

ACCESSION NO: AP3001240

S/0033/63/040/003/0466/0476

AUTHOR: Ginzburg, V.I.; Syrovatskiy, S.I.

69
67

TITLE: On cosmic rays in metagalactic space

SOURCE: Astronomicheskiy zhurnal, v. 40, no. 3, 1963, 466-476

TOPIC TAGS: cosmic rays, metagalactic cosmic rays, galactic cosmic rays, electron component cosmic rays, metagalactic radio emission, metagalactic gamma emission, cosmology

ABSTRACT: This theoretical discussion attempts an assessment of the possible energy density of cosmic rays in metagalactic (MG) space within the framework of evolutionary cosmology. In all probability, the energy density of MG cosmic rays is substantially smaller than that of galactic (G) cosmic rays; the latter approximates 10-sup-minus-12 erg/cc. The best possible estimate for the energy density of MG cosmic rays, at this time, approximates 10-sup-minus-15 to 10-sup-minus-16 erg/cc. The energy density of MG thermal radiation is estimated at 10-sup-minus-15 erg/cc or 10-sup-minus-3 eV/cc. A discussion is also given of models in which the MG cosmic-ray energy density is comparable with that of G cosmic rays only within the limits of a local supergalaxy or a local group of

L 1191-63
ACCESSION NR: AP3001240

galaxies. Such models may provide an explanation for the measured cosmic-ray intensities in the Galaxy only if a large number of far-reaching supplementary assumptions are made, all of which appear to have a low probability. By contrast, the galactic theory of the origin of the major portion of the cosmic rays observed on Earth does not encounter any serious objections. The paper discusses also a number of problems of the electron component of cosmic rays and radio emission and gamma emission of MG origin. There are 14 numbered equations.

ASSOCIATION: Fizicheskiy in-t (Physics Institute nauk SSSR, imeni P.N. Lebedevol, Akademi Academy of Sciences, SSSR)

SUBMITTED: 30 Oct 62

DATE ACQD: 01 Jul 63

ENCL: 00

SUB CODE: AS, FH

NO REF SOV: 008

OTHER: 009

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Card 2/2

L 13837-63
EWI(1)/BDS/EGG(b)-2 AFTC/ASD/ESD-3 GG/K

ACCESSION NR: AP3003143
8/0056/63/044/006/2104/2106

AUTHOR: Ginzburg, V. L.

59
58

TITLE: Account of the pressure in the theory of second-order phase transitions
(with application to the case of superconductivity)

SOURCE: Zhurnal eksper. i teor. fiziki, v. 44, no. 6, 1963, 2104-2106

TOPIC TAGS: superconductivity, second-order phase transition, pressure dependence, BCS microscopic theory

ABSTRACT: The effect of pressure on second-order phase transitions is considered with application to superconductors. The reasons for the study are that the theory is nearly exact in the region of superconducting transition at nonzero temperature, the microscopic Bardeen-Cooper-Schrieffer theory yields in the case of superconductivity additional information, and the pressure dependence of the various quantities can be easily determined by experiment in the case of superconductors. The characteristic parameter used in the theory is the energy-spectrum gap. Particular attention is paid to the dependence of the critical temperature on the critical pressure. Comparison of the method

L 13837-63

ACCESSION NR: AP3003145

employed with the phenomenological approach is made. Orig. art. has: 6
formulas.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva Akademii nauk SSSR
(Physics Institute, Academy of Sciences, SSSR)

SUBMITTED: 26Jun63

DATE ACQ: 23Jul63

ENCL: 00

SUB CODE: 00

NO REF SQ: 011

OTHER: 003

Card 2/2

Card 2/2

GINZBURG, V.L.; SYROVATSKIY, S.I.

Gamma rays and magnetic bremsstrahlung X rays of galactic and
metagalactic origin. Zhur. eksp. i teor. fiz. 45 no.2:353-363
Ag '63. (MIRA 16:9)

1. Fizicheskiy institut imeni P.N.Lebedeva AN SSSR.
(Cosmic rays)

BULAYEVSKIY, L.N.; GINZBURG, V.L.

Temperature dependence of the shape of the domain wall in
ferromagnetics and ferroelectrics. Zhur. eksp. i teor. fiz. 45
no. 3:772-779 S '63. (MIRA 16:10)

1. Fizicheskiy institut imeni P.N. Lebedeva AN SSSR.
(Domain structure)

GJNZBURG, V.L.

What is proved by measurements of gravitational frequency
shifting? Usp. fiz. nauk 81 no.4:739-743 D '63.

(MIRA 17:1)

AMBARTSUMYAN, V.A., akademik; GINZBURG, V.L.; ZEL'DOVICH, Ya.B.,
akademik; FONTEKORVO, B.E.; MORGBINSKIY, Ya.A., doktor
matem. nauk, prof.; FOK, V.A., akademik. CHERNOV,
A.G.; FAYNBOYM, I.B., red.

[Birth and evolution of the galaxies and stars; the third
discussion] Rozhdenie i evoliutsiia galaktik i zvezd; be-
seda tret'ia. [by] V.A.Ambartsumian i dr. Moskva, Izd-vo
"Znanie," 1964. 27 p. (Novoe v zhizni, nauke, tekhnike.
Seria IX: Fizika, matematika, astronomiia, no.12)

(MIRA 17:6)

1. Chlen-korrespondent AN SSSR (for Ginzburg, Fontekorvo).

GINZBURG, V.I.

Some results of radio astronomical research. Vest. AN SSSR 34
no. 2:17-21 F '64. (MIRA 17:5)

1. Chlen-korrespondent AN SSSR.

ACCESSION NR: AP4034067

S/0126/64/017/004/0631/0633

AUTHORS: Bulayevskiy, L. N.; Ginzburg, V. L.

TITLE: Possibility of the existence of surface ferromagnetism

SOURCE: Fizika metallov i metallovedeniye, v. 17, no. 4, 1964, 631-633

TOPIC TAGS: ferromagnetism, surface property, rare earth, magnetic moment

ABSTRACT: The possible existence of antiferromagnetic or paramagnetic crystals whose surface layer is ferromagnetic is discussed; such would be the case if there were uncompensated magnetic moments at the surface. The particular case of helical structure (structure type of rare-earth metals) is considered, with the axis of the helix assumed perpendicular to the surface and the spins directed along the surface. The angle between the directions of the average values of spins in adjacent layers is θ (neglecting surface effects). Considering only the interaction between nearest and next nearest neighbors, the exchange energy is

$$E = -N_s S^2 \sum_{k=1}^{n-1} (J_1 \cos \theta_k + J_2 \cos (\theta_k + \theta_{k+1}))$$

where θ_k is the angle between the spins of the k layer and those of the $k + 1$

ACCESSION NR: AP4034067

layer, N_0 is the number of unit cells in a layer perpendicular to the helix axis, S is the spin per cell, $I_1 > 0$ and $I_2 < 0$ are the parameters of the exchange interaction, and the thickness of the sheet is na , a being the lattice constant in the direction of the helix axis. Minimizing E over all θ_k results in a system of $n-1$ coupled homogeneous equations, the solution of which is sought in the form

$$\theta_k = \theta + \alpha_k, \text{ where } \theta = \arccos\left(-\frac{I_2}{4I_1}\right)$$

Assuming $\alpha_k \ll 1$, a finite-difference equation for α_k is written from which

$$\alpha_k = \alpha_1 e^{-1(k-1)} \text{ for } 1 < k < n/2;$$
$$\alpha_k = \alpha_1 e^{1(k+1-n)} \text{ for } n-1 > k > n/2;$$

where $\text{ch } \lambda = \frac{1}{2\cos^2 \theta - 1}$ and $\alpha_1 = \frac{\sin \theta}{\sin \theta + \cos \theta}$. Thus, for sufficiently small θ , the surface affects strongly distort the helix near the surface, and the distortion decreases exponentially with depth. A rough calculation shows that the uncompensated magnetic moment per unit surface area can reach $3 \cdot 10^{-4}$ erg/gauss for $\theta \sim 0.1$.
Orig. art. has: 18 equations.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva AN SSSR (Institute of Physics

ACCESSION NR: AP4034067

AN SSSR)

SUBMITTED: 23Nov63

ENCL: 00

SUB CODE: SS, MM

NO REF SOV: 001

OTHER: 001

MEMORANDUM FOR THE DIRECTOR, CIA.

