of an electric field \underline{E} which, in the first approximation, is independent of the velocity \underline{u} and is given by Eq (12). An analysis of the above theory leads to the conclusion that the formation and motion of ionization irregularities in the F-layer is not a hydrodynamic problem and should be considered with the

inclusion of the terrestrial magnetic field $\underline{\underline{H}}_{0}$, the electric field $\underline{\underline{E}}$ and the difference between the velocity Card4/5 of the gas as a whole, \underline{u} , and the velocity of the ionized

On the Formation of Ionospheric Irregularities 66330

SOV/141-2-1-2/19

component \underline{u}_p . If one neglects the velocity gradients, the quantity \underline{u}_p is determined by the two quantities \underline{u} and \underline{E} and usually to a good approximation by \underline{E} only (cf Eq (12)). In order to solve the ionospheric wind and the ionospheric irregularity problems, the character of distribution of the field \underline{E} and the velocity \underline{u} must be known on a scale comparable with the dimensions of the terrestrial globe. There are 12 references, of which 4 are Soviet and 8 are English.

ASSOCIATION: Issledovatel'skiy radiofizicheskiy institut pri Gor'kovskom universitete (Research Radio-Physical Institute of Gor'kiy University)

SUBMITTED: August 25, 1958.

Card 5/5

24(5) AUTHORS: Ginzburg, V. L., Eydman, V. Ya.

SOV/56-36-6-28/66

TITLE:

The Radiative Force For a Charge Moving

(O sile reaktsii izlucheniya pri dvizhenii in a Medium

zaryada v srede)

PERIODICAL:

Zhurnal eksperimental'noy i teoreticheskoy fiziki, 1959,

Vol 36, Nr 6, pp 1823-1833 (USSR)

ABSTRACT:

In the present paper the radiative force for a non-punctiform charge moving in a generally anisotropic and gyrotropic medium is investigated. The radiative force in a medium may play a considerable role when the particle moves in a magnetoactive plasma, in channels and slits in dielectrics and also in wave guides. At velocities larger than the phase velocity of light in the medium the radiative force, which changes the amplitude of the oscillations and which is related to the emission of anomalous Doppler frequencies, possesses a different sign than that of radiative friction due to the emission of normal Doppler frequencies. The total radiative force which is responsible for the change in the amplitude of the oscillations of a particle in an isotropic medium

Card 1/2

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The Radiative Force For a Charge Moving in a Medium

SOV/56-36-6-28/66

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corresponds to friction also in the case of super-light motion. However, this friction may be appreciably smaller than the radiative friction encountered at sublight velocities. In an anisotropic medium amplification of the oscillations may occur instead of friction. The decrease of radiative friction or the appearance of the amplification may be related to the peculiarities of the anomalous Doppler effect as revealed by a quantum mechanics amalysis and also to the instability of the super-light particle beams. The theoretical considerations are based upon the results obtained by a large number of previous papers (Ginzburg et al), and, in the course of the final discussion, the resulting conclusions are discussed. There are 15 Soviet references.

ASSOCIATION:

Radiofizicheskiy institut Gor'kovskogo gosudarstvennogo universiteta (Radiophysics Institute of Gor'kiy State University)

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SUBMITTED:

December 20, 1958

Card 2/2

24.6820

67524 SOV/141-2-3-1/26

AUTHORS:

Ginzburg, V.L. and Eydman, V.Ya.

TITLE:

On Some Peculiarities of Electromagnetic Waves Radiated

by Particles Moving Faster Than Light

PERIODICAL: Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika,

1959, Vol 2, Nr 3, pp 331 - 343 (USSR)

ABSTRACT:

The paper was presented at the Ministry of Higher Education Conference on Radio-electronics, Kiyev, 1959.

The classical treatment of this problem yields the

Vavilov-Cherenkov radiation Condition in:

 $\cos \Theta_0 = c/n(\omega)v$ (1)

is the angle between the particle velocity $\bar{\mathbf{v}}$

and the wave-vector \vec{k} of the Cherenkov wave, $n(\omega)$ is the refractive index at the frequency $\,\omega\,$, the medium being isotropic. In this paper quantum representations

are used because they are so fruitful of interesting results. The fundamental conclusion is that for particles

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moving faster than light the reaction force of the radiation, changing the amplitude of particle vibration.