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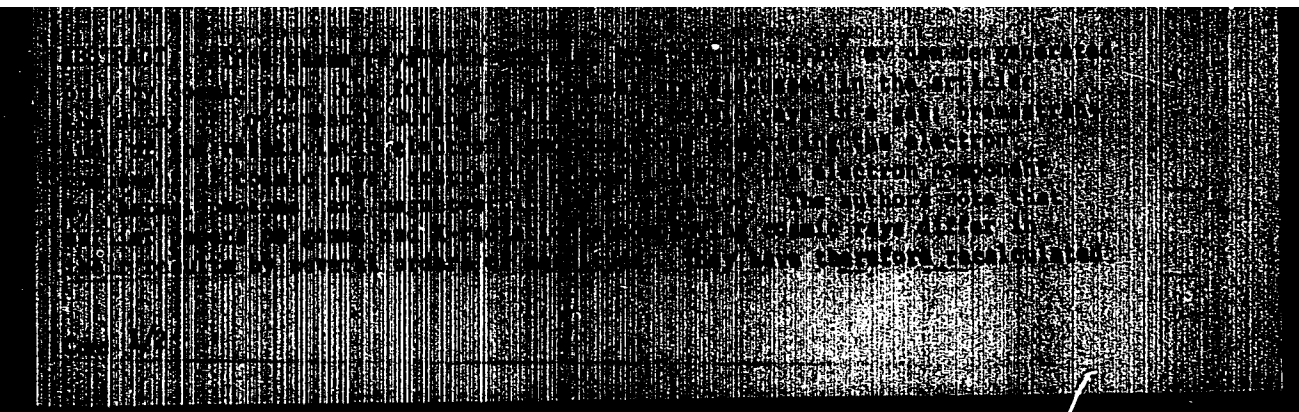
BARBER, R. L. 10/10/54

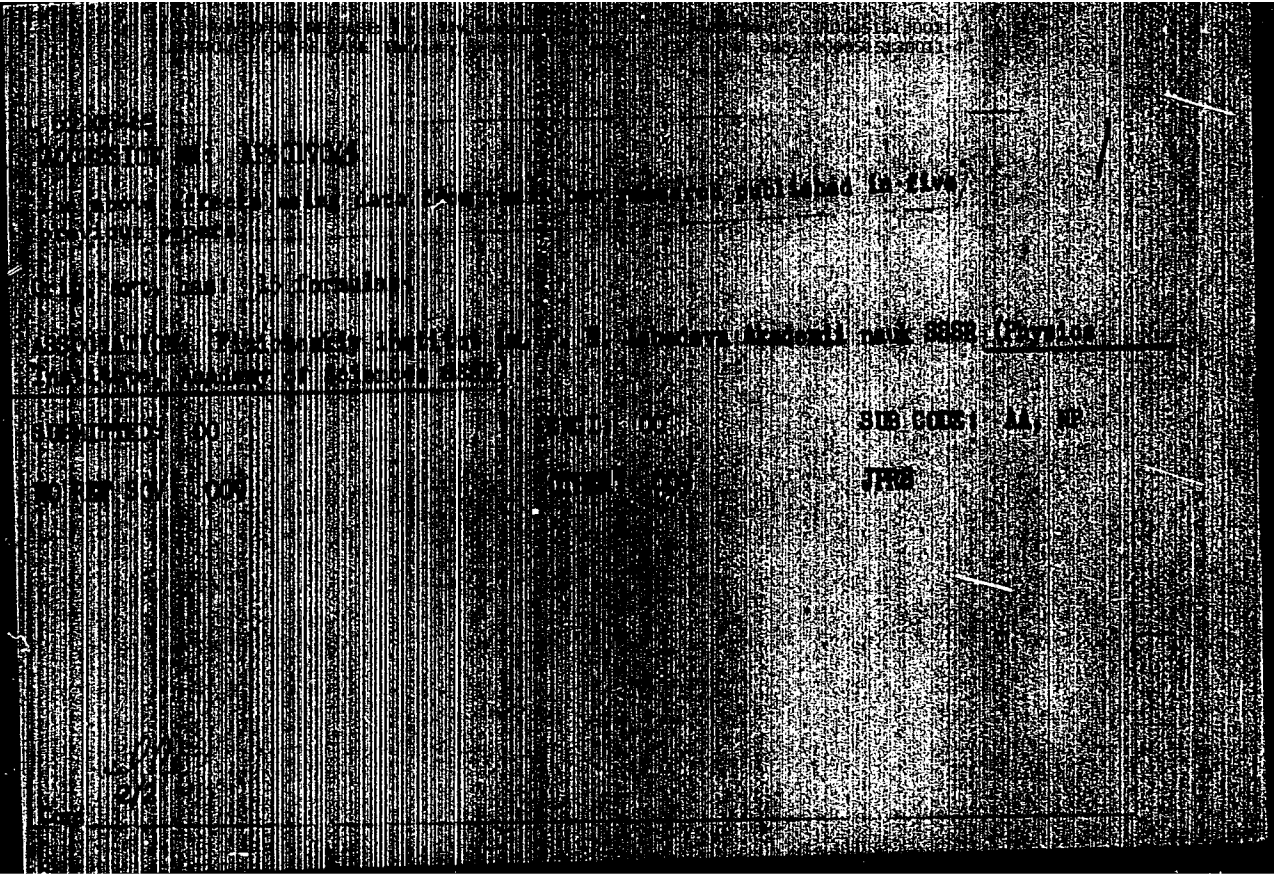
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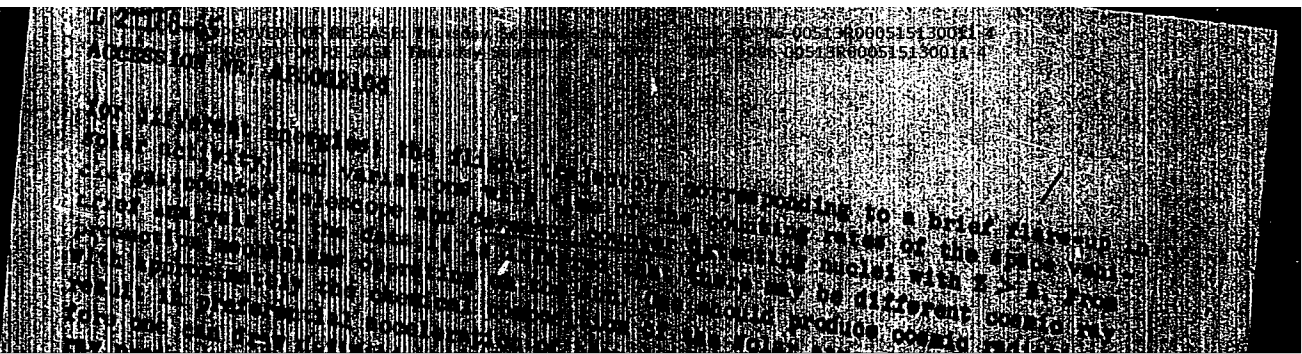


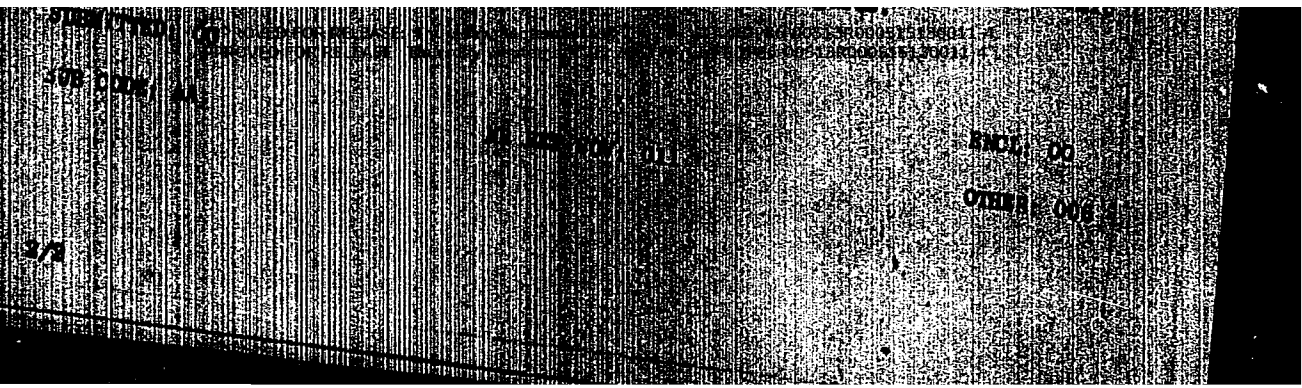
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... cosmic rays in the region of
TOPIC ... space vehicles, and comparative
... in the region of
... another table gives the values of
... nuclei referred to the total
... on the 1/3 ratio are reasonably
... for the percentages. Figures give
... nuclei with Z 2-5 and nuclei with Z 2-13





ACCESSION NR: AP4040841

S/073/64/041/003/0430/0445

AUTHOR: Ginzburg, V. L.

TITLE: Secondary electron component of cosmic rays and the spectrum of general galactic radio emission

SOURCE: Astronomicheskly zhurnal, v. 41, no. 3, 1964, 430-445

TOPIC TAGS: astrophysics, cosmic ray electron component, cosmic ray, galactic radio emission, radio emission, Galaxy, interstellar space, galactic halo, supernova, interstellar gas, synchrotron radiation

ABSTRACT: The authors computed the energy spectrum and spatial distribution of secondary electrons forming at the time of nuclear interactions of cosmic rays in the interstellar gas. The motion of electrons in interstellar space is considered as isotropic diffusion. Synchrotron, Compton and ionization losses of electron energy are taken into account. The authors also computed the intensity of the synchrotron radiation of secondary electrons. This intensity is appreciably less (by at least two orders of magnitude for wavelengths $\lambda \approx 10$ m) than the observed value. It therefore follows that the general galactic radio emission is generated by primary electrons which enter interstellar space from certain sources (in a stationary model, from the envelopes of supernovae). There is a discussion of

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ACCESSION NR: AP4012551

S/0056/64/046/001/0243/0253

AUTHORS: Ginzburg, V. L.; Meyman, N. N.

TITLE: On the dispersion relations for the refraction and absorption indices

SOURCE: Zhurnal eksper. i teoret. fiz., v. 46, no. 1, 1964, 243-253

TOPIC TAGS: dispersion relations, anisotropic medium, refractive index, absorption index, electrodynamics, macroscopic electrodynamics, dielectric tensor, plasma, magnetoactive plasma

ABSTRACT: This work is mainly a continuation of the work of one of the authors (V. L. Ginzburg) to which many references are given, particularly to two articles by V. M. Agranovich and V. L. Ginzburg (UFN, v. 76, 643, 1962 and v. 77, 663, 1962). Dispersion relations for the complex index of refraction in electrodynamics are considered. It is pointed out that the transition from dispersion

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ACCESSION NR: AP4012551

relations for the dielectric tensor to dispersion relations for the index of refraction is not in general as trivial a matter as it is for an isotropic medium. Analyticity questions for appropriate functions in anisotropic media, for example, magnetoactive plasmas, are discussed, and conditions are given under which the dispersion relations for the index of refraction are expected to hold even for such media. It is also observed that in optical problems spatial dispersion, if present, is usually sufficiently weak so that dispersion relations for those indices of refraction that refer to the behavior of normal waves (i.e., waves present also when there is no spatial dispersion) remain valid at least within a certain domain of complex values of the wave vector. Orig. art. has: 22 equations.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva AN SSSR
(Physics Institute, AN SSSR); Institut teoreticheskoy i eksperimental'noy fiziki (Institute of Theoretical and Experimental Physics)

Card 2/2

ACCESSION NR: AP4012571

S/0056/64/046/001/0397/0398

AUTHORS: Ginzburg, V. L.; Kirzhnits, D. A.

TITLE: Superconductivity of electrons in surface levels

SOURCE: Zhurnal eksper. i teoret. fiz., v. 46, no. 1, 1964, 397-398

TOPIC TAGS: superconductivity, surface superconductivity, surface electron, surface electron state, surface electron level, Cooper effect, Cooper pairs, metal superconductivity, dielectric superconductivity

ABSTRACT: The question is raised whether the Cooper phenomenon (Phys. Rev., v. 104, 1957, 1189) is possible for surface electrons (electrons at surface levels) which, as pointed out by I. Ye. Tamm (Phys. Zs. Sowjetunion, v. 1, 1932, 733), can be localized on crystal surfaces. It is pointed out that an interaction Hamiltonian can be formulated also for the two-dimensional case, that a negative value for the interaction constant cannot be excluded, and that the electrons then pass into a superconducting state. The possible existence of surface superconductivity is also considered for dielectrics. The effect of a magnetic field on surface conductivity is qualitatively considered. Orig. art. has: 5 formulas.

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ACCESSION NR: AP4012571

ASSOCIATION: Fizicheskij institut im. P. N. Lebedeva AN SSSR (Physics
Institute, AN SSSR)

SUBMITTED: 26Nov63

DATE ACQ: 26Feb64

ENCL: 00

SUB CODE: PH

NO REF SOV: 006

OTHER: 002

ACCESSION NR: AP4037602

S/0056/64/046/005/1865/1879

AUTHOR: Ginzburg, V. L.; Syrovatskiy, S. I.

TITLE: Gamma and x-ray radiation associated with galactic and metagalactic cosmic rays

SOURCE: Zh. eksper. i teor. fiz., v. 46, no. 5, 1964, 1865-1879

TOPIC TAGS: gamma ray, x-ray, cosmic ray, galactic cosmic ray, relativistic electron, interstellar gas, intergalactic gas

ABSTRACT: The intensity of γ -rays produced as a result of various interaction processes between cosmic rays (including the electron component) and the interstellar and intergalactic gas, or between cosmic rays and thermal radiation is calculated. It is shown that the main contribution to the γ -ray intensity is from the scattering of relativistic electrons by thermal photons, although γ -rays from the decay of π^0 -mesons and electron bremsstrahlung may also be important. Comparison of calculations with experimental data indicates

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ACCESSION NR: AP4037602

that the intensity of the electron component of cosmic rays in the Metagalaxy is at least one-and-a-half or two orders smaller than that in the Galaxy. The region of ultrahigh-energy electrons which could yield synchrotron x-rays and produce atmospheric showers with an anomalously small number of mesons is discussed. Orig. art. has: 49 formulas, 1 figure, and 2 tables.

ASSOCIATION: FIAN

SUBMITTED: 06Dec63

DATE ACQ: 09Jun64

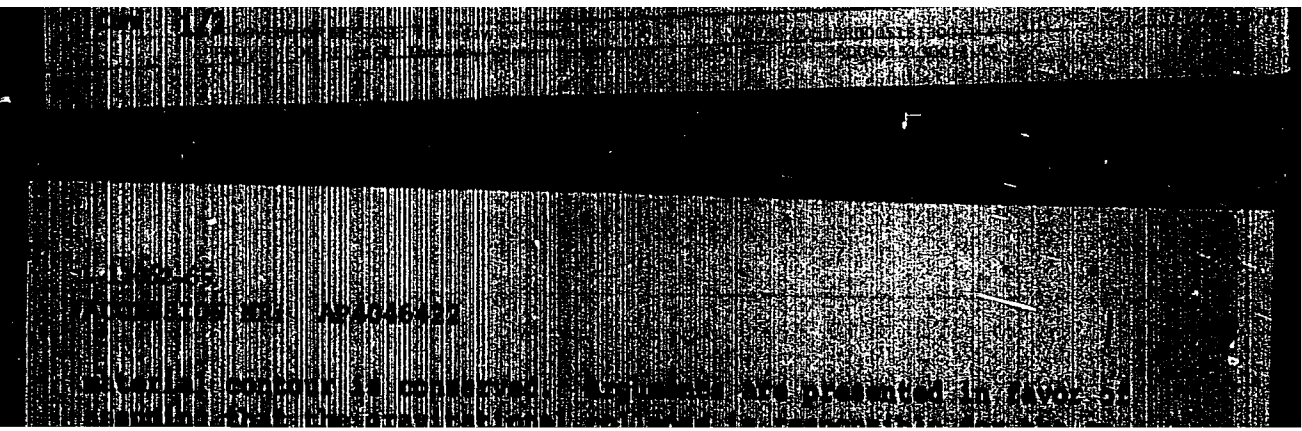
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NO REF SOV: 016

OTHER: 011

Card 2/2



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PROCESSES IN THE SOLAR WIND

The solar wind is a supersonic plasma that flows from the Sun into the solar system. It is composed of electrons and protons, with a temperature of approximately 10^6 K. The solar wind is a result of the high temperature and low density of the solar corona, which allows particles to escape the Sun's gravitational pull.

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СДЗ 2/2

Accession No. A500183

1005/2/07/005/229/228

Author: Wilhelm, V. J.
On the interaction of a proton with a magnetic field

Source: Journal of Experimental Physics, v. 47, no. 6, 1961, p. 228

Subject: Synchrotron radiation; cyclotron; beta radiation; proton interaction

ABSTRACT: The authors consider the processes associated with cyclotron radiation. The interaction of a proton with a magnetic field is considered in a 0-order approximation. It is shown that the radiation in a 0-order approximation is a function of the angle between the proton velocity and the magnetic field. It is possible for the radiation to be observed in principle, the authors believe, especially under high conditions where the proton velocity is close to c . The authors believe that at the magnetic fields involved in certain stars, even near collapsing stars, the

Page 1/2

6/0056/6-1047/006/2279/228

Accession No. A-100123

Author: Chandrasekhar, S.

Title: Proton and beta radiation from a neutron star in a magnetic field.

B

Source: Journal of Astrophysical Physics, v. 17, no. 6, 1967, p. 229-230.