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On Some Peculiarities of Electromagnetic Waves Radiated by Particles Moving Faster Than Light

is less compared with that for velocities less than light and, in an anisotropic medium, can even change sign. The force corresponds, therefore, not to "friction" but to an excitation of the vibrations. This effect is obviously directly connected to the instability of faster-thanlight particle beams. A point charge moving uniformly in an isotropic medium radiates energy, as a result of the Vavilov-Cherenkov effect, at a rate given by Eq (2). If the radiated frequency is ω_0 , then as a result of the Doppler effect, the apparent frequency at an angle Θ given by Eq (3). Within the so-called Cherenkov cone the Doppler effect is anomalous since w increases with Θ and, if n is constant, $\omega \to \infty$ when $\Theta \to \Theta_0$. In practice, the effect is of interest for particle beams passing through narrow slots or close to delaying systems or for beams in magneto-active plasma where the losses are low. From a quantum point of view, the kinematics of radiation are determined by the laws of conservation of

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APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R00051673(

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SOV/141-2-3-1/26
On Some Peculiarities of Electromagnetic Waves Radiated by Particles Moving Faster Than Light

The changes in energy and momentum energy and momentum. as a result of radiation are given in Eqs (4) and (5), respectively. A system which moves uniformly in vacuo can only radiate as a result of a change in its interval state (thus, for example, an electron cannot radiate in vacuo if moving uniformly). In the general case, when $n \neq 1$, the radiation condition, in quantum terms, is that given by Eq (6). The advantage of the latter representation is that it shows the normal Doppler effect to involve an energy transition from an upper to a lower level, while the anomalous effect requires the reverse transition. A system which has only two discrete energy levels can exhibit both kinds of Doppler effect. In systems with many energy levels the anomalous effect leads to the possibility of exciting transverse radiation. Two cases exist, corresponding to an increase and decrease, respectively, of the system energy. The calculation of the transition probabilities which determine how a system will behave may be carried out by classical means; quantum methods

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SOV/141-2-5-1/25

On Some Peculiarities of Electromagnetic Waves Radiated by Particles Moving Faster Than Light

offer no advantage. The absorption coefficient, in the "normal" process, is that given by Eq (9) while the anomalous value is Eq (10). The latter expression is useful where the production of microwaves is considered. In particular, the case of a magneto-active plasma medium is applicable to sporadic solar radiation. In an anisotropic medium the phase and group velocities of a wave need not have the same direction. Figure 2 shows the effect of the sign of $d\omega/dk_r$ on the generation of the Cherenkov radiation.

As a rule, the radiation forces are small compared with the retarding forces but may become significant when motion occurs in narrow channels or in plasma.

There are 2 figures and 26 references, 25 of which are Soviet and 1 Hungarian.

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On Some Peculiarities of Electromagnetic Waves Radiated by Particles Moving Faster than Light

ASSOCIATION: Issledovatel'skiy radiofizicheskiy institut

pri Gor'kovskom universitete (Radiophysics Research

Institute of Gor'kiy University)

SUBMITTED: February 25, 1959

Card 5/5

3(1)

AUTHORS:

Ginzburg, V.L., Zheleznyakov, V.V.

SOV/33-36-2-5/27

TITLE:

On the Propagation of Electromagnetic Waves in the Solar Corona Taking Into Account the Influence of the Magnetic Field

PERIODICAL:

Astronomicheskiy zhurnal, 1959, Vol 36, Nr 2, pp 233-246 (USSR)

ABSTRACT:

The present note has preparatory character. In a following article the authors intend to investigate the influence of the magnetic field of the corona on the sporadic solar radiation. In this connection the influence of the magnetic field on the propagation and emission of the electromagnetic waves of the corona is considered as a preparation. The authors compile well-known results of western and Soviet scientists and complete them in a form necessary for the following article. In particular they consider the emission from the corona caused by the interaction of normal waves and caused by their dispersion on the fluctuations of the electron density; conditions of emission are given. Furthermore the authors describe the propagation of the electromagnetic waves in the corona under the influence of a strong sunspot magnetic field. N. A. Mityakov is mentioned in the paper.