Subject: radiation; neutron stars; beta radiation; proton

Abstract: The radiation from a neutron star is associated with cyclotron emission from the surface. The manifestation of a magnetic field is the emission of a proton and a beta particle from a neutron star. It is shown that the emission of a proton and a beta particle is possible for a neutron star with a magnetic field of the order of 10^{10} gauss. The conditions where the proton and beta radiation are emitted are discussed. In principle, the conditions where the proton and beta radiation are emitted are discussed. The results show that at the magnetic field strengths prevailing in the space near collapsing stars, the



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10076/6/07/006/2210/228- in a magnetic field

10076/6/07/006/2210/228- 17, no. 5, 156,

10076/6/07/006/2210/228-

10076/6/07/006/2210/228- proton

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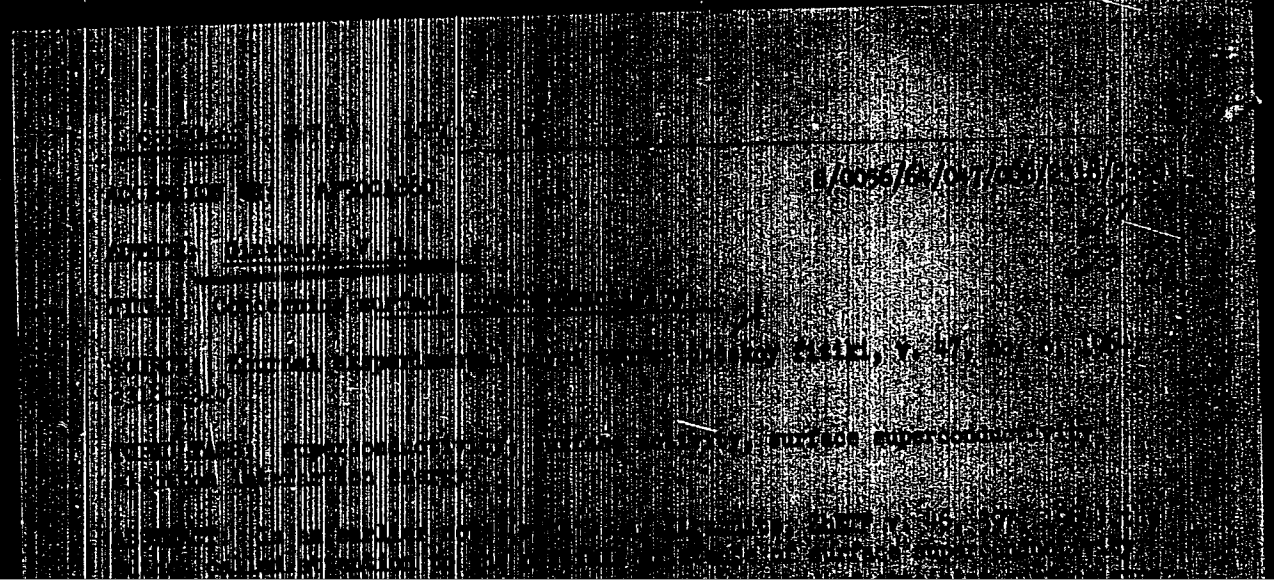
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ACCESSION NR: AP4031621

S/0053/64/082/004/0585/0647

AUTHOR: Ginzburg, V. L.; Kurnosova, L. V.; Razorenov, L. A.; Fradkin, M. I.

TITLE: Investigations of the nuclear component of cosmic radiation performed on Soviet satellites and rockets

SOURCE: Uspekhi fizicheskikh nauk, v. 82, no. 4, 1964, 585-647

TOPIC TAGS: cosmic ray, satellite measurement, space probe, cosmic ray charge distribution, cosmic ray flux, cosmic ray energy spectrum, solar cosmic ray, primary cosmic radiation, nuclear active component, electron positron component, galactic cosmic ray

ABSTRACT: This review summarizes results of measurements of cosmic-ray particle fluxes, cosmic-ray energy spectra, and intensity variations of the cosmic-ray components performed by the authors with satellite-borne equipment and reported in various publications (Geomagnetizm i aeronomiya, v. 2, 193, 1962. Iskustvennyye sputniki zemli, no. 2, 70, 1958; no. 5, 20, 1960; no. 8, 87, 1961; no. 12, 16, 1961; no. 5., 131, 1961; no. 12, 31, 1961; no. 15, 66, 1962. J. Phys. Soc. Japan v. 17, Suppl. A-II, 315, 1962. Izv. AN SSSR ser. fiz. v. 26, 782, 1962).

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ACCESSION NR: AP4031621

The experimental results are compared with the data by others. In addition, some problems and possibilities of cosmic-ray research outside the earth's atmosphere and magnetic field are also discussed. The advantages and limitations of satellite and rocket studies are briefly enumerated. Certain features of Cerenkov counters, which provided the bulk of the information, are discussed. Difficulties in the comparison of the results of different researches and the effect of the solar-activity cycle and of the individual solar flares are extensively dealt with. The correlation with solar radio emission is also discussed in connection with the electron-positron component of cosmic radiation. The section headings are: Introduction. I. Investigation of the nuclear component of cosmic rays with Soviet satellites and space probes. 1. Procedure. 2. Chemical composition of cosmic rays, fluxes of different nuclear groups and their energy spectra. 3. Variations of the flux of the nuclear cosmic-ray component and nuclei of solar origin. II. Use of satellites and rockets to study primary cosmic radiation. 4. Nuclear component of galactic cosmic rays. 5. Solar cosmic rays and high-latitude cutoff. 6. Electron-positron component of cosmic rays. Bibliography. Orig. art. has: 31 figures, 15 tables, and 6 formulas.

ACCESSION NR: AP4031621

ASSOCIATION: None

SUBMITTED: 00

ATD PRESS: 3059

ENCL: 00

SUB CODE: SV, AA

NO REV SOV: 063

OTHER: 074

Card 3/3

AUTHOR: STRAUSS, V. L.

TITLE: SOME PROBLEMS IN GAMMA RAY ASTRONOMY

SOURCE: REPORT (LABORATORY OR OTHER), 1964, 201-242

TOPIC TERMS: cosmic gamma ray production, cosmic ray, cosmic ray particle, x ray, astronomy, gamma astronomy, galactic, metagalactic radiation

ABSTRACT: The author discusses various mechanisms for the production of gamma rays in galaxies and the intergalactic space. Principal attention is given to the production of gamma rays by cosmic rays. The converse problem, that of obtaining information on cosmic rays from observations of the intensity and spectrum of the meta-galactic gamma rays, is also discussed. The sensitivity of x-ray astronomy is limited by the background radiation from the galaxy. Whether cosmic rays contribute to the background radiation is discussed. The existence of particles of high energy in intergalactic space. The existence of particles of high energy in intergalactic space. The existence of particles of high energy in intergalactic space. The existence of particles of high energy in intergalactic space.

CONFIDENTIAL

ACCESSION NR: AP4013323

S/0020/64/154/003/0557/0560

AUTHORS: Ginzburg, V.L. (Corresponding member); Ozernoy, L.M.;
Syrovatskiy, S.I.

TITLE: On the radiation mechanism of galaxy 3C 273-B

SOURCE: AN SSSR. Doklady*, v. 154, no. 3, 1964, 557-560

TOPIC TAGS: extragalactic, radiation source 3C 273-B, metagalactic object, optical radiation, bremsstrahlung, luminosity, Compton loss, free-free transition, bound-free transition, relativistic electron, photon, visible frequency

ABSTRACT: The extragalactic radiation source 3C 273-B, identified as an emissive star-shaped object of the 12th magnitude (M. Schmidt, Nature, 197, 1040/1963/), is one of the recently discovered new types of metagalactic objects. The exceptionally high luminosity and irregular changes in the brilliance of that galaxy are indicative of an unusual nature of the radiating object. The latter is

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ACCESSION NR: AP4013323

probably a "super star" rather than a cluster of stars, but the word "galaxy" is used here for want of a better term. It would be interesting to find out whether the continuous optical radiation from the galaxy 3C 273-B is due to magnetic bremsstrahlung. The optical radiation of galaxy 3C 273-B is not polarized, and could therefore easily be characterized also as non-magnetic bremsstrahlung. A contrary assumption would of course be wrong inasmuch as magnetic bremsstrahlung can, for a number of reasons, be completely depolarized. If the radiation from the object is of a braking nature (free-free and bound-free transitions), it cannot be considered as black body radiation in view of its spectral characteristics. Although the possibility of the bremsstrahlung nature of the mentioned radiation is not excluded, the spectrum of the other star-shaped extragalactic sources makes such a hypothesis considerably less probable (in the opinion of I. Shlovskiy). Orig. art. has: 11 formulas and 1 table.

ASSOCIATION: Fizicheskiy institut im. P.N. Lebedeva Akademii nauk

ACCESSION NR: AP4013323

SSSR (The P.N. Lebedev Physics Institute, Academy of Sciences SSSR)

SUBMITTED: 10Oct63

DATE ACQ: 26Feb64

ENCL: 00

SUB CODE: PH

NO REF SOV: 002

OTHER: 009

ACCESSION NR: AP4035806

S/0020/64/156/001/0043/0046

AUTHOR: Ginzburg, V. L. (Corresponding member)

TITLE: On Magnetic Fields of Collapsing Masses, and the Nature of Superstars

SOURCE: AN SSSR. Doklady*, v. 156, no. 1, 1964, 43-46

TOPIC TAGS: superstar, collapsing protostar, superstar magnetic field, general relativity theory, cosmic x ray, cosmology, cosmic radio wave

ABSTRACT: The author points out that the observation of a protostar (a gas cloud with a mass of many million times that of the sun) in an advanced stage of collapse will be difficult, if not impossible, because of the decrease of surface and the curvature of light waves. Therefore, one should look for other phenomena that might be connected with a collapsing superstar, such as a magnetic field. A magnetic field existing before the collapse, may be expected to be preserved during the collapse if a reasonable electroconductivity of the medium is assumed. The presence of a magnetosphere may cause a noticeable Zeeman splitting of spectral lines and a rotation of the plane of polarization of radio waves. Particles in this field will be the source of electromagnetic waves in the radio, optical,

Card 1/2

ACCESSION NR: AP4035806

and x-ray range (cyclotron radiation). This may explain the recently discovered discrete source of x-rays. Orig. art. has: no figures, 1 table, 3 equations.

ASSOCIATION: Fizicheskiy Institute im P. N. Lebedeva Akademii nauk SSSR
(Physics Institute, Academy of Sciences SSSR)

SUBMITTED: 31Jan64

DATE ACQ: 26May64

ENCL: 00

SUB CODE: B, AA, GP

NO REF SOV: 006

OTHER: 004

SECRET

Ro-4/Pa-5/Po-4/Pl-4/Pb-7/
8/0020/6A/158/004/0808/0810

AVIATION BR. 1540873
AUTHOR: G. I. GINSBURG (CONTRIBUTOR: V. L. GINSBURG), SY-KOVRALIK, V. I.

TITLE: Decay sources of x-radiation

SOURCE: AVIATION BR. 1540873, No. 4, 1964, 808-810

SYNOPSIS: Cosmic x-radiation decay source, magnetronstrahlung, relativistic electrons, magnetic field, electron spectrum, correlation of the cosmic rays.

Abstract: Recently, the cosmic origin was detected. The decay sources of x-radiation are unknown. V. L. Ginsburg and G. I. GINSBURG have proposed hypotheses about the source of cosmic x-radiation. The source is emitting magnetronstrahlung from a volume filled with relativistic electrons. The possibility of the magnetronstrahlung in one small part of this volume is shown. It is shown that the occurrence of magnetronstrahlung is a result of a hypothesis which expresses the magnetronstrahlung flux. Another formula represents the ratio of two

Card 1/2

00051190014

17/0020/64/158/004/080A/0810

ACQUISITION NO. 17/0020/64/158/004/080A/0810

AUTHOR: GINSBURG, V. L. (Corresponding member AN SSSR); SYROVICKIY, S. I.

TITLE: Discrete sources of x-radiation

ISSUES: AN SSSR, Doklady, vol. 122, no. 4, 1964, 808-810

B

TOPIC: Cosmic x-radiation, discrete source, magnetobremstrahlung, relativistic electron, magnetic field, electron spectrum, collisional energy transfer.

ABSTRACT: Recently, x-radiation of cosmic origin was detected. The discrete nature of this radiation is unknown. V. L. Ginsburg and S. I. Syrovitskiy have proposed a hypothesis about the source of cosmic x-radiation. They consider the source as emitting magnetobremstrahlung of electrons filled with relativistic electrons. It is shown that the magnetic field in one small part of this volume exceeds a critical value which favors the occurrence of magnetobremstrahlung. A formula was developed which expresses the magnetobremstrahlung flux. Another formula represents the ratio of two

Card 1/2

ACQUISITION

AUTHOR: GINSBURG, V. L. (Corresponding member AN SSSR); SYROVETSKIY, E.

TITLE: Discrete sources of x-radiation

SOURCE: AN SSSR, Doklady, v. 159, no. 4, 1964, 808-810

B

TOPIC: Cosmic x-radiation; discrete source; magnetobremstrahlung; magnetic field; electron spectrum; collisional excitation

ABSTRACT: Recently, x-radiation of cosmic origin was detected. The discrete sources of this radiation are unknown. V. L. Ginsburg and E. Syrovetskiy put forward a hypothesis about the source of cosmic x-radiation. They consider the source as emitting magnetobremstrahlung produced by collisions with relativistic electrons. It is shown that only one small part of this volume of the universe is required to explain the occurrence of magnetobremstrahlung. A formula is developed which expresses the magnetobremstrahlung flux. Another formula represents the ratio of two

Card 1/2

AUTHOR: GINZBURG, V. L. (Corresponding member AN SSSR); Syrovatskiy, S. I.

TITLE: Discrete sources of x-radiation

SOURCE: AN SSSR, Doklady*, v. 158, no. 4, 1964, 808-810

TOPIC TAGS: cosmic x radiation, discrete source, magnetobremstrahlung, relativistic electron, magnetic field, electron spectrum, collapsed star, neutron star

ABSTRACT: Recently, x-radiation of cosmic origin was detected. The discrete sources of this radiation are unknown. V. L. Ginzburg and S. I. Syrovatskiy have proposed a hypothesis about the source of cosmic x-radiation. Both authors consider the source as emitting magnetobremstrahlung from a volume filled with relativistic electrons. The intensity of the magnetic field in one small part of this volume attains a very high strength which favors the occurrence of magnetobremstrahlung. A formula was developed which expresses the magnetobremstrahlung flux. Another formula represents the ratio of two

L 12053-65

ACCESSION NR: AP4047317

Fluxes from different parts of the emitting volume where the electron spectra differ and the difference is characterized by a coefficient. The size of the emitting volumetric part is determined with assumed numerical values for the electron energy and the intensity of the magnetic field. Strong magnetic fields may exist only in parts of collapsed stars as neutron stars. Orig. art. has: 4 formulas.

ASSOCIATION: Fizicheskiy Institut im. P. N. Lebedeva Akademii nauk SSSR (Institute of Physics, AN SSSR)

SUBMITTED: 10JUL68 AND PRESS: 3120 ENCL: 00
SUB CODE: AA NO KEY SOV: 005 OTHER: 00A

Card 2/2

L 1888-66 EWT(1)/EWT(m)/FCC/T/EWA(h) IJP(c) GS/GW

ACCESSION NR: AT5022822

UR/0000/65/000/000/0008/0022

AUTHOR: Gineburg, V. L.; Kurnosova, L. V.; Logachev, V. I.; Razorenov, L. A.;
Fradkin, N. I.

TITLE: Primary component¹⁹ of cosmic rays

SOURCE: Vsesoyuznoye soveshchaniye po kosmofizicheskomu napravleniyu issledovaniy kosmicheskikh luchey. 1st, Yakutsk, 1962, Kosmicheskiye luchy i problemy kosmofiziki (Cosmic rays and problems in cosmophysics); trudy soveshchaniya. Novosibirsk, Redizdat Sib. otd. AN SSSR, 1965, 8-22

TOPIC TAGS: primary cosmic ray, cosmic ray particle, cosmic ray measurement, cosmic radiation composition

ABSTRACT: The article is a survey of reported experimental data on the composition of cosmic rays. The following groups of nuclei (other than protons and alpha particles) with charge $Z \geq 3$ are considered: (1) light nuclei with charge $3 \leq Z \leq 5$ (group L); (2) nuclei of the middle group with $6 \leq Z \leq 9$; (3) heavy nuclei with $Z > 10$ (group H). The symbol S is also used and designates nuclei with $Z \geq 6$ ($S = M + H$). It is shown that fluxes of different nuclei (including protons) should be compared for a given value of their hardness. As a rough general rule, nuclei of elements with atomic number Z are Z times more

L 1888-66

ACCESSION NR: AT5022822

frequent in cosmic rays than in nature. Difficulties involved in measurements of fluxes of the different groups of nuclei are described. High-altitude experiments definitely indicate the presence of lithium, beryllium, and boron nuclei (20-30% of the quantity of heavier nuclei) in the primary component of cosmic rays in the vicinity of the earth. Findings concerning the electron-positron component of cosmic rays are discussed, and the chemical composition of solar cosmic rays is considered. Differential energy spectra of protons and nuclei and their hardness spectra are analyzed. On the basis of the body of data accumulated thus far it is now possible to state that not only protons, but also multiply-charged nuclei are accelerated on the sun; however, this mechanism of particle acceleration is still unknown, and several such mechanisms may exist. Orig. art. has: 14 figures and 3 tables.