Cord 1/2

ASSOCIATION: Sci. Res. Inst. of Radiophysics, Gor'kiy Univ.

24 (5), 24 (8)

AUTHOR: Ginzburg, V. L.

SOV/56-36-6-46/66

TITLE:

Comparison of the Macroscopic Theory of Superconductivity With Experimental Data (O sravnenii makroskopicheskoy teorii sverkhprovodimosti s eksperimental nymi dannymi)

PERIODICAL:

Zhurnal eksperimental'noy i teoreticheskoy fiziki, 1959, Vol 36, Nr 6, pp 1930 - 1932 (USSR)

ABSTRACT:

L. P. Gor'kov (Ref 1) (see this periodical p 1918) showed that the macroscopic equations for superconductors set up by Landau and Ginzburg (Ref 2) can be deduced from the modern microscopic theory of superconductivity. In the equations deduced by Gor'kov the charge eff was put equal to the double electron charge corresponding to the Cooper pairs (eff is identical with the quantity denoted by Gor'kov as e*). The phenomenological constant we obeys the equation (1):

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into the massive metal at the given temperature T). The measured

Comparison of the Macroscopic Theory of Superconductivity With Experimental Data

SOV/56-36-6-46/66

penetration depth δ is equal to δ_L if T is equal to the cuitical temperature T_c . Within the range of critical temperature, accuracy differs, for Sn $\delta \approx \delta_L$ at Δ T = T_c -T \lesssim 0.1°, at Al $\delta \approx \delta_L$ if Δ T \lesssim (10⁻³)°. Further, a number of empirical relations is set up, viz. for δ , H_{cm} , κ , T_c and Δ and for Sn, Sn-In (2.3% In). The experimentally obtained values are compared with theoretical values. Thus, for Sn T_c = 3.73° and κ = 0.153 is obtained, according to Faber 0.15 and for the isotropic model according to reference 5, 10: κ = 0.149, so that the value κ = 0.15 - 0.16 may be considered to be correct (both in the macroand in the microscopical theory). If, for the surface energy, it holds that $\sigma_c = H_{cm} \Delta/\delta \pi$, one obtains with κ = 0.158 : Δ = 6.5 $\delta_L \approx$ 1.66.10⁻⁵. $\sqrt{T_c/(T_c - T)}$; experimentally, however, the values 2.5 (according to Sharvin, reference 8) and 1.38 (Faber, reference 9) are obtained instead of 1.66. All data hitherto

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Comparison of the Macroscopic Theory of Superconductivity With Experimental Data

SOV/56-36-6-46/66

mentioned refer to Sn. Analogous comparisons are drawn for Al. There are 14 references, 7 of which are Soviet.

Fizicheskiy institut im. P. N. Lebedeva Akademii nauk SSSR ASSOCIATION:

(Physics Institute imeni P. N. Lebedev of the Academy of

Sciences, USSR)

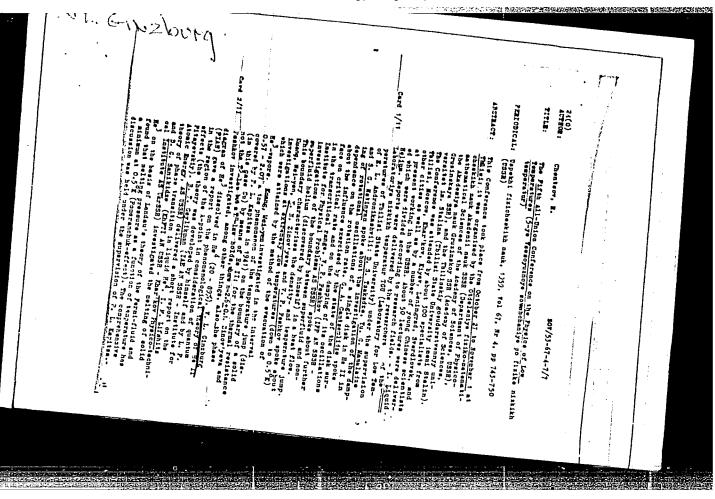
SUBMITTED:

February 19, 1959

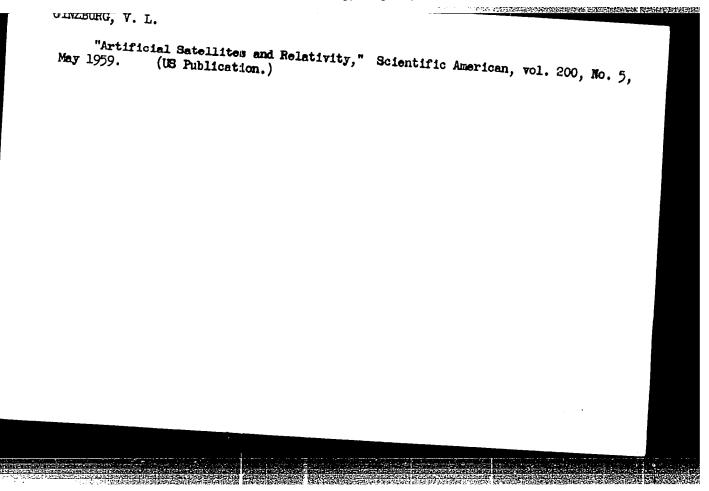
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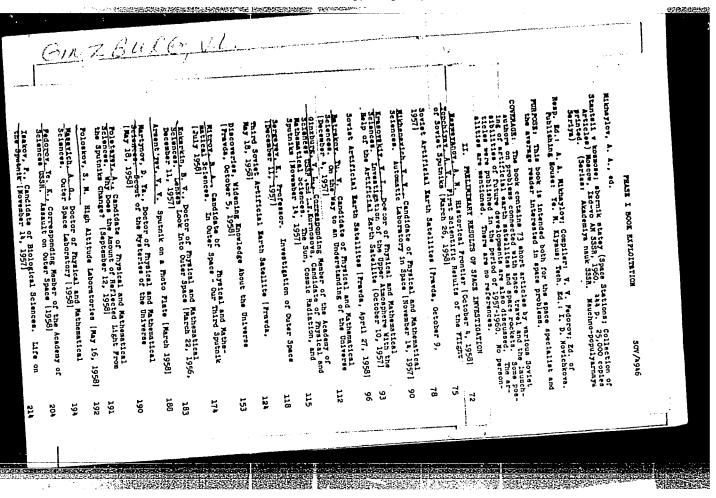


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	depersions that are not very near absolute acro. 1. 7. Shyke and L. K. Gurwich (FTI M SSER) probe about the sufface apergy on the boundary between the suprescenductive and norsal phases. 2. 7. Zubarev and U. A Terforvidov (antathobek) in stitut M SSER 2 Saheanics in Strikuter state (Fromhichek) to the thermotromates of the suprescenductive state (Fromhichek). 2. 7. Polanche (FLE) investigated the problem of collective accitations in a supreconductor. 2. 7. Shirkov (Ob-yedinanny institut yedernych isoletonan; - John: Institute of Noclear antatho speke about consideration of collective fasterns in semiconductors. The problem of consideration of electrons in semiconductors. The problem of consideration of the Coulomb interaction semiconductors. The problem of consideration of the Coulomb interaction semiconductors.	Oi - 4.28. A. A. Arikeger, L. P. Coriby and I. E. Balat- niker (IF) theoretically investigated the behavior of a giperconductor in the high-frequency field. T. L. Cinching and G. P. Enarter (FIRS) deals with the microscopical theory, and Ginburg discussed among other things the part played by fluctuations in phase transitions of the second kind. I. Lights (Energy about that it follows from the modern theory of superconductivity in consideration of the enten- tropy of seals that, in principle, the office enter ontucing is possible which are super-conductive only with- on listed range of the pressions of the file of the pre- balos the critical ones; D. T. Goglikum and T. E. Ersin (ILE) investigated the silvaton- and phonon thermal conductivi- ty of superconductors by means of the picconcepted there is	The Conference was aliended by about 300 specialists from Theirs, Serdowsk, and Theirs, Lyre, Lenharzi Serdowsk, and Obbar cities as well as by a number of young Chinese scientist at present working the two serd delivered to the serdowsky of the serious at the serious at the object on the field of which were strength to research fields. It appends on experiental invertigation of purposed on this field of which we were statement invertigations of purposed of the theory of the context the serious of the context of the serious	Temperatury ()-ye feesyutnoye moministative transporture () **Temperatury () **Temperatury () **Temperatury () **This Conference took place from October 27 to Mograber () **This Conference took place from October 27 to Mograber () **This Conference took place from October 27 to Mograber () **This Conference took place from October 27 to Mograber () **This Conference took place from October 27 to Mograber () **This Conference took place from October () **This Conference took place from October () **Entropy () **Entropy () **This Conference took place from October () **Entropy () **En	Chentery, R. The !!!!h !!! - Snion Conference on the Physics of Low	
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PHASE I BOOK EXPLOITATION