ASSOCIATION: Fizicheskii institut im. N. P. Lebedeva AN SSSR (Physics
Institute, AN SSSR)

SUBMITTED: 29Oct64

ENCL: 00

SUB CODE: AA, NP

NO REF SOV: 014

OTHER: 020

272

L 2326-66 ENT(1)/FCC/EWA(h) GS/GW
ACCESSION NR: AT5023626

UR/0000/65/000/000/0486/0501

AUTHORS: Ginzburg, V. I.; Kurnobova, L. V.; Razorenov, L. A.; Syrovatskiy, S. I.;
Fradkin, M. I. 36
ETA

TITLE: Some problems and perspectives in the investigation of primary cosmic rays

SOURCE: Vsesoyuznaya konferentsiya po fizike kosmicheskogo prostranstva. Moscow,
1965. Issledovaniya kosmicheskogo prostranstva (Space research); trudy konferentsii.
Moscow, Izd-vo Nauka, 1965, 486-501

TOPIC TAGS: cosmic ray, gamma ray, x ray, solar activity, antiparticle

ABSTRACT: Problems associated with the investigation of primary cosmic rays and gamma rays are presented in a three-part report. Part I deals with the proton-nucleus component of the cosmic rays, Part II covers the electron-positron component, and Part III discusses cosmic gamma- and x-rays. Although the proton-nucleus component of primary cosmic rays has been studied quite completely, a group of problems still remains unanswered. Eight such problems discussed in Part I are:
1) energetic spectra of protons and nuclei in the energy interval below 100 Mev/nucleon. These spectra are represented by the form $N(E) \sim E^{1.8}$. 2) The relationship between fluxes of different nuclei groups (L, M, H) in the energy range 55 to 550 Mev/nucleon, which is still not well known. 3) Isotopic components of primary

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L 2326-66

ACCESSION NR: AT5023626

cosmic rays. This would require the measurement of three independent parameters such as dE/dx , E , and pc . 4) The presence of high speed antiprotons generated by the interaction of cosmic rays with interstellar media. Some measurements place the percent composition of antiparticles at 0.23%. 5) The verification of the presence of superheavy nuclei, $Z > 30$. 6) Estimates of the time rate of change of the fluxes in primary nuclei components which have their origin either in solar bursts or in modulated galactic cosmic rays. These intensity variations should be recorded continuously, outside the terrestrial atmosphere. 7) Intensity gradients of cosmic rays in the solar system as evidenced by data from Pioneer-5 and Mariner-1. 8) Anisotropy among particle fluxes of low, near-threshold energies. Two similar problems are discussed in Part II. Here the flux and energy spectra of primary cosmic ray electron-positron components are analyzed first, where data are shown to be rather scant. Next, the relationship between positron and electron fluxes is considered by measuring the charge composition of the primary cosmic rays. In Part III, calculation results of expected γ - and x-ray intensities from important galactic sources are considered. The γ -ray generation is attributed to processes such as π^0 -meson decay, bremsstrahlung radiation of relativistic electrons and positrons, and Compton γ -rays by the scattering of photons on x-ray electrons. Experiments indicate $I_\gamma (> 50 \text{ Mev}) \approx 3.5 \times 10^{-4} \text{ photons/cm}^2/\text{sec/stere}$ which is larger than expected galactic estimates. This then implies γ -rays of

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L 2326-66

ACCESSION NR: AT5023626

meta-galactic origin. For lower energies (0.51 Mev) $I_{\gamma} = 1.2 \text{ to } 300 \times 10^{-6}$
photons/cm²/sec/stere. Orig. art. has: 6 tables, 2 figures, and 4 formulas. [04]

ASSOCIATION: none

SUBMITTED: 02Sep65

ENCL: 00

SUB CODE: AA, NP

ATD PRESS: 4/107

NO RRF SCV: 020

OTHER: 046

Card 3/3

L 28865-66

EW (M) / 1

SOURCE CODE: UR/0357/65/002/006/1103/1114

ACC NR: AP6018855

AUTHOR: Ginsburg, V. L.; Man'ko, V. I.

ORG: Physics Institute Im. P. N. Lebedev, AN SSSR (Fizicheskij institut AN SSSR)

TITLE: Relativistic oscillator models of elementary particles. This paper was given at the 14th Annual Conference on Nuclear Spectroscopy, Tbilisi, February 1964.

SOURCE: Yadernaya fizika, v. 2, no. 6, 1965, 1103-1114

TOPIC TAGS: nuclear particles, nuclear spin

ABSTRACT: Relativistic invariant equations are considered for particles described by the "center of mass" coordinates x_i and three 4-vectors u_i ($i = 1, 2, 3, 4$; $\alpha = 1, 2, 3$); u_i are the intrinsic variables which obey the oscillator equations. The solutions of these equations express the group $SU(3)$ in the rest system. It is possible to write the equations which give the solution corresponding to the octet and decuplet for the case of half-integer spin and to the octet and singlet for the case of integer spin. The authors thank M. A. Markov for his discussions. Orig. art. has: 25 formulas and 3 tables. Based on authors' Eng. abst. JPRS

SUB CODE: 20 / SUBM DATE: 15 May 65 / ORIG REF: 008 / OTH REF: 006

Card 1/1 UC

15-071-65 EIT(17)/1/200(1)-2 P1-1 IFF(1) 00
ACCESSION NO: AP5015941 02/0017765/013/005/0175/0210

25

AUTHOR: Agabekov, V. M. Glazovskiy, V. L.

13

TITLE: Some problems in crystal optics with spatial dispersion and the exciton theory

SOURCE: Fortschritte der Physik, v. 13, no. 5, 1965, 175-210

TOPIC TAGS: crystal optics, exciton, crystal structure, crystal theory

ABSTRACT: The purpose of this article, which was translated from the Russian by K. H. Herrmann, was to provide additional data and clarification for the authors' article published earlier in Fortschritte der Physik v. 11, 1963, p 163. The salient features of both the original article and the present article are scheduled for early publication by John Wiley and Sons. The following subjects are discussed: introduction of the $\epsilon_{ij}(\omega, k)$ tensor into crystals, separation of this tensor from the transverse field E , dispersion relations for the refraction index of light waves in the various ray direction and group velocity in an absorbing medium, relations between the refractive index of light waves and frequency considering spatial dispersion and adsorption both in optically active and

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U 53871-02
ACCESSION NO: AP5015961

optically inactive media, the boundary conditions, and the macroscopic theory of surface excitons. Orig. art. lang: Russ. 79 formulas and 7 figures.

ASSOCIATION: None

SUBMITTED: 00

ENCL: 00

SUB CODE: OP

NO REF SOV: 022

OTHER: 000

Card 2/2

ANNILOV, A. A., inzh.; BARALEYNIK, Ye. M., inzh.; BIRGEN, D. I.,
inzh.; BIRK, I. S., inzh.; BIL'OT, A. I., inzh.; BIL'BERG, V. L.,
inzh.; ZABELIN, V. I., inzh.; ZHELEZNY, Ye. G., inzh.; ISAYEV,
E. V., inzh.; KLIMOVITSKIY, A. M., inzh.; KNYAZEV, V. V., inzh.;
KOTOV, V. A., inzh.; LEYDERMAN, A. Ye., inzh.; POGORELICKIY,
M. L., inzh.; SAZHAEV, V. I., inzh.; SEVAST'YANOV, V. V., inzh.;
SIL'NIKOV, S. K., inzh.; FROBERG, A. K., inzh.; SHENYEROV, M. S.,
inzh.; KALIKH, A. M., inzh.; VEREBOVSKIY, M. I., red.; TUBKOV,
G. A., red.; KALINA, T. O., red.; GUMBAROV, Ye. Ya., red.;
ANTONOV, D. I., vel. red.

[Low means of automatic and centralized control for nonfer-
rous metal mines] Novye sredstva avtomatizatsii i dispetcher-
skoy upravleniya dlya ruznikov tsvetnogo metallurgii. Moskva,
Izd. 1968. 96 p. (SIA 18:4)

L 6948-006

ACC NR: AP 5026225

SOURCE CODE: UR/0048/65/029/010/1819/1824

AUTHOR: Ginzburg, V.L.; Syrovatskiy, S.I.

29
B

ORG: Physics Institute in. P.N.Lebedev, Academy of Sciences,SSSR(Fizicheskiy institut Akademii nauk SSSR)

TITLE: Fundamental problems in cosmic ray astrophysics /Report, All-Union Conference on Cosmic Ray Physics held at Apatity, 24-31 August 1984/

SOURCE: AN SSSR. Izvestiya. Seriya fizicheskaya, v.29, no.10,965,1819-1824

TOPIC TAGS: Primary cosmic ray, galaxy, cosmology, astrophysics

ABSTRACT: Recent literature on the astrophysical aspects of cosmic ray physics is reviewed and some related questions, particularly that of the origin of cosmic rays, are discussed. Advances in astronomy and radioastronomy in the past decade have attracted attention to the possible cosmological significance of cosmic rays. Cosmic rays, for example, might serve as a mechanism for carrying off the energy released in the gravitational collapse of stars or galaxies. There is convincing evidence that the energy density of cosmic rays is much less in metagalactic space than in our Galaxy. Arguments to the contrary by G.R.Burbidge and F. Hoyle (Proc.Phys.Soc.,84 141 (1964) are specifically criticized. Most of the galactic cosmic rays, therefore, must be of galactic origin. The electron component of the primary cosmic radiation has been shown to arise mainly from the direct acceleration of electrons rather than

Card 1/2

0902 0060

ACC NR: AP 8028225

from meson decay. The stationary model of cosmic ray origin given by the authors (Proiskhozheniye kosmicheskikh luchey, Izd. AN SSSR, 1963), according to which most cosmic rays originate in supernovae, has not encountered any difficulties but cannot be regarded as proved. There is no convincing evidence that an explosion of the nucleus of our Galaxy has occurred in the past 10^9 years. The explosions observed in M82 and NGC 5128 do not indicate that such explosions are probable in our Galaxy, for M 82 and NGC 5128 are not spirals and differ considerably from our Galaxy. It has been argued that if all cosmic rays originated simultaneously, the relativistic particles would have travelled farther and encountered more interstellar matter than the low-energy particles, and would therefore contain a larger fraction of light nuclei, whereas such observational evidence as exists indicates that the higher energy cosmic rays have a smaller fraction of light nuclei. This argument is not decisive, however, because the high-energy cosmic rays may have travelled through regions in which the density of interstellar matter is less than in those through which the low-energy cosmic rays passed. It may be possible soon to determine whether the cosmic ray intensity was significantly greater some 10^8 years ago than now by investigating meteorites as proposed by the authors (Proc. Internat. Conf. Cosmic Rays, Jaipur, 3 301 (1964); Izv. AN SSSR. Ser.fiz., 28, 1910 (1964); Astron.zh. 41, 430 (1964). Although the authors do not see any real arguments in favor of a non-stationary model for the origin of cosmic rays in the Galaxy, they agree that further investigation of this matter is justified. Orig.art. has 2 formulas and 1 table

SUB CODE: AA SUBM DATE: 00/--Oct65 ORIG.REF: 012 OTH REF: 007

Card 2/2

beh

MEMORANDUM FOR THE DIRECTOR, CIA

Re: [Illegible] (MIRA 18:10)

1. [Illegible]

GINZBURG, M.L.; SMIRINA, G.I.

Use of various light sources for atomic absorption analysis. Zav. lab.
31 no.2:249-250 '65. (MIRA 18:7)

1. Konstruktorskoye byuro "TSvetmetavtomatika".

GINZBURG, V. I. [Russian] and [Russian] PEYINOV, I. M., ed.

[Modern astrophysics; some results and prospects. Tendency of development.] Sovremennaya astrofizika; nekotorye rezultaty i perspektivy. Tendentsiya razvitiia. Moskva, Znanie, 1965. 22 p. (Novoe v zhizni, nauke, tekhnike. IX Seriya: Fizika, matematika, astronomiia, n.10) (MIRA 18:7)

1. [Russian] [Russian] (for Ginzburg).

ZOGANOVICH, Vladimir Melnyevich; GINZBURG, Vitaliy Lazarevich;
VIRKO, I.G., red.; RAYSKAYA, N.A., red.

[Crystal optics with allowance for spatial dispersion and
excitation theory] Kristallogoptika s uchetom prostranstvennoi
dispersii i teoriia eksitonov. Moskva, Nauka, 1974. 37 p.
(SIRA 18:5)

AKHIEZER, I.I., akademik; KEMYSH, M.V., akademik; KREIN, I.I., akademik;
KULIKOV, B.P.; KUSHNAROV, I.F.; LITVINOV, A.A.; LITVINOV, A.N.,
akademik; SHCHUKIN, D.V., akademik; SHCHUKIN, A.P., akademik;
AMBARTSUMYAN, V.A., akademik; TRUDYKIN, Ya.B.; GOSMAN, M.M.,
akademik; KOTLYNIKOV, V.A., akademik; LEPKOV, I.M.; KRESIN, V.I.,
akademik; GIMBERG, V.I.; MILIKHOVICH, M.P., akademik

Some problems in the development of modern physics: discussion of
the work of the Department of General and Applied Physics. Vest.
Sverdlovsk. no.2:3-46. 1965. (Sov. Phys. Usp. 8:3)

1. Citeny-korrespondenty AN SSSR (for Vol. V, reprints in, Leningrad,
Lifshits, Ginzburg).

[Faint, illegible text, possibly bleed-through from the reverse side of the page]

IN THE MATTER OF THE ESTATE OF J. Edgar Hoover, deceased.

As per agreement in resolution. Mar. 11, 1964. (100-101)

L 3676-66 EWT(n)/EWP(w)/T/EWP(t)/EWP(b) JD

UR/0020/65/163/006/1352/1355

ACCESSION NR: AP5021885

AUTHORS: GINSBURG, Y. I., (Corresponding member AN SSSR); Motulevich, G. P.; Pitayevskiy, I. P. 77
55 8

TITLE: Optical properties of polyvalent metals and interelectronic interaction

SOURCE: AN SSSR. Doklady, v. 163, no. 6, 1965, 1352-1355

TOPIC TAGS: Fermi surface, optical electron, valence electron, polyvalent metal, polyvalent metal electron

ABSTRACT: A critical examination of literature data pertaining to the ratio of optical electrons to valence electrons for Au, Ag, Al, Sn, and Pb is presented. The data are examined in terms of the Fermi surface integral

$$\oint v dS \approx \frac{4\pi p_F^3}{m} \approx \frac{3(2\pi\hbar)^3}{2m} N_{val}$$

where m and p_F are the mass and momentum on the Fermi surface of the free electron and N_{val} the number of valence electrons per cm^3 . It is concluded that the observed decrease of N optical from N valence in Al, Sn, and Pb is most probably due to interelectronic interactions. However, two difficulties regarding the

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L 3676-66

ACCESSION NR: AP5021885

6
above explanation are noted. For all polyvalent metals, the number of optical electrons per atom $n_{opt} \approx 1$, a fact the authors are unable to explain. The assumption that interelectron interaction exerts a strong influence on the ratio n_{opt}/n_{val} is not supported by any existing theoretical model. The authors thank M. Ya. Azbel⁵ and D. Pajns for their helpful criticism. Orig. art. has: 2 tables and 7 equations. 55

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva, Akademi nauk SSSR (Physical Institute, Academy of Sciences, SSSR); Institut fizicheskikh problem, Akademi nauk SSSR (Institute for Physical Problems, Academy of Sciences, SSSR) 55

SUBMITTED: 04 Jun 65

ENCL: 00

SUB CODE: MM, OP

NO REF SOV: 013

OTHER: 004

MC
Card 2/2

ACC NR: AR6019473

SOURCE CODE: UR/0269/66/000/002/0037/0037

AUTHOR: Ginzburg, V. L.

TITLE: Space-physics trend in cosmic-ray research

SOURCE: Ref. zh. Astronomiya, Abs. 2.51.313

REF SOURCE: Sb. Kosmich. luchy i probl. kosmofiz. Novosibirsk, Sib. otd. AN SSSR, 1965, 5-6

TOPIC TAGS: cosmic ray, electromagnetism, radiation belt, *INTERPLANETARY SPACE*

ABSTRACT: A brief historical review of cosmic-ray research is given. It is noted that in the 1955-57 period a new era in cosmic-ray research began, i.e., a study of the origin of cosmic rays and the relationship between the variations of cosmic rays and electromagnetic phenomena in interplanetary space, in the radiation belts of the earth and of the planets. The significance of the new space-physics aspect of cosmic-ray research is stressed in conjunction with the launching of artificial satellites and space rockets. N. K. Translation of abstract

SUB CODE: 04/

Card 1/1

UDC: 523.165

SOURCE CODE: UR/0033/66/043/005/0097/0092

AUTHOR: Ginsburg, V. L.

ORG: Physics Institute im. P. N. Lobodova, Academy of Sciences SSSR (Fizicheskii institut, Akademii nauk SSSR)

TITLE: X ray emission of radio galaxies

SOURCE: Astronomicheskii zhurnal, v. 43, no. 5, 1966, 897-902

TOPIC TAGS: galaxy, galactic radiation, x ray emission, bremsstrahlung

ABSTRACT: The cause of the intense x ray emission from radio galaxies Virgo and Cygnus is discussed. Objections are made to magnetic and Compton mechanisms and to quasar type nuclei in the radio galaxy or to stellar groups as explanations of the source of the x ray emission. Instead, the bremsstrahlung mechanism is proposed. In the case of such emission from large clouds of hot gas, the internal energy of the gas would be of the same order of magnitude as the total cosmic ray energy in the source and the gas could be heated by the change in the energy of the cosmic rays during their isotropization. The difficulty with the proposed mechanism is that for an extended source the hot gas mass should be larger. X ray emission of the Crab nebula is also discussed in the light of the proposed small hot gas cloud. Need for additional measurements of polarization, dimensions, shape and spectra of the x ray

UDC: 523.164.4

ACC NR: AT6027583

SOURCE CODE: UR/0000/66/000/000/0015/0020

AUTHOR: Ginzburg, V. L. (Corresponding member AN SSSR)

ORG: none

TITLE: New trends in astronomy

SOURCE: Zvezdy i vselennaya (Stars and the universe). Moscow, Izd-vo Znaniye, 1966, 15-20

TOPIC TAGS: descriptive astronomy, radio astronomy, primary cosmic ray, neutrino, scintillation, galactic nebula, supernova

ABSTRACT: The first revolution in astronomy, followed by a rapid progress of the science, occurred in 1610 when Galileo observed the sky for the first time with an armed eye. This opened up the optical era in astronomy. The second revolution occurred in 1945 when optical astronomy became universal. Now, twenty years after this second revolution, six new branches of astronomy have been developed: astronomy of radiofrequency waves; x-ray and gamma-ray astronomy; study of primary cosmic rays; optical astronomy from satellites in visible, infrared, and ultraviolet rays; direct methods using satellites (e.g., photographing the Moon and Mars from short distances) and reception of neutrinos (they pass through the Sun and Earth in small but detectable amounts). These new methods have already yielded significant results. The development

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of new techniques has eliminated the disadvantage of the small angular resolution of radiofrequency astronomy and made its angular resolution even greater than that of optical astronomy. This is done by three methods: the use of radiofrequency interferometry (placing several long-wave telescopes at large distances from each other), use of overlapping the sources of radiofrequency-wave radiation by the Moon, and scintillation twinkling of very small radiation sources. A number of significant discoveries were made in 1965 based on new astronomical methods. The small source, especially powerful on long waves, was discovered by the British (and confirmed in the USSR) in a nebula remaining from a supernova star, the envelope of which expanded at the rate of 1000 km/sec. It is possible that this source is a remnant of the supernova star. This fact could possibly solve the problem of the nebulae. Other important discoveries made in 1965 included the detection of quasiastral galaxies, among them the 3S-9 quasar, which is moving away from us at a velocity equal to 80% of the speed of light, and a most spectacular discovery of radiation that was emitted 10 billion years ago.