Rasprostraneniye elektromagnitnykh voln v plazme (Propagation of Electromagnetic Waves in Plasma) Moscow, Fizmatgiz, 1960. 552 p. Ginzburg, Vitaliy Lazarevich

Errata slip inserted. 8,000 copies printed.

PIRPOSE: This book is intended for scientific workers, aspirants, V.D.Kozlov; Tech. Ed.: K.F.Brudno. and university students enrolled in advanced courses in physics

COVERAGE: The book discusses the propagation of electromagnetic Waves of various frequencies in an isotropic as well as in a waves of various frequencies in an isotropic as well as in a magnetically active plasma, the propagation of electromagnetic waves of various types, radio waves, plasma waves, magnetohydrodynamic waves in plasma, and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and the behavior of a plasma in an analysis and analysis and an analysis a dynamic waves in plasma, and the behavior of a plasma in an The author electric field varying in time, but uniform in space.

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TOR RELEASE: Thursday, July 27, 2000

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also treats the case of a homogeneous and inhomogeneous medium, Propagation of Electromagnetic (Cont.) also treats the case of a nomogeneous and innomogeneous medium, and gives special attention to the propagation of radio waves in the analysis and in cosmic space. incomphere and in cosmic space. The author reports a number of basic results and conclusions from research in plasma physics. lonosphere and in cosmic space. The author reports a number of basic results and conclusions from research in plasma physics, placing emphasis on investigations and research in which he took part. Where constitute he has horrowed material for the placing emphasis on investigations and research in which he took part. Where possible, he has borrowed material for the present work from his Teoriya rasprostraneniya radiovoln to the propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the Propagation of Radio Waves in the longsfere (Theory of the Propagation of Radio Waves in the Propagation of Radio Wave present work from his reorlya rasprostranentya radiovoli versionosfere (Theory of the Propagation of Radio Waves in the Jonosphere) (Gostavhizdat Joho) and from the second hard Tonosphere) ionosfere (Theory of the Propagation of Radio Waves in the Ionosphere) (Gostekhizdat, 1949), and from the second part of Radio (Gostekhizdat, radiovoln (Propagation of With the book, Rasprostraneniye radiovoln (Propagation with use as the book, Gostekhizdat, 1953), written in collaboration use as To facilitate its use as Waves) (Gostekhizdat, L. Feynberg. To facilitate in various Ya. L. Al'pert and Ye. L. Feynberg. been repeated in various reference book, certain formulas have been repeated. reference book, certain formulas have been repeated in various reference book, certain formulas have been referenced to sections of the book reference book, certain formulas have been repeated in various to reference to the book. The bibliography includes reference, the sections of the book. The bibliography work, for example, the works on problems not discussed in this work, statistical propagation of radio waves in the presence of statistica propagation or radio waves in the presence of statistical B.N. Gersh-irregularities. The author thanks Ye.A. Benediktov, M. M. M. Lyakov Tregularities. The author thanks Ye.A.Benediktov, B.N.Gershman, A.V.Gurevich, N.G.Denisov, V.V.Zheleznyakov, N.A.Mityakov, man, A.V.Gurevich, N.G.Denisov, V.V.Zheleznyakov, N.A.Mityakov,

s/035/61/000/012/013/043 A001/A101

AUTPHOR:

Ginzburg, V.L.