SUB CODE: 03/ SUBM DATE: 22Apr66

S/115/62/000/011/002/008
E192/E382

AUTHOR: Ginzburg, V.M.

TITLE: Relationship between information and energy during
measurements

PERIODICAL: Izmeritel'naya tekhnika, no. 11, 1962, 6 - 8

TEXT: The quality of measuring equipment can be estimated by the formula $\eta = I/W$, where I is the desired information and W is the energy consumed during measurement leading to the recovery of information. The problem consists of calculating η and determining the conditions for which η is a maximum. The measured system is in the form of material particles which are either stationary or mobile. It is found in the case of stationary particles that: 1) the quantity of energy required for measuring the position of a finite number of particles in an enclosed volume is independent of the number of particles and is directly proportional to the quantity of the resolution elements in volume; 2) for a known number of particles in the observation volume the information per unit of energy is a maximum when the number of particles n is approximately equal to the overall number of the resolution elements

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S/115/62/000/011/002/008
E192/E382

Relationship between

N in the volume; 3) if any number of particles from 0 to N is equally probable, the information per unit energy decreases monotonically with increasing N from $1 - 1/2$ bit; 4) if a probability distribution $p(n)$ is given, where $0 < n < N$ and its maximum occurs at $n = n_1$, the information is a maximum for N slightly smaller than $2n_1$. In the case of moving particles, the measurement is effected by equipment producing probe pulses of duration τ_p and a period T . Two cases are possible. In the first case, the character of motion of the particles is not known and each of the particles can be found in any of the N elements contained in the volume under observation. The measurements are conducted for this case in the same manner as that used for the stationary particles. In the second case the particles can appear only in a portion $N_1 < N$ of the volume under observation and the maximum time during which the particles remain inside the volume is known (T_n). Now, the analysis shows that there exists an optimum measurement period when a minimum energy is necessary for securing the required information; the optimum period decreases with increasing velocity of the particles. Secondly, the gain due to the correct choice of

Card 2/3

S/115/62/000/011/002/008
E192/E382

Relationship between

parameters of the measuring equipment increases with the number
of particles passing through the volume under observation.

Card 3/3

L 18398-c3 EWT(1)/EWG(k)/BDS/EEC-2/EEC(b)-2/ES(w)-2 AFFTC/ASD/ESD-3/
APGC/AFWL/IJP(C)/SSD Pz-h/Pi-h/Pj-h/Po-h/Pab-h AT
ACCESSION NR: AP3003729 S/0109/63/008/007/1269/1274

87

AUTHOR: Ginzburg, V. M.

TITLE: Electromagnetic energy propagation in a circular waveguide filled with gas discharge plasma whose parameters are variable

SOURCE: Radiotekhnika i elektronika, v. 8, no. 7, 1963, 1269-1274

TOPIC TAGS: waveguide propagation, circular waveguide, gas discharge plasma

ABSTRACT: A method for determining average electron concentration by the deviation of a beam in a plasma-containing slot antenna was investigated theoretically and experimentally. A formula for average electron concentration is developed for the case of a discharge tube that is also a metal cylindrical waveguide. The formula was experimentally verified by a metal-coated (Ar, 0.09-mm Hg) tube with a longitudinal slot which acted as a directional antenna. Tabulated data shows good agreement between the values of electron concentration measured

L 18398-03

ACCESSION NR: AP3003729

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by the author's method and those measured by the conventional Langmuir's probe method. Orig. art. has: 4 figures, 14 formulas, and 2 tables.

ASSOCIATION: none

SUBMITTED: 05Jul62

DATE ACQ: 02Aug63

ENCL: 00

SUB CODE: PH, GE

NO REF SOV: 007

OTHER: 001

Card 2/2

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G. N. B. V. M. I. P. S. O. V. , .

Electromagnetic waves in a plane waveguide with nonuniformity
in the direction of propagation. Radiotekhnika i Elektronika, No. 5, p. 8-12, 1965.
(USSR 1965)

GINZBURG, V. M.

V. M. GINZBURG, I. N. Belova; "Computation of the directivity pattern of image antennas." Scientific Session Devoted to "Radio Eng.", May, 1970, Truzhervindat, Moscow, 9 Sep. 50

A method is proposed to compute the directivity pattern of image antennas with beam hunting on an electronic digital computer.

Results are presented of a computation on the BESM of the spatial amplitudes and phases of the directivity pattern of parabolic antennas for the main and cross-polarization field components. The results of the computation of the amplitude patterns for the main polarization are confirmed experimentally to hunting angles of 30° for a beam width of about 1.

Universal dependences of the direction of the fundamental maximum, the side lobe levels, beam width, drop in gain, phase at the maximum and width of the vector directivity patterns on the displacement of the emitter from the focus are constructed for various antenna parameters for hunting angles which do not exceed ten times the beam width.

GINZBURG, Vera Moiseyevna; BELOVA, Inessa Nikolayevna; ALEKSANDROVA,
A.A., red.; SMUROV, B.V., tekhn.red.

[Calculation of parabolic antennas] Raschet parabolicheskikh
antenn. Izd-vo "Sovetskoe radio," 1959. 249 p. (MIRA 13:2)
(Antennas (Electronics))

L 20957-66 EMI(1)/ETC(f)/EPF(n)-2/EWG(m)/EWA(h) IJP(c) CG/AT

ACCESSION NR: AP5013340

UR/0109/65/010/005/0868/0878

621.372.822.09

AUTHOR: Ginzburg, V. M.; Maripov, A.

TITLE: Electromagnetic waves in a flat waveguide containing heterogeneous isotropic plasma

SOURCE: Radiotekhnika i elektronika, v. 10, no. 5, 1965, 868-878

TOPIC TAGS: flat waveguide, plasma, rectangular waveguide

ABSTRACT: The propagation of a magnetic wave in a flat laminated waveguide, whose internal layer is represented by a plasma, is theoretically considered. A method is developed for calculating the propagation constant of such a waveguide which contains a heterogeneous plasma with a supercritical concentration of electrons at the axis. The possibility is considered of using the method for designing a rectangular waveguide with a round discharge tube; a small distortion of the field by the plasma and limited electron concentration are assumed.

Card 1/2

I. 20957-66

ACCESSION NR: AP5013340

Formulas are derived for computing the distribution of field amplitude and phase over the waveguide cross-section, and an example of the application of these formulas for the cases of low and supercritical concentrations in the center of the discharge is given. Orig. art. has: 7 figures, 25 formulas, and 1 table.

ASSOCIATION: none

SUBMITTED: 07Mar64

ENCL: 00

SUB CODE: EC, EM

NO REF SOV: 005

OTHER: 000

Card 2/2

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I 33158-66 ENT(1) 00
ACC NR. AP6014239

SOURCE CODE: UR/0100/66/0117005/88507085

AUTHOR: Ginzburg, V. M.; Yakovleva, Ye. A.

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B

ORG: none

TITLE: Electromagnetic field in a coaxial line containing an inhomogeneous isotropic dielectric

SOURCE: Radiotekhnika i elektronika, v. 11, no. 5, 1966, 850-859

TOPIC TAGS: electromagnetic field, isotropic dielectric, dielectric loss, dielectric constant

ABSTRACT: A method is described for computing the electromagnetic wave-propagation constant through a coaxial line containing an isotropic dielectric with inhomogeneous losses over the cross section. An analysis is made of a case in which a line filled with gas-discharge plasma, whose dielectric constant changes along the cross section, approximately following the parabolic law from the minimal value of less than one on the axis line up to one along its walls. A study has been made of the distribution of the amplitude and the field phase in this line. For comparative purposes, a field has been considered in a coaxial line filled with an inhomogeneous dielectric whose constant changes according to the same law from the maximal value of more than one at the axis to one along the walls. Orig. art. has: 6 figures and 23 formulas. [Translation of author's abstract.] [AM]

SUB CODE: 20,09/ SUBM DATE: 24Aug64/ ORIG REF: 008/

Ord 1/15

UDC: 621.372.224.09

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(MIRA 13:11)

(Epoxy resins)

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SO: Letopis Zhurnal'nykh Statey, Vol. 44.

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(Altai Mountains--Anthropology) (Anthropology--Altai
Mountains)

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