TITLE:

Some questions in the theory of cosmic ray origin

PERIODICAL:

Referativnyy zhurnal. Astronomiya i Geodeziya, no. 12, 1961, 79. abstract 12A329 ("Tr. Mezhdunar, konferentsii po kosmich, Ducham,

1959, V. 3", Moscow, AN SSSR, 1960, 200 - 208)

The mean life time of cosmic ray nuclei in the Galaxy is estimated. The mean life time of cosmic ray nuclei in the Galaxy is estimate for protons it amounts to 3.8×10^9 years, for nuclei of the M group (C. N. 0) - 3.6×10^9 and for Fe = 1.4 × 100 years. Since the existence time of the Galaxy is about 1010 years, the flux of Fe nuclei, for instance, should have been attenuated about 10- years, the risk of remucies, for instance, should have been assented and during this time by a factor of 1025. Therefore, the hypothesis that cosmic rays were generated in the early stage of galactic evolution should be rejected. The necessary summary power of cosmic ray sources (~10%) - 10 erg/sec) can be necessary summary power of cosmic ray sources (Figure 10 to energy into energy of cosmic rays is quite plausible. It a preferential acceleration of heavy nuclei of the N group $(Z \geqslant 10)$ takes place in this process, then

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APPROVED FOR RELEASE: Thursday, July 27, 2000

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S/035/61/000/012/013/043 AU01/A101

Some questions in the theory of cosmic ray origin

the observed proton flux is a secondary one, due to disintegration of heavy nuclei. The magnitude of the ratio of proton flux to the flux of heavy nuclei, which is equal to 15 - 20, agrees with this assumption. Diffusion of cosmic rays generated near the galactic plane to the halo boundaries restricts their life time in the Galaxy. Therefore it is presumed that magnetic fields in the Galaxy orme in one dataxy. Incretore to is presumed that magnetic fields in the dataxy are in such a disordered state ("knot" of force lines) that ~ 1% of cosmic rays, are in such a disordered state ("knot" of force lines) that ~ 1% of cosmic rays, are in such a disordered state ("knot" of the halo. In this case the life life, 10° particles/sec can go out of the halo. In this case the life time of protons relative to nuclear absorption is comparable to their life time relative to the leakage from the Galaxy. If there are galacto fields of 10 garas, heavy nuclei possessing energies up to 1019 ev will seen leave the mi considerably slower than vanish in nuclear collisions. Taking into account heir leaving the system due up drifting in a non-homogeneous magnetic (all these no alter the conclusions drawn. The problem of anisotropy segree of costs that items is discussed, which is due to ordering of the magne in freed in the garactic spiral. Radio astronomical data, as well as data on subscript degree of ocean rays near the Earth, show that effect of aparal arms on the motion of scenar that is insignificant. The problem of the origin of cosmic ray electron communist is briefly discussed. If the adiabatic invariant Sin20 H is preserved in the raid the motion of cosmic rays will proceed with low values of Tite. The ree field

Card 2/3

s/2831/60/000/002/0013/0018

ACCESSION NA: AT3012749

AUTHORS: Gershman, B. W.; Ginzburg, V. L.

TITLE: Formation of ionospheric inhomogeneities

SOURCE: AN SSSR. Mezhduvedomst. komit. po prov. mezhdunarodn. geofizich. goda. 5 razdel program. MGG: Ionosfera. Sb. statey, no. 2, 1960, 13-18

TOPIC TAGS: ionosphere, ionospheric inhomogeneities, F layer, E layer, ionization wind, plasma velocity, ionization wind velocity, altitude variation of ionosphere

ABSTRACT: Some of the hypotheses recently advanced to explain the mechanism whereby inhomogeneities are produced in the F layer are discussed, with emphasis on the inhomogeneities that cause flicker of radio stars and diffuse reflection from the F layer. The authors analyze the conditions which determine the motion of an ionized gas

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AT3012749 ACCESSION NR:

component in the ionosphere, since a clarification of this question can serve as a premise for the construction of a theory of "ionization winds" and formation of inhomogeneities in the F layer. It is concluded from various estimates of the possible plasma and wind velocities and the resultant variations of the electric and magnetic fields in the ionosphere that the formation and motion of inhomogeneities of ionization in the F layer is not a hydrodynamic problem but must be solved with allowance for the earth's magnetic field, the electric, field, and the fact that the gas as a whole does not move with the same speed as its ionized component. nature of the ionospheric winds and ionospheric inhomogeneities must be ascertained by clarifying the character of the distribution of the electric field and the velocity on scales that are comparable, with the dimensions of the earth's sphere. The particular partial problems still to be considered are: the transport of the electric field in the F layer from the diurnal region and the related circulation in the F layer, the passage of different types of low fre-

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APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R00051673(

ACCESSION NR: AT3012749

quency waves from the E layer into the F layer (capable of becoming propagated in a weakly ionized gas), the influence of the earth's magnetic field, and the inhomogeneity of the atmosphere with altitude. Orig. art. has: 12 formulas.

ASSOCIATION: None

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ENCL: 00

SUB CODE: AS, AI

NO REF SOV: 005

OTHER: 006

Card 3/3

811062 S/181/60/002/009/003/036 B004/B056

Ginzburg, V. L.

AUTHOR:

TITLE:

Some Remarks on Phase Transitions of Second Kind and the Microscopic Theory of Seignettoelectrics 1

PERIODICAL:

Fizika tverdogo tela, 1960, Vol. 2, No. 9, pp. 2031-2043

TEXT: The author discusses the problem as to whether transitions of second kind have the same character in superconductors, liquid helium, ferromagnetics, and seignettoelectrics, and differ from one another only by the value of one parameter. Corresponding to the theory of phase transitions developed by L. D. Landau (Ref. 2), the author writes down the series: $\Phi = \Phi_0 + \alpha \eta^2 + (\beta/2)\eta^4 + (\gamma/6)\eta^6 + \delta(\operatorname{grad} \eta)^2$ (1) for the thermodynamic potential Φ , where Φ_0 , α , β , γ ; δ , are functions of temperature and pressure, and $\delta(\operatorname{grad}\eta)^2$ is described as correlation energy. Because of . thermal motion, η fluctuates round a mean value η_0 . The series (1) in first approximation gives correct results as long as $(\Delta \eta)^2 \sim (\Delta \eta)^2_0$ and $(\overline{\Delta \eta})_{\mathrm{T}}^2$, respectively, are small as compared to $\eta_0^2 \equiv (\overline{\eta})^2$. The condition Card 1/3

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Some Remarks on Phase Transitions of Second S/181/60/002/009/003/036
Kind and the Microscopic Theory of Seignetto- B004/B056

electrics $\Delta T/\theta \gg \kappa^2/(\delta/\alpha_0^4 \theta)^3 \ (\Delta c)^2 = \int_T (5b) \ is \ written \ down \ (\kappa = Boltzmann \ constant, \ \theta = temperature of the transition point). \ \Delta T = \theta - T, \ T \simeq \theta,$ stant, \ \theta = temperature of the transition thus depends only on \(\int_T \cdot For \) \(\Delta c = (\alpha_0^1)^2 \theta_0^{-1} \). The character of transition thus depends only on \(\int_T \cdot For \) \(\Delta c = (\alpha_0^1)^2 \theta_0^{-1} \). The character of transition thus depends only on \(\int_T \cdot For \) \(\Delta c = (\alpha_0^1)^2 \theta_0^{-1} \). The character of transition thus depends only on \(\int_T \cdot For \) \(\Delta c = (\alpha_0^1)^2 \theta_0^{-1} \). Therefore, no anomalies of specific heat are observed in superconductors. Likewise, \(\int_T \) is anomalously low for ferro-observed in superconductors. Likewise, \(\int_T \) is anomalously low for ferro-observed in superconductors. Likewise, \(\int_T \) is anomalously low for ferro-observed in superconductors, above \(\frac{1}{11} \) BaTiO_3, it is found that the magnetics. For seignettoelectrics, above \(\frac{1}{11} \) BaTiO_3, it is found that the parameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, whereas the fluctuation is reparameter 12 = \(\delta/\alpha_0^1 \) is anomalously great, wher

"APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R00051673

S/181/60/002/609/003/036 B004/B056 Some Remarks on Phase Transitions of Second Kind and the Microscopic Theory of Seignetto-

electrics

ASSOCIATION:

Fizicheskiy institut im. P. N. Lebedeva, Moskva (Institute of Physics imeni P. N. Lebedev, Moscow)

SUBMITTED:

February 10, 1960

Card 3/3

80876

3,1700

s/141/60/003/02/023/025 E032/E314

AUTHOR: TITLE:

On the Possibility of a Determination of the Magnetic Field in the Outer Solar Corona by Examining the Polarized Radiation due to Discrete Sources Transmitted Through it

Izvestiya vysshikh uchebnykh zavedeniy, Radiofizika, 1960, Vol 3, Nr 2, pp 341 - 342 (USSR) PERIODICAL:

ABSTRACT: The presence of an ordered magnetic field in the solar corona can lead to a rotation of the plane of radio emission passing through the corona. The radio emission of the Crab nebula Tis the radiation in question. It passes through the corona during June and has a polarization of approximately 7% at 3 cm. The polarization is characterised by a position angle of W = 148-149. For 10 cm waves the polarization is $3 \pm 0.5\%$ and $U = 142 \pm 5$. There is evidence that at long wavelengths the polarization is smaller. In the corona (in the plane of the solar equator) the electron concentration is (Ref 7) N \sim 7 x 10 cm at $\eta = 5$, N \sim 10 cm $\eta = 10$ and N \sim 2.5 x 10^3 cm at $\eta = 20$. It

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On the Possibility of a Determination of the Magnetic Field in the Outer Solar Corona by Examining the Polarized Radiation due to Discrete Sources Transmitted Through it

follows that for $\eta = 5$ to 20, $\omega_0^2 = 4\pi r e^2 \text{N/m} = 3.18 \times 10^9$, $N \sim 2 \times 10^{14} - 8 \times 10^{12}$ and $\omega_H = \text{eH/mc} = 1.76 \times 10^7 \text{ H} \sim 10^5 - 10^5 \text{ sec}^{-1}$. Moreover, the

frequency of the radio emission $\omega = 2\pi c/\lambda \sim 2 \times 10^{10}$

 $\rm sec^{-1}$ at $\lambda \cong 10$ cm. Under these conditions the propagation of radio waves may be looked upon as quasilongitudinal for practically all angles α between the magnetic field \underline{H} and the direction of the wave normal. The difference between the refractive indices no for normal circularly polarized waves is $\Delta n = \omega_{H}^{2} \cos \alpha/\omega^{3} = 5.6 \times 10^{16} \text{ HN } \cos \alpha/\omega^{3} .$

rotation of the plane of polarization after passage through the plasma layer is given by Eq (1), where the integration is carried out along the ray which in the present case can be considered to be rectilinear. In order to estimate this

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On the Possibility of a Determination of the Magnetic Field in the Outer Solar Corona by Examining the Polarized Radiation due to Discrete Sources Transmitted Through it

effect it is assumed that:

$$\Delta \Psi \sim \frac{10^6 \text{ HNL cos } \alpha}{\omega^2}$$

where L is a certain effective path length. With $\eta = 5$, $H \sim 10^{-2}$ Oe, $N \sim 10^5$ cm⁻³, $\cos \alpha \sim 1$ and $_{L} \sim _{\eta}R_{0} \sim 3 \times 10^{11}$ cm the rotation of the plane of polarization is $\Delta \Psi = 60^{\circ}$. When $\eta \approx 10$, $H \sim 10^{-3}$ $N \sim cm^{-3}$, $\cos \alpha \sim 1$ and $L \sim 10R_Q$, the rotation is The relatively strong dependence of the rotation on ω should serve as a useful additional rotation. There are 9 references, 5 of which are Soviet effect. and 4 English.

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"APPROVED FOR RELEASE: Thursday, July 27, 2000 CIA-RDP86-00513R00051673

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On the Possibility of a Determination of the Magnetic Field in the Outer Solar Corona by Examining the Polarized Radiation due to Discrete Sources Transmitted Through it

: Nauchno-issledovatel'skiy radiofizicheskiy institut pri Gor'kovskom universitete (Scientific-research ASSOCIATION: Radiophysics Institute of Gor'kiy University)

SUBMITTED: March 25, 1960

Card 4/4

Achievements and objectives of the industry of rubber products for engineering uses in the German Democratic Hepublic. Kauch. (MIRA 13:11)
i rez. 19 no. 11:29-32 N '60.

1. TSentral'naya nauchno-issledovatel'skaya loboratoriya rezinovogo zavoda "El'be", Germanskaya Demokraticheskaya Respublika.

(Germany, East-Rubber goods